

ALS 8.3.1 Team

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Acknowledgements

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8.3.1 PRT: Tom Alber

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Center for HIV Accessory and Regulatory Complexes
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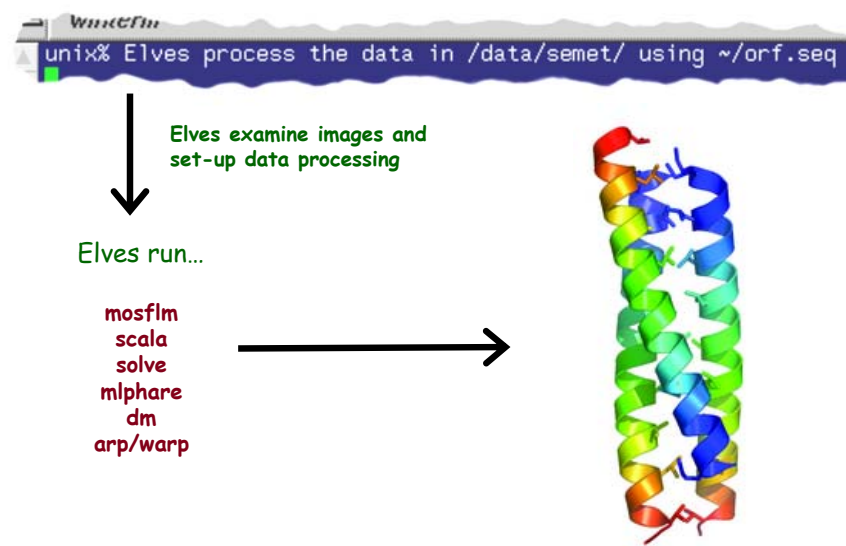
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University of California San Francisco

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Elven Automation



Holton and Alber *PNAS USA* **101** 1537-42 (2004)

Elven Automation

How often does it
really work?

what are my chances?

100 s/dataset

what are my chances?

100 s/dataset

200 days/year

~150 beamlines

what are my chances?

100 s/dataset

200 days/year

~150 beamlines

~26,000,000 datasets/year

what are my chances?

100 s/dataset

200 days/year

~150 beamlines

~26,000,000 datasets/year

7000 PDBs in 2008

Turning Data into Models

35 operating US beamlines
in 2003

Turning Data into Models

35 operating US beamlines

2×10^{13} ph/s

<http://biosync.sdsc.edu/>

Turning Data into Models

35 operating US beamlines

$\sim 10^{11}$ ph/ μm^2 exposure limit

2×10^{13} ph/s

Henderson et al (1990)

Turning Data into Models

35 operating US beamlines

$\sim 10^{11}$ ph/ μm^2 exposure limit

2×10^9 ph/ $\mu\text{m}^2/\text{s}$

<http://biosync.sdsc.edu/>

Turning Data into Models

35 operating US beamlines

$\sim 10^{11}$ ph/ μm^2 exposure limit

$\div 2 \times 10^9$ ph/ $\mu\text{m}^2/\text{s}$

= 400,000 datasets/year

Turning Data into Models

35 operating US beamlines

$\sim 10^{11}$ ph/ μm^2 exposure limit

$\div 2 \times 10^9$ ph/ $\mu\text{m}^2/\text{s}$

$\sim 200,000$ datasets/year

Turning Data into Models

35 operating US beamlines

$\sim 10^{11}$ ph/ μm^2 exposure limit

$\div 2 \times 10^9$ ph/ $\mu\text{m}^2/\text{s}$

$\sim 100,000$ datasets/year

Turning Data into Models

35 operating US beamlines

$\sim 10^{11}$ ph/ μm^2 exposure limit

$\div 2 \times 10^9$ ph/ $\mu\text{m}^2/\text{s}$

$\sim 100,000$ datasets/year

$\div 1324$ str in 2003

Jiang & R.M. Sweet (2004)

Turning Data into Models

35 operating US beamlines

$\sim 10^{11}$ ph/ μm^2 exposure limit

$\div 2 \times 10^9$ ph/ $\mu\text{m}^2/\text{s}$

$\sim 100,000$ datasets/year

$\div 1324$ str in 2003

$\sim 2\%$ efficient

Turning data into models

Turning data into models

8.3.1 in 2003

Number	Description	Percent
	Images	

Turning data into models

8.3.1 in 2003

Number	Description	Percent
446028	Images (~7 TB)	33%

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	MAD/SAD	

Turning data into models

8.3.1 in 2003

Number	Description	Percent
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449	MAD/SAD (1:2)	19%

Turning data into models

8.3.1 in 2003

Number	Description	Percent
446028	Images (~7 TB)	33%
2346	Data sets	47%
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	Published	

Turning data into models

8.3.1 in 2003

Number	Description	Percent
446028	Images (~7 TB)	33%
2346	Data sets	47%
449	MAD/SAD (1:2)	19%
104	Published	4.4%

DVD data archive



Elven Automation

Apr 6 – 24 at ALS 8.3.1

27,686	images collected
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Elven Automation

Apr 6 – 24 at ALS 8.3.1

27,686	images collected
148	datasets (15 MAD)

Elven Automation

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148	datasets (15 MAD)
31	investigators

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5 KDa – 23 MDa	asymmetric unit

Elven Automation

Apr 6 – 24 at ALS 8.3.1

27,686	images collected
148	datasets (15 MAD)
31	investigators
56	unique cells
5 KDa – 23 MDa	asymmetric unit
0.94 – 32 Å	resolution (3.2 Å)

Elven Automation

Apr 6 – 24 at ALS 8.3.1

148 datasets

Elven Automation

Apr 6 – 24 at ALS 8.3.1

148	datasets
117	succeeded

Elven Automation

Apr 6 – 24 at ALS 8.3.1

148	datasets
117	succeeded
~3.5 (0.1-75)	hours

Elven Automation

Apr 6 – 24 at ALS 8.3.1

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~3.5 (0.1-75)	hours
31	failed

Elven Automation

Apr 6 – 24 at ALS 8.3.1

148	datasets
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~3.5 (0.1-75)	hours
31	failed
~61 (0-231)	hours

Elven Automation

Apr 6 – 24 at ALS 8.3.1

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~61 (0-231)	hours
2 / 15	MAD structures

Why do structures fail?

Overlaps

Radiation Damage

Signal to noise

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Signal to noise

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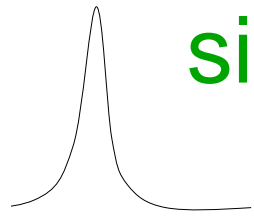
Radiation Damage

Signal to noise

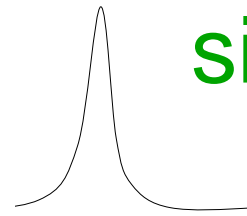
Elven Automation

Apr 6 – 24 at ALS 8.3.1

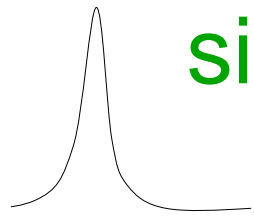
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117	succeeded
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31	failed
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2 / 15	MAD structures



signal



signal vs noise

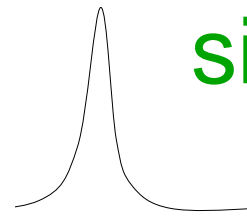


signal vs noise



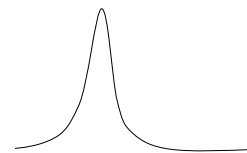
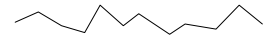
“If you don’t have
good data,
then you have
no data at all.”

-Sung-Hou Kim



signal vs noise

easy

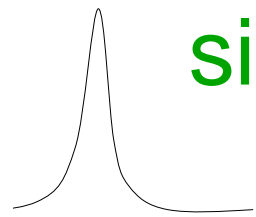


hard



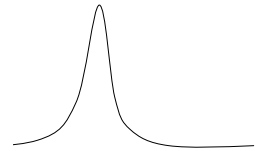
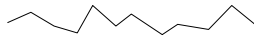
impossible





signal vs noise

easy



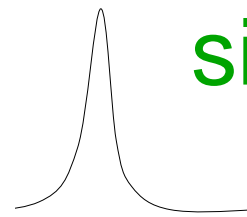
hard



threshold of "solvability"



impossible



signal vs noise



"If you don't have
good data,
then you must
learn statistics."

-James Holton

Adding noise

Adding noise

$$1 + 1 = 1.4$$

Adding noise

$$1 + 1 = 1.4$$

$$\sigma_{\text{total}}^2 = \sigma_1^2 + \sigma_2^2$$

Adding noise

$$1^2 + 1^2 = 1.4^2$$

$$\sigma_{\text{total}}^2 = \sigma_1^2 + \sigma_2^2$$

Adding noise

$$1^2 + 1^2 = 1.4^2$$

$$3^2 + 1^2 = 3.2^2$$

$$\sigma_{\text{total}}^2 = \sigma_1^2 + \sigma_2^2$$

Adding noise

$$1^2 + 1^2 = 1.4^2$$

$$3^2 + 1^2 = 3.2^2$$

$$10^2 + 1^2 = 10.05^2$$

Sources of noise

“photon counting”	$\sigma(N) = \sqrt{N}$
Read-out noise	rms 11.5 e-/pixel
Shutter jitter	rms 0.57 ms
Beam flicker	0.15 %/$\sqrt{\text{Hz}}$
spot shape	pixels? mosaicity?
radiation damage	B/Gray?

Simulated diffraction image

Simulated diffraction image

Arndt & Wonacott (1977)

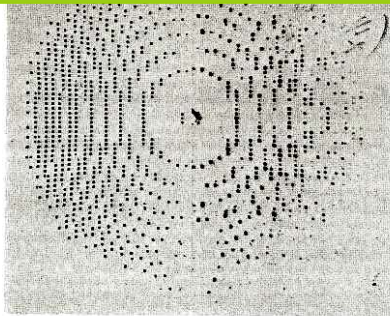


Fig. 7.12. Comparison of simulated diffraction pattern with small-angle rotation photograph.

The most practical means of testing for over reflexions is to specify a minimum separation between

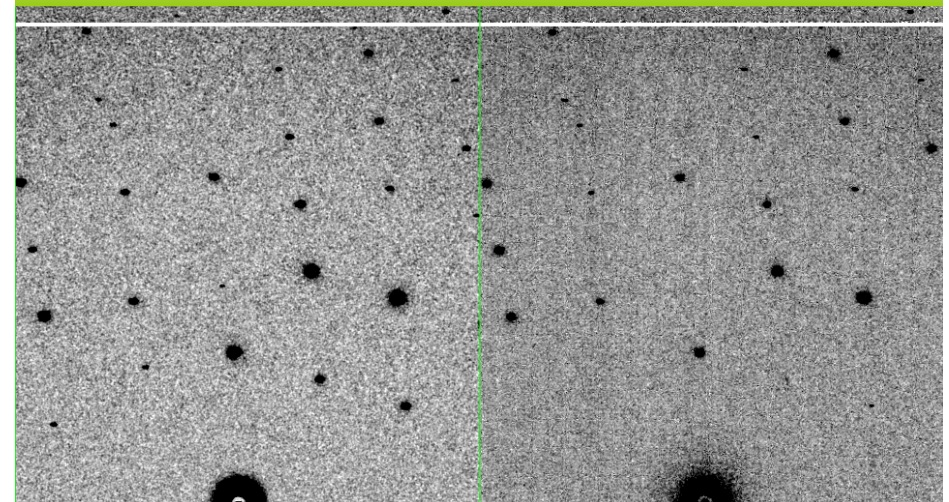
7.6. Choice of Crystal Orientation

Several factors must be considered in choosing a method used for the collection of a three-dimensional data set. These factors depend on the individual characteristics of the object being measured. The general applicability can, however, be made.

In principle, data could be collected from a crystal with a preferred orientation but this creates certain practical difficulties. It is necessary to choose an appropriate oscillation range for each photograph. This is more difficult and the total range of rotation required is larger. The reciprocal space of an asymmetric unit of reciprocal space may be greater than the reciprocal space of the crystal as a whole, so that when the crystal is rotated about a particular axis, the reciprocal space of the unit is not necessarily rotated in the same way as the reciprocal space of the crystal.

Recording a complete data set will, in general, require the use of more crystals than necessary, it being necessary to rotate the crystal to within, say, one degree of the desired azimuth from a still photograph. Then angular increments for adjacent crystals need only be minimal in order to avoid overlap of the data.

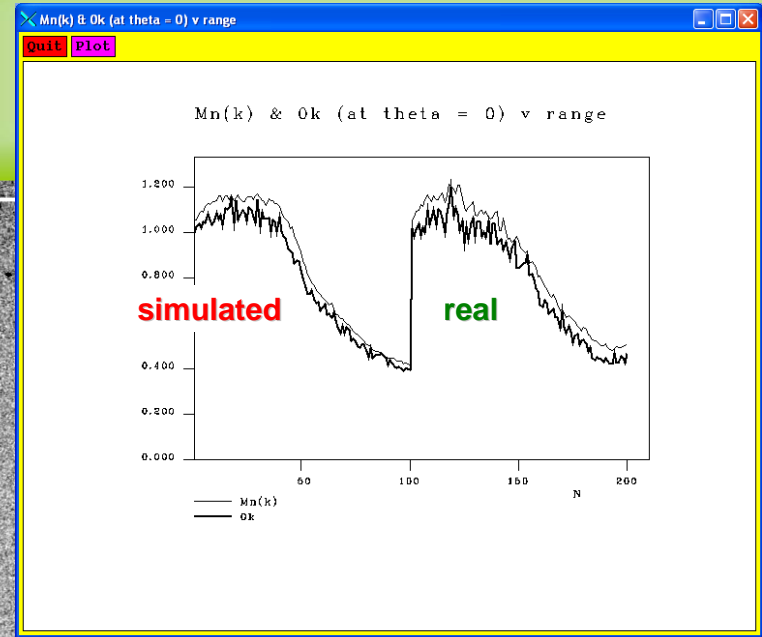
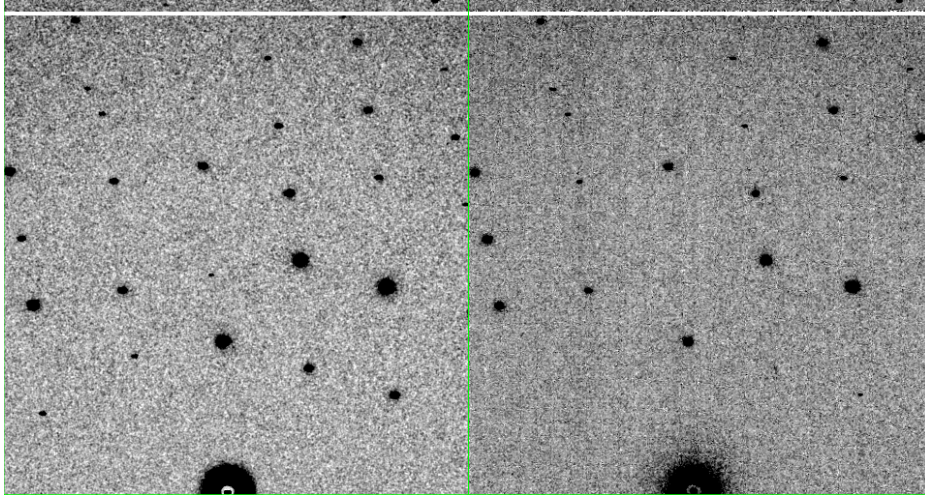
Simulated diffraction image
MLFSOM



Simulated diffraction image MLFSOM

simulated

real



Simulated

statistic

Real

5.4%	R_{merge}	6.2%
18.9	I/sd	16.2
1.9	I/sd (1.4 Å)	1.6
1.7 4.0 0.020	SDCORR	1.0 2.2 0.065
36.8	PADFPH	31.46
3.871	mlphare f''	3.476
0.178	FOM	0.192
0.545	FOMDM	0.664
0.6270	CC(1H87)	0.6090
0.170	R_{cryst}	0.184
0.205	R_{free}	0.231

Reference experiment

Beamline:

Crystal:

ALS 8.3.1

0.1 s exposures

total dose = 0.47 MGy

PDB id 1H87

tetragonal HEWL

soaked with
Gd complex (2 sites)

Detector:

Gd f'' at 12660 eV ~ Se

ADSC Quantum 315r

Bijvoet ratio: 3.8%

What goes into the simulation?

What goes into the simulation?

Crystal:

- “pristine” PDB
- “decayed” PDB
- Crystal size (μm)
- orientation matrix/angles
- mosaic spread (deg)
- absorption envelope (facet planes)

- Wavelength (\AA)
- dispersion ($\Delta\lambda/\lambda$)
- divergence (deg)
- flicker noise ($\%/\sqrt{\text{Hz}}$)
- shutter jitter (ms)

Detector:

- Distance, beam center
- tilt, twist, omega (deg)
- pixels x pixels, pixel size
- E-O gain, amplifier gain
- DQE, vignette
- read noise (rms e^-/pixel)
- “ripple” noise (%)
- point-spread function
- PSF variation

Non-crystal:

- “air” path (mm)
- cryo types & thicknesses (μm)
- ice rings

Beam:

- Flux (photons/s)
- beam size (μm)

What goes into the simulation?

What does **not** go into the simulation?

Crystal:

- “pristine” PDB
- “decayed” PDB
- Crystal size (μm)
- orientation matrix/angles
- mosaic spread (deg)
- absorption envelope (facet planes)

- Wavelength (\AA)
- dispersion ($\Delta\lambda/\lambda$)
- divergence (deg)
- flicker noise ($\%/\sqrt{\text{Hz}}$)
- shutter jitter (ms)

Detector:

- Distance, beam center
- tilt, twist, omega (deg)
- pixels x pixels, pixel size
- E-O gain, amplifier gain
- DQE, vignette
- read noise (rms e^-/pixel)
- “ripple” noise (%)
- point-spread function
- PSF variation

Non-crystal:

- “air” path (mm)
- cryo types & thicknesses (μm)
- ice rings

Beam:

- Flux (photons/s)
- beam size (μm)

Crystal:

- “one dimensional” (but variable)
- uniform damage
- assumed “top hat” mosaic spread
- kinematical diffraction

Beam:

- assumed uniform square “top hat” profiles for beam size, divergence, dispersion
- flicker and jitter follow normal distribution

Non-crystal:

- no polarization correction
- no 2θ support

Detector:

- no spatial transformation
- no flood/dark correction
- no read-noise anisotropy
- no electron-counting noise

Spot Shape

Arndt & Wonacott (1977)



Oscillation photograph of Tobacco Mosaic Virus protein. Crystal rotated by 0.9° ($P2_1, a = 226 \text{ \AA}, b = 224 \text{ \AA}, c = 174 \text{ \AA}$).

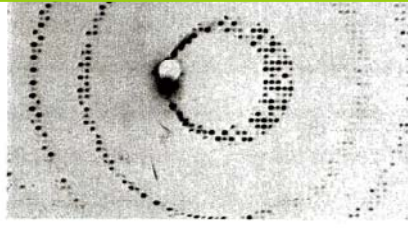


Fig. 2.4. Enlarged central portion of the pattern shown in Fig. 2.3 to show partially recorded reflexions at the edges of the lunes.

2.8. The Torus of Reflexion

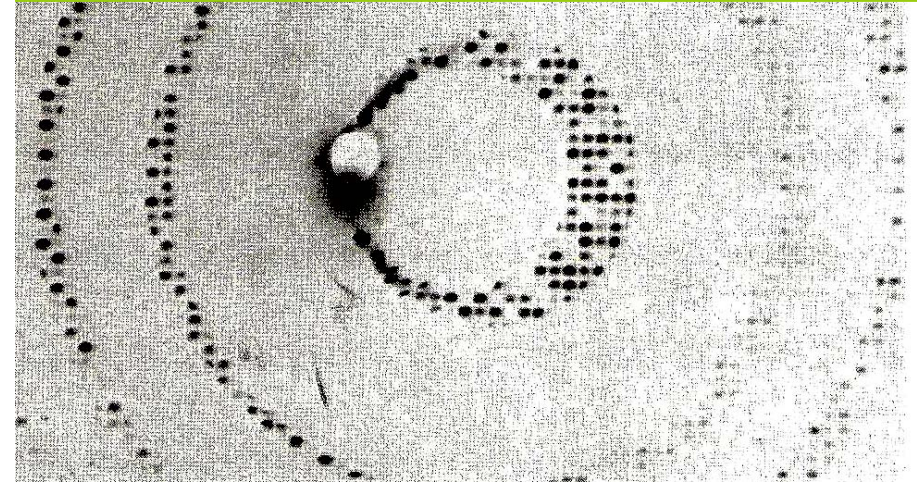
It is important to consider which reciprocal lattice points come into the sphere and which remain outside the sphere during one complete rotation of the crystal about the oscillation axis. Fig. 2.6 shows the "torus" of reflexions swept out by a complete rotation of the sphere of reflexion, the reciprocal lattice remaining stationary. For protein crystals, where the resolution limit is normally about 2.9 \AA and, at best, 1.5 \AA , only part of the "torus" is accessible and the sphere centred on O represents the d^* limit. The region that is inaccessible has to be investigated by rotation about another axis is a cusp centred on the oscillation axis. For 3 \AA data this volume amounts to only 2% of the total volume within the d^* limiting sphere (see Fig. 7.5).

It should be noted that blind regions in screenless oscillation photographs cannot be avoided by inclining the crystal axis to the X-ray beam. In inclined beam photography only two reciprocal lattice layers, the "equi-inclination"

ned by the intersection of the Ewald sphere with a particular reciprocal level, $l = 0, 1, 2, 3$ etc., at the beginning and the end of the oscillation. particular photograph the lunes are highly pronounced because of the density of reciprocal lattice points and because the levels are nearly parallel to the incident X-ray beam. Often the lunes are less obvious, or not lunes are interlaced, as in Figs. 2.5c and d. ally recorded reflexions can be seen on the edges of the lunes in Fig. d, more clearly in Fig. 2.4. Such reflexions at the edges of the lunes are out radially, but when a reflexion has passed through the sphere in a on perpendicular to its surface the fact that it has been cut is not obvious by inspection. It is often only possible to identify which of a recorded area correspond to the start and which to the end of a on by comparing the photograph with adjacent ones. lunes can often be followed clearly on a series of adjacent photo-

Spot Shape

Arndt & Wonacott (1977)



Spot shape

Ewald sphere

(h,k,l)

Spot shape

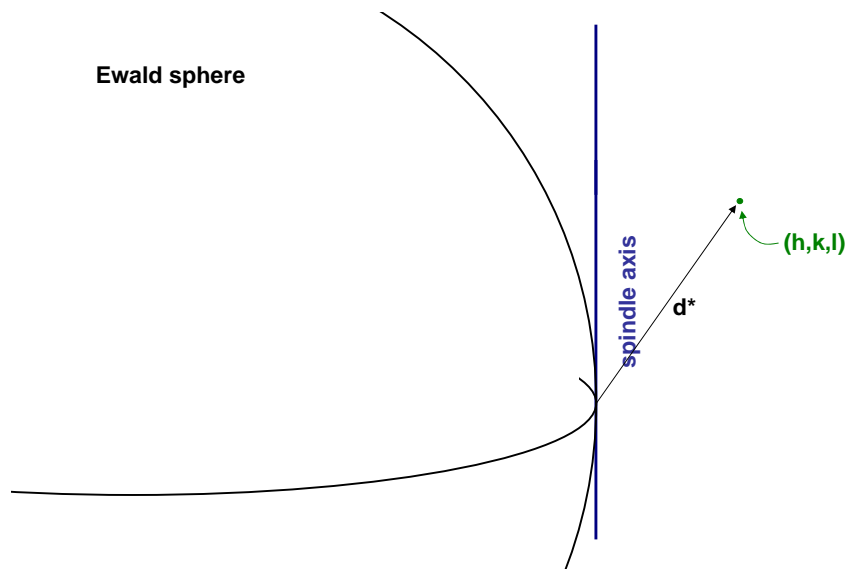
Ewald sphere

d^*

(h,k,l)

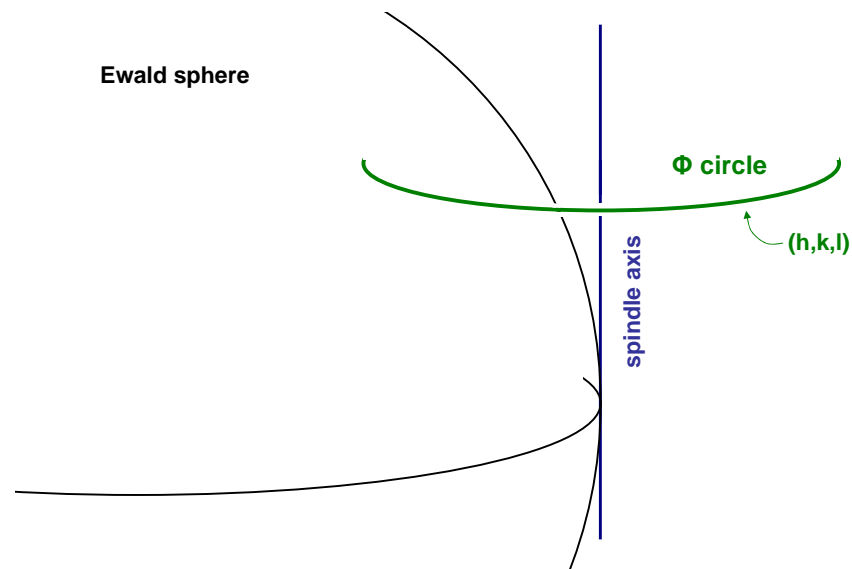
Spot shape

Ewald sphere



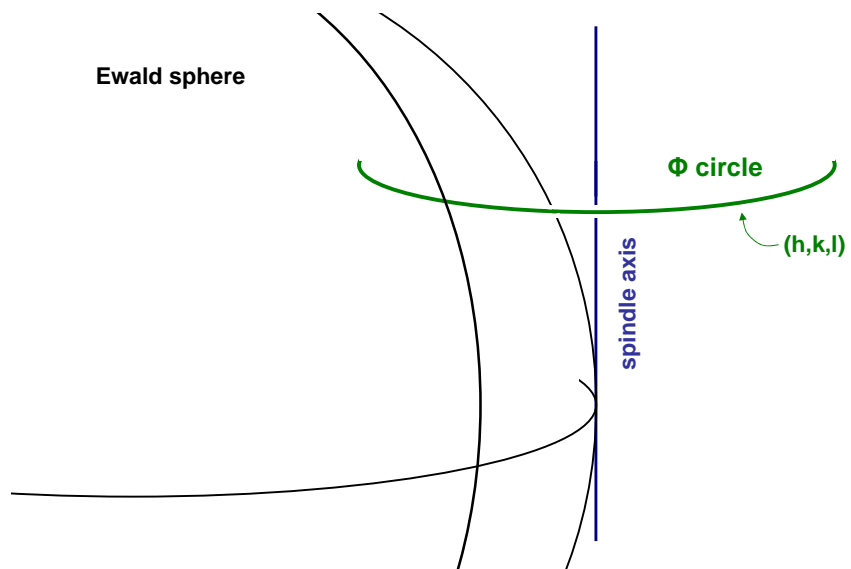
Spot shape

Ewald sphere



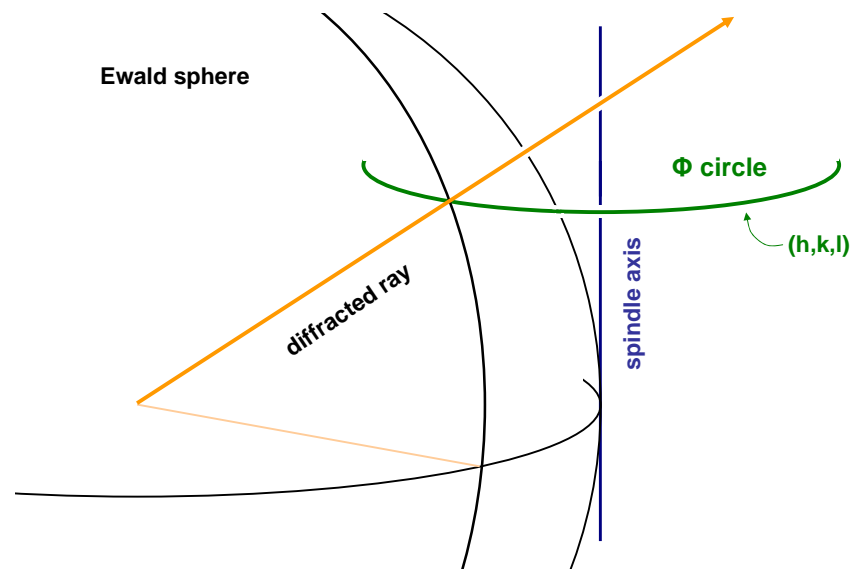
Spot shape

Ewald sphere

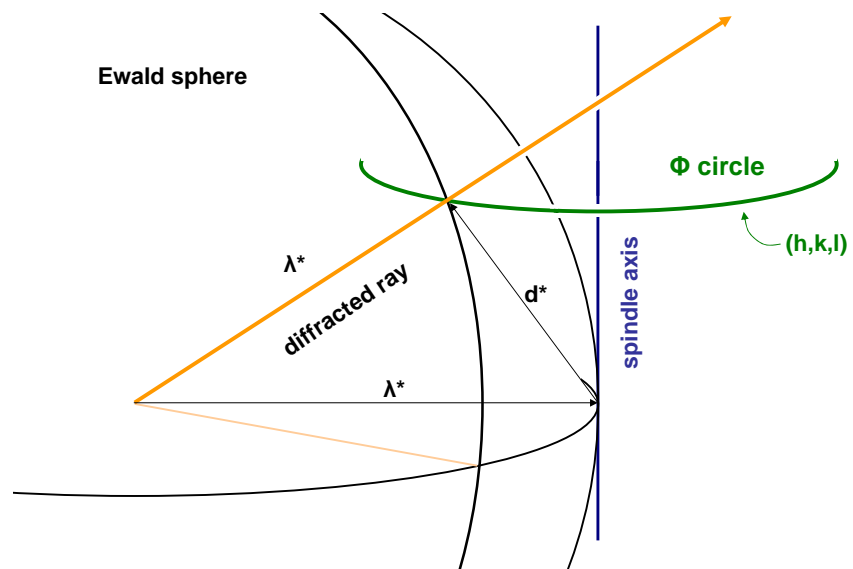


Spot shape

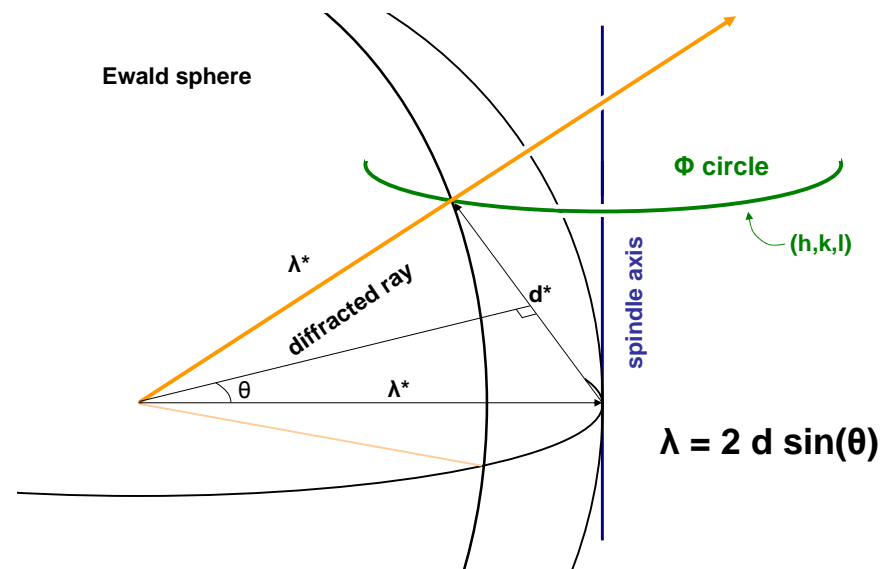
Ewald sphere



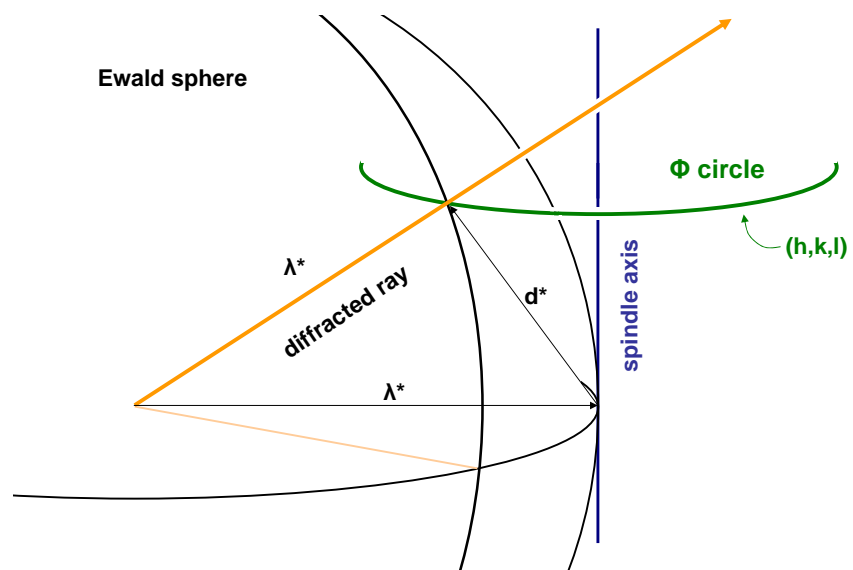
Spot shape



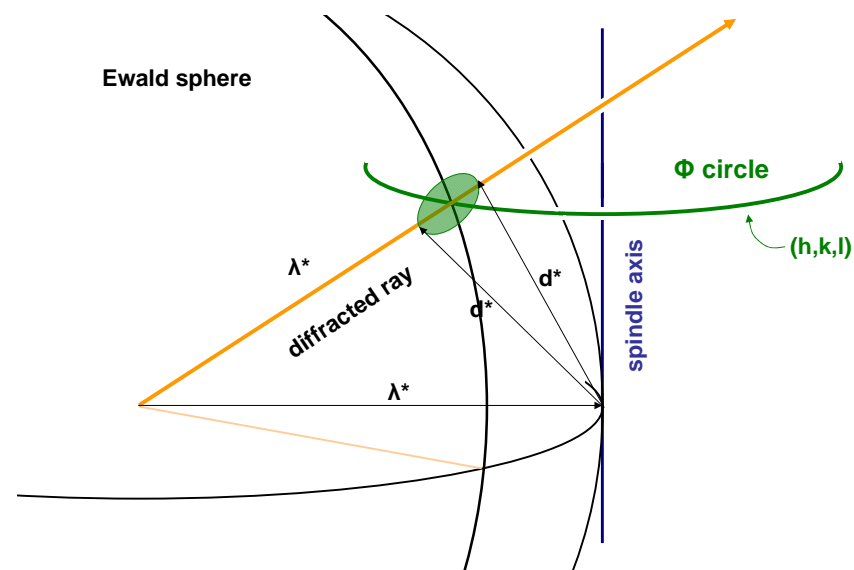
Spot shape



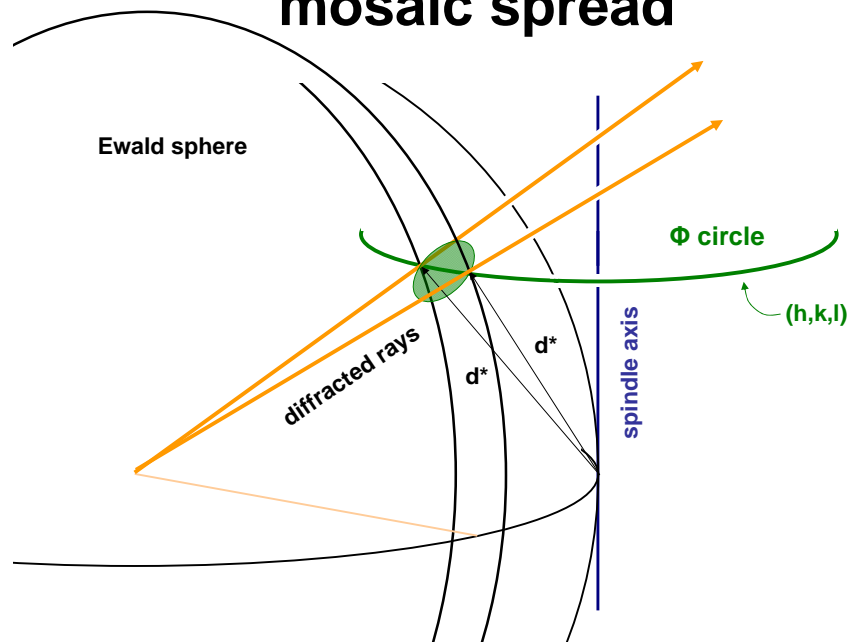
Spot shape



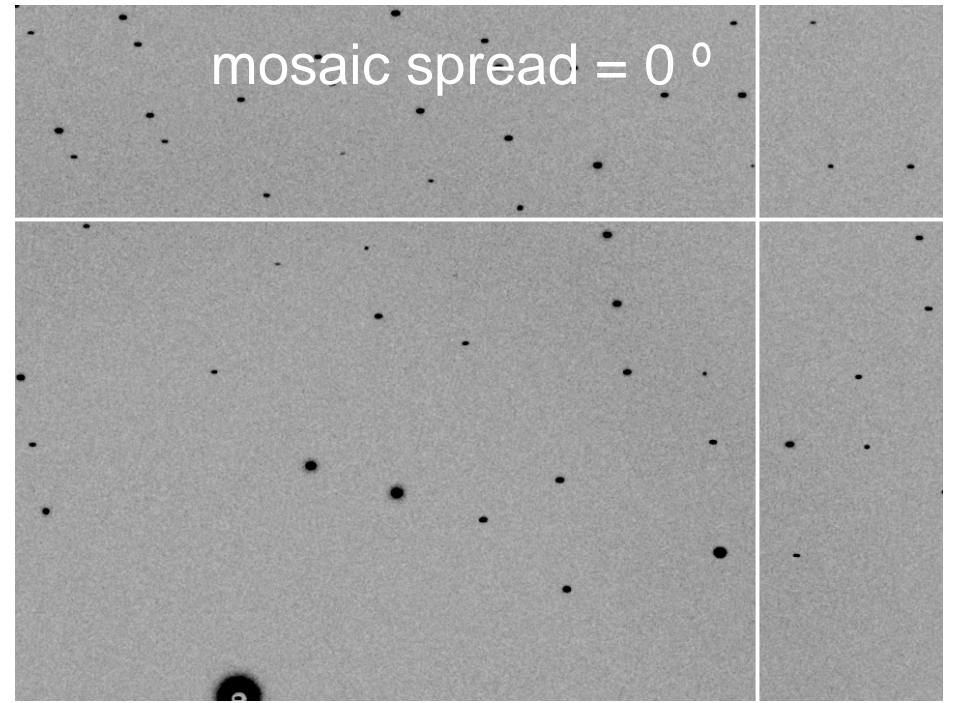
mosaic spread



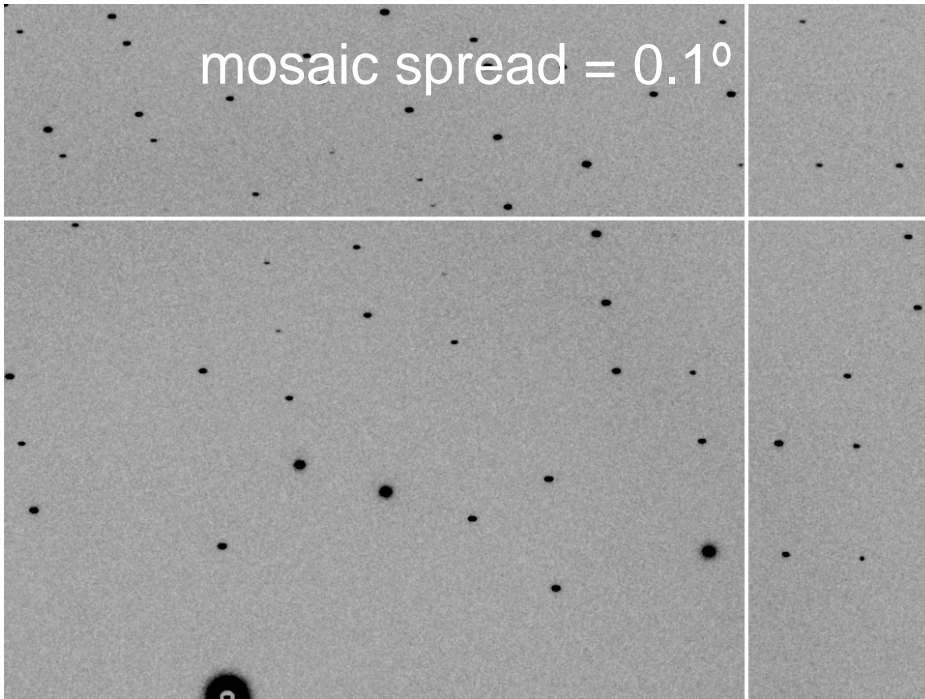
mosaic spread



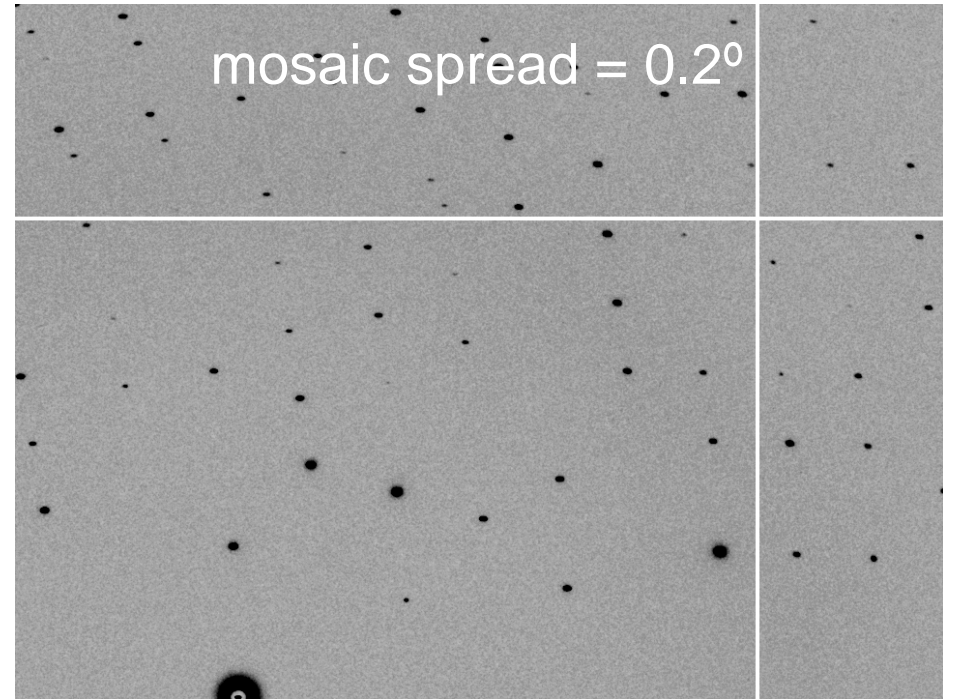
mosaic spread = 0°



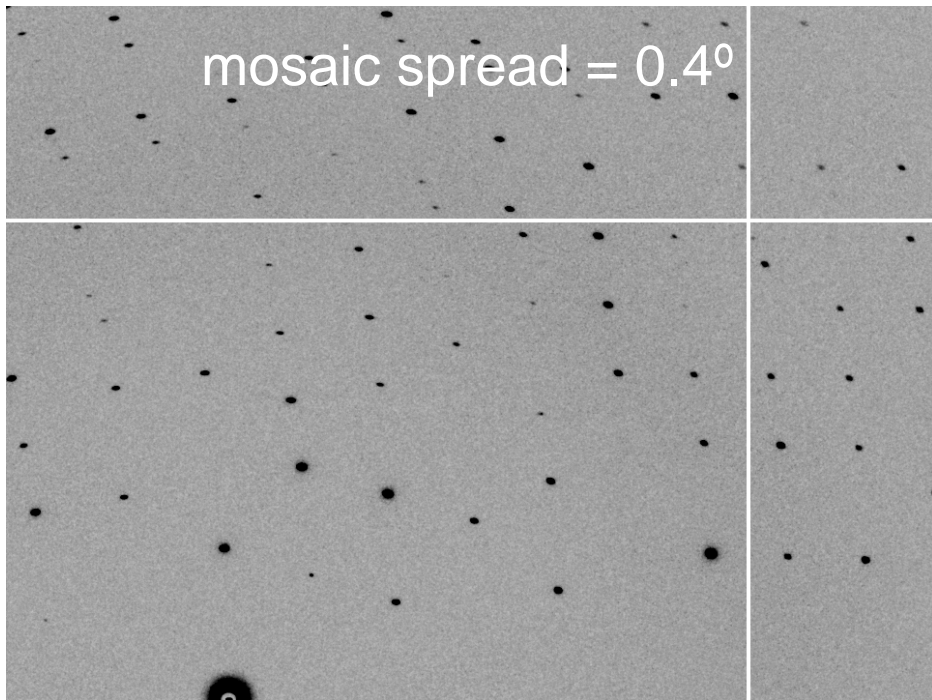
mosaic spread = 0.1°



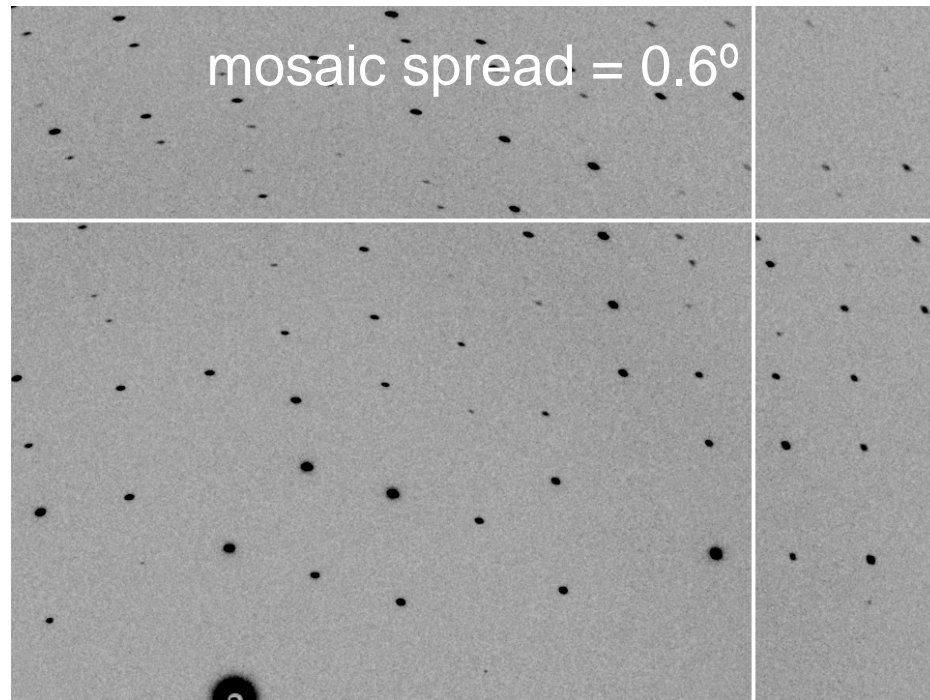
mosaic spread = 0.2°



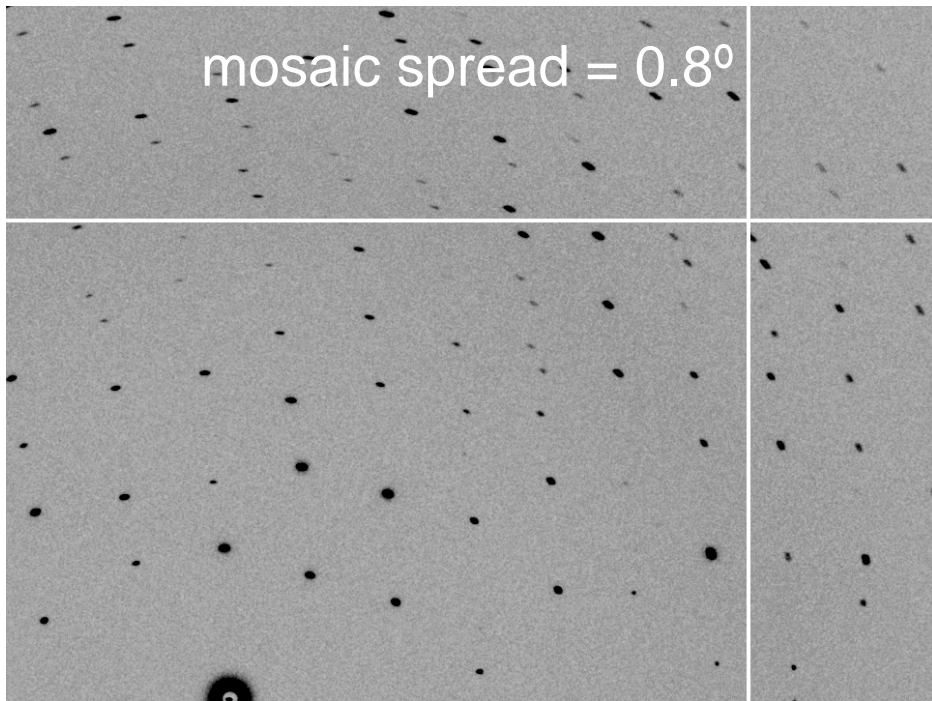
mosaic spread = 0.4°



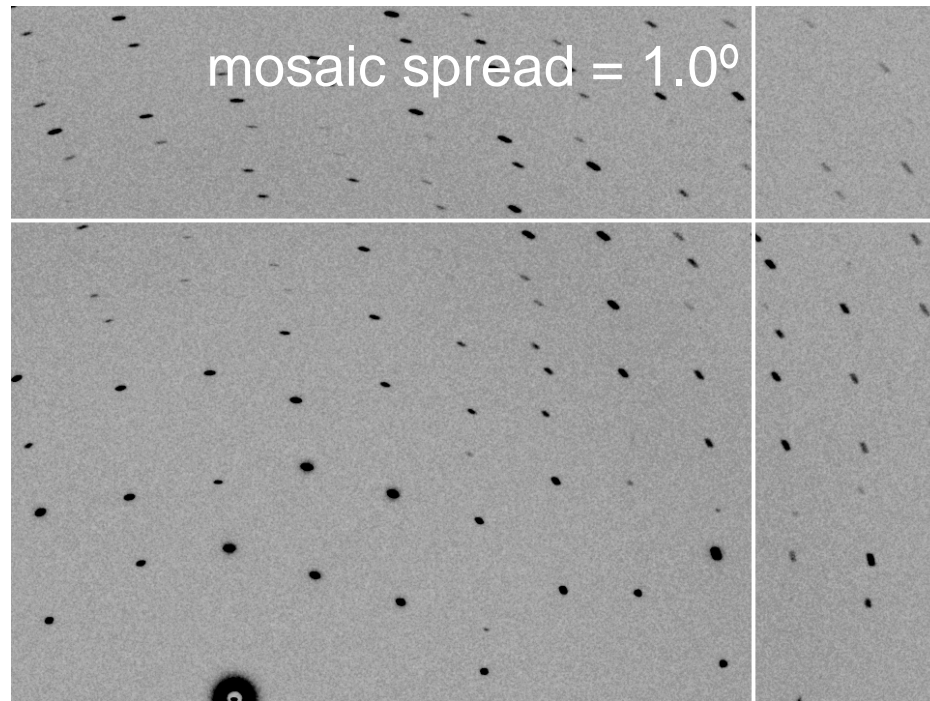
mosaic spread = 0.6°



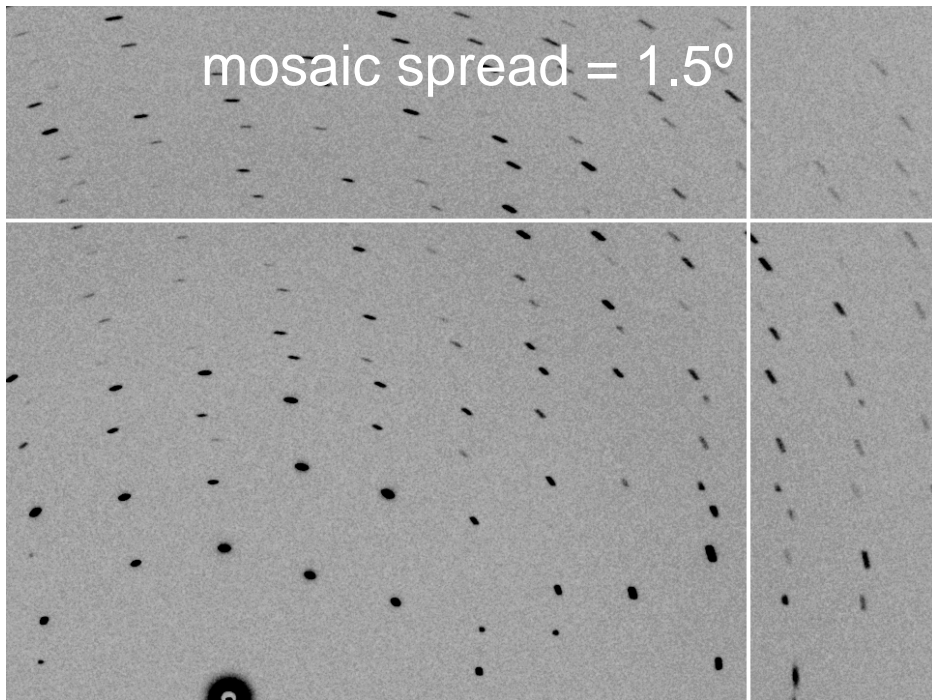
mosaic spread = 0.8°



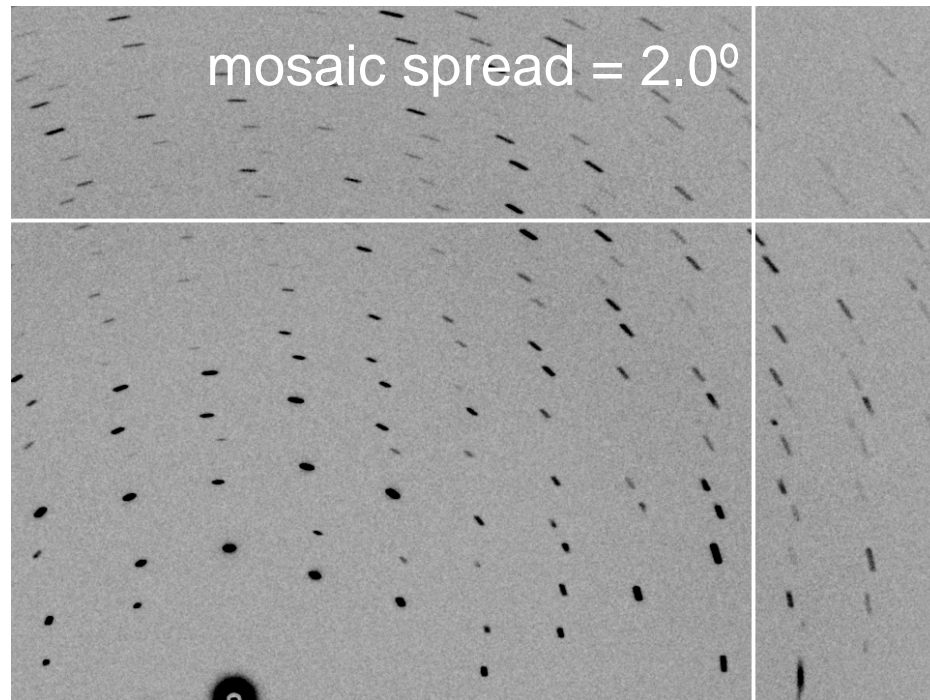
mosaic spread = 1.0°



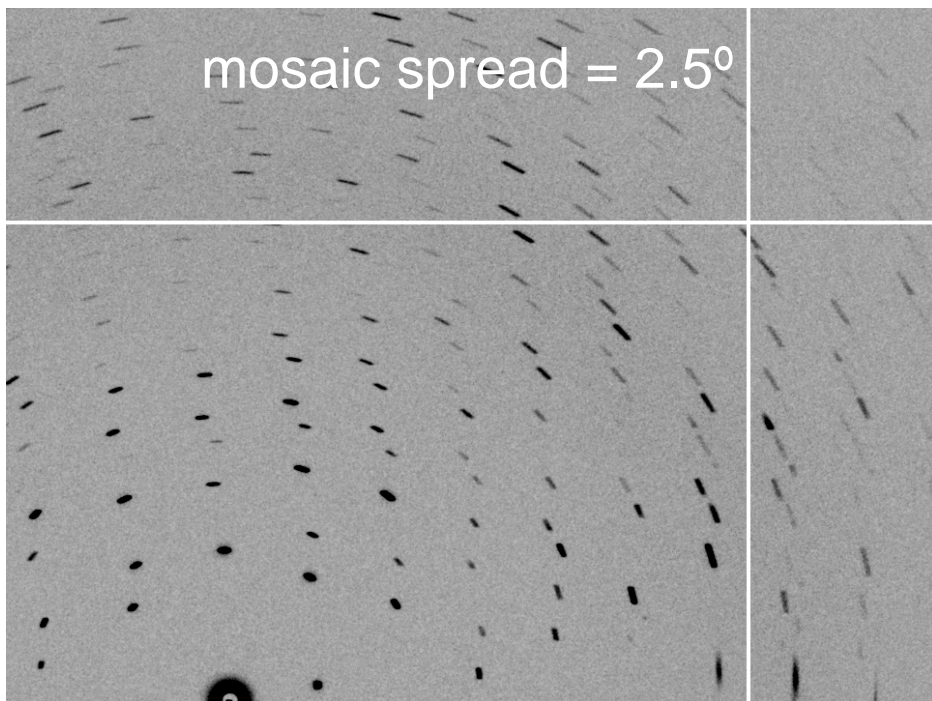
mosaic spread = 1.5°



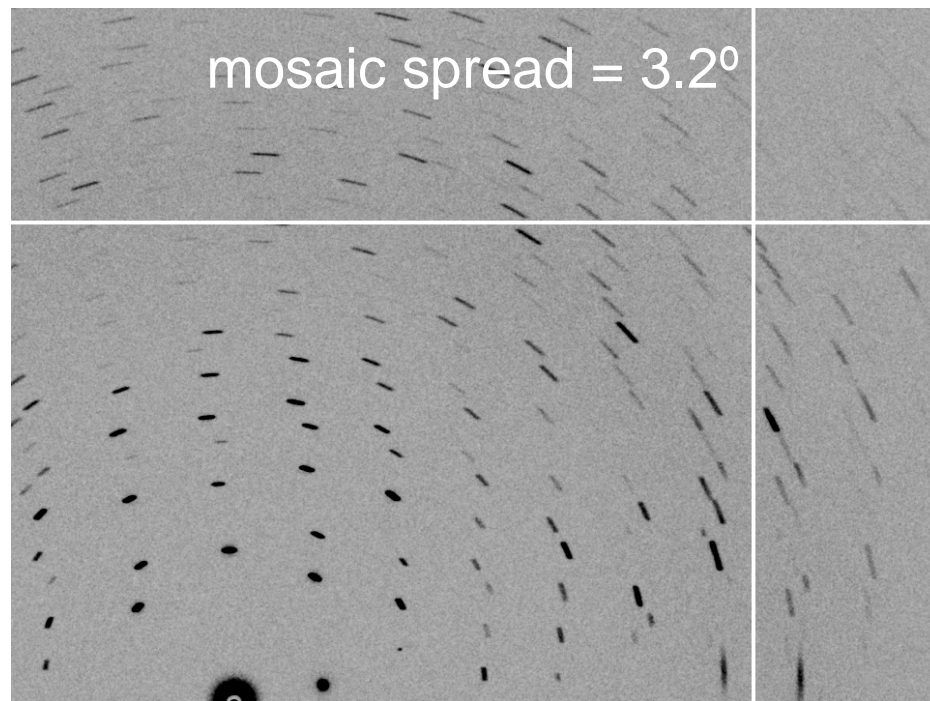
mosaic spread = 2.0°



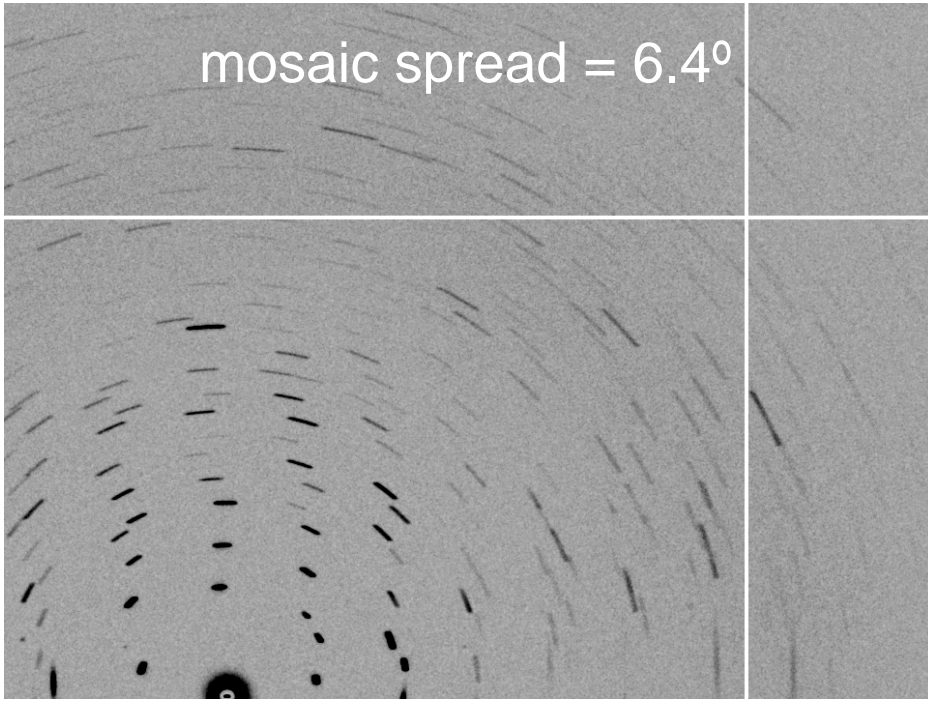
mosaic spread = 2.5°



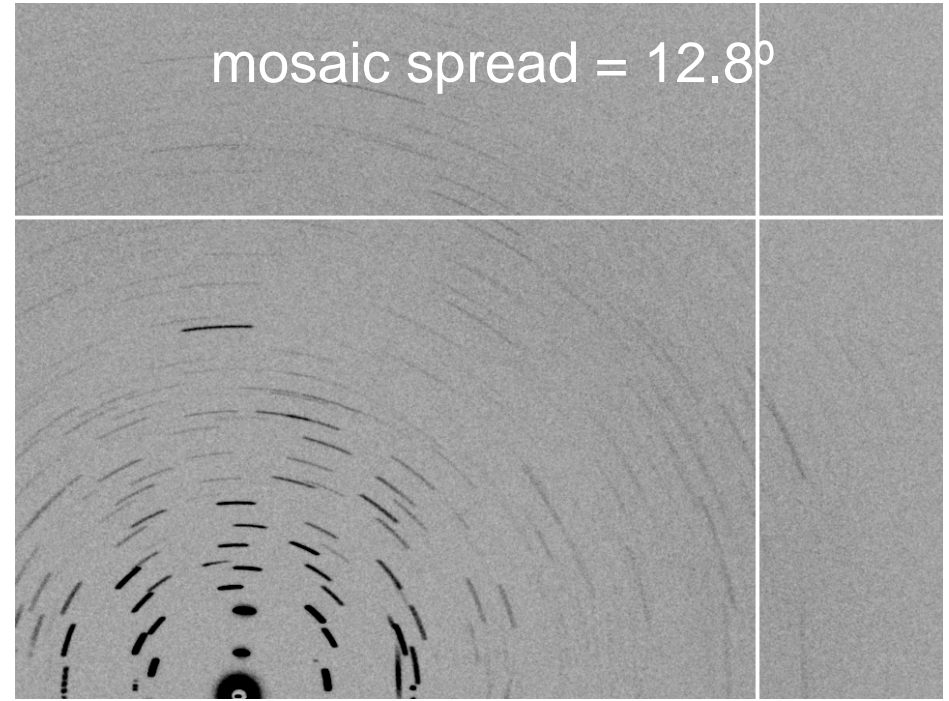
mosaic spread = 3.2°



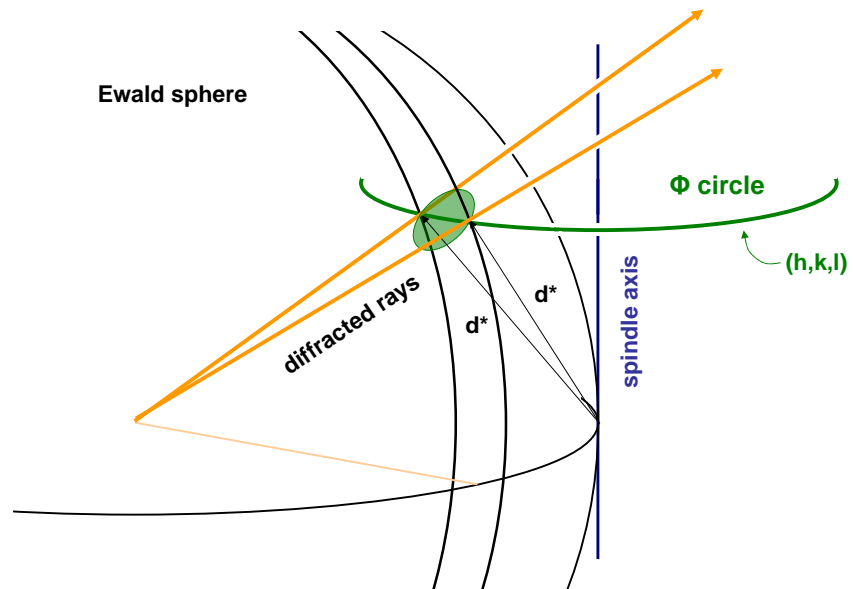
mosaic spread = 6.4°



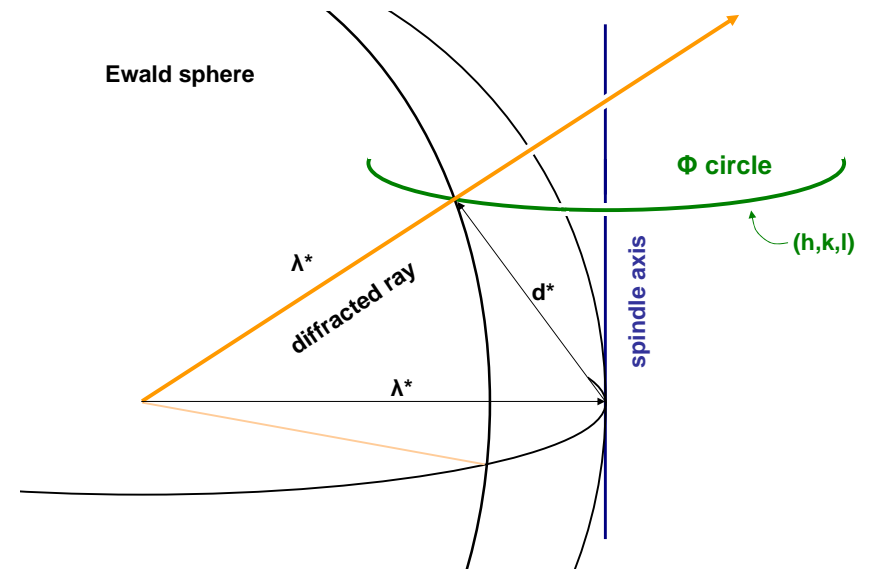
mosaic spread = 12.8°



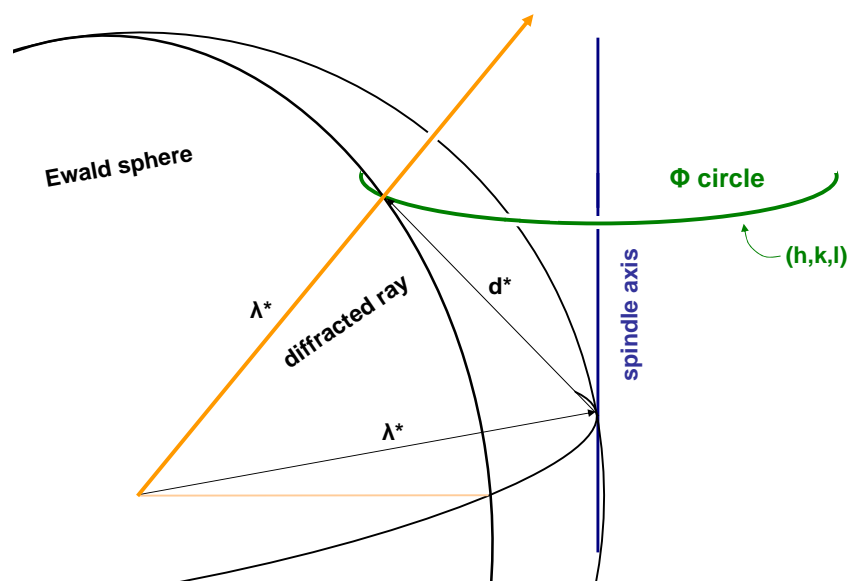
mosaic spread



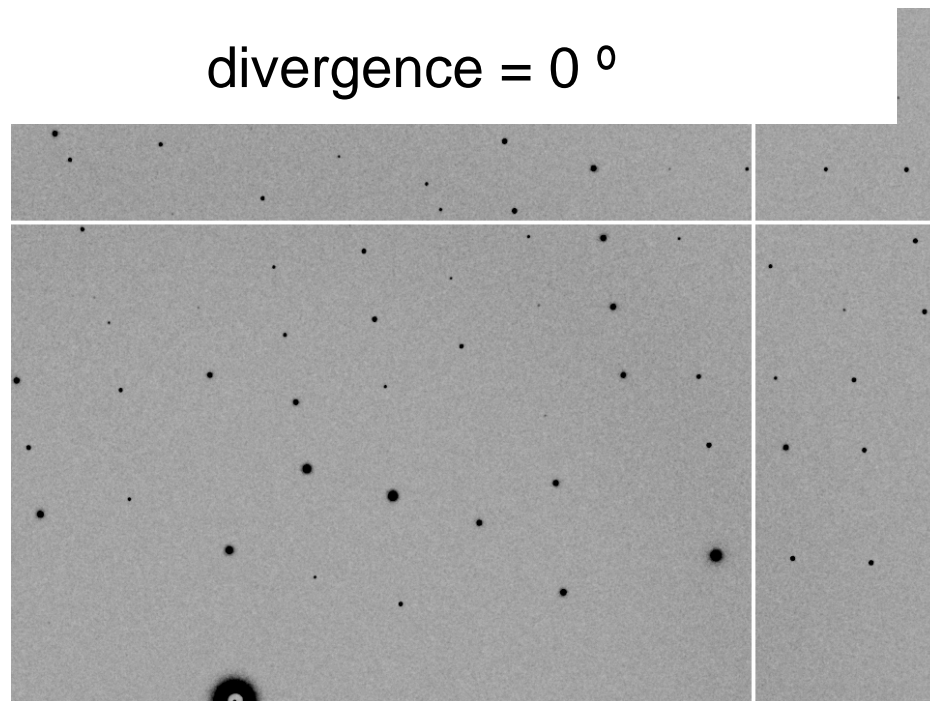
beam divergence



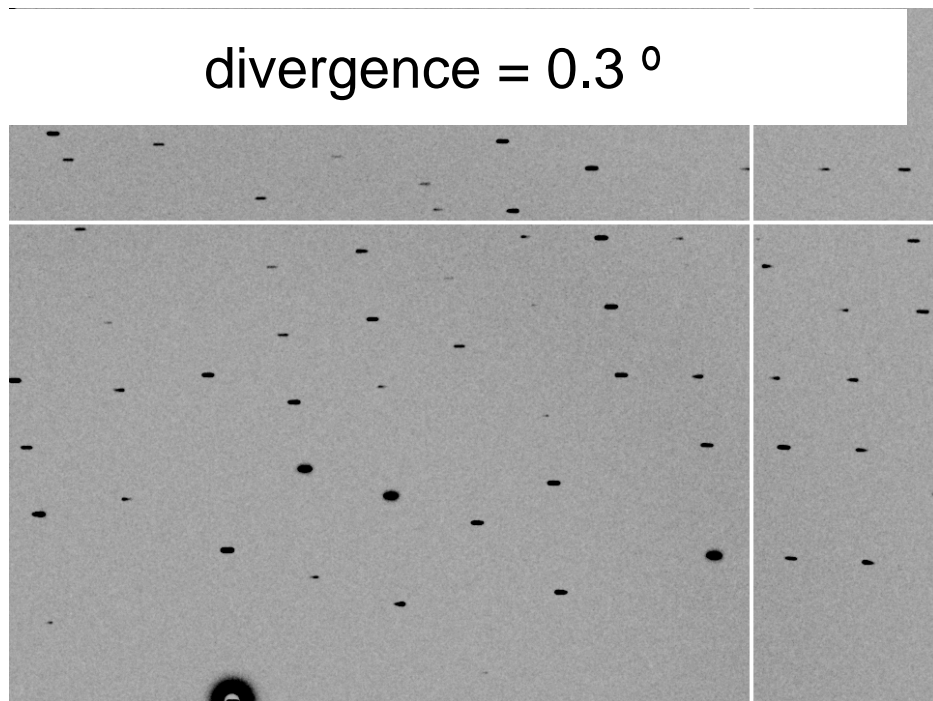
beam divergence



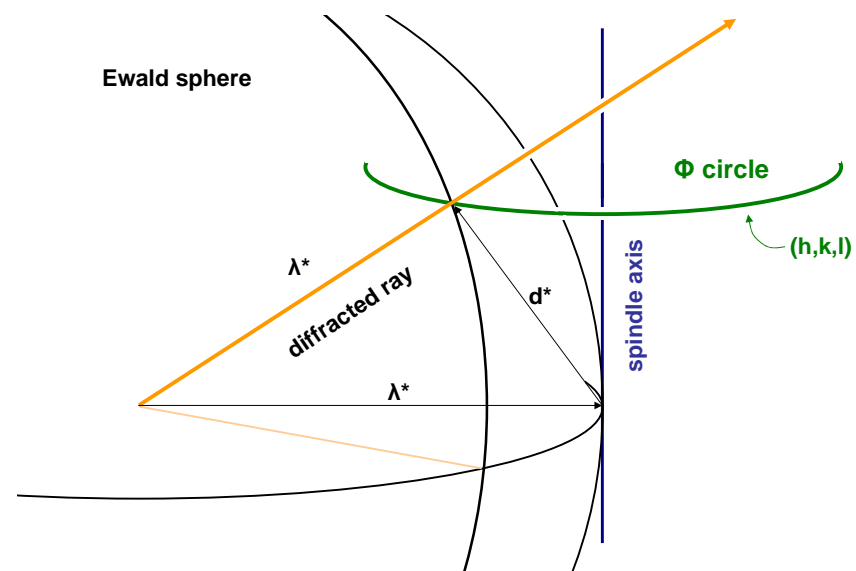
divergence = 0°



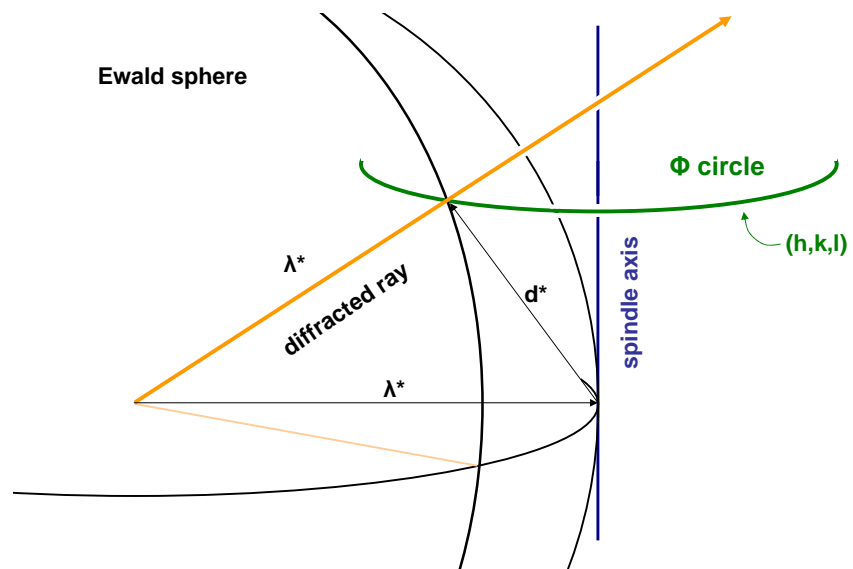
divergence = 0.3°



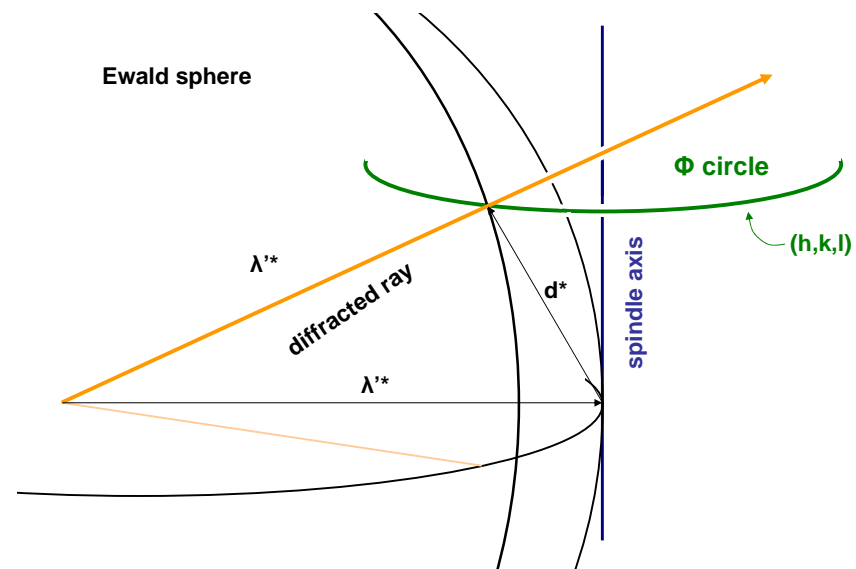
beam divergence



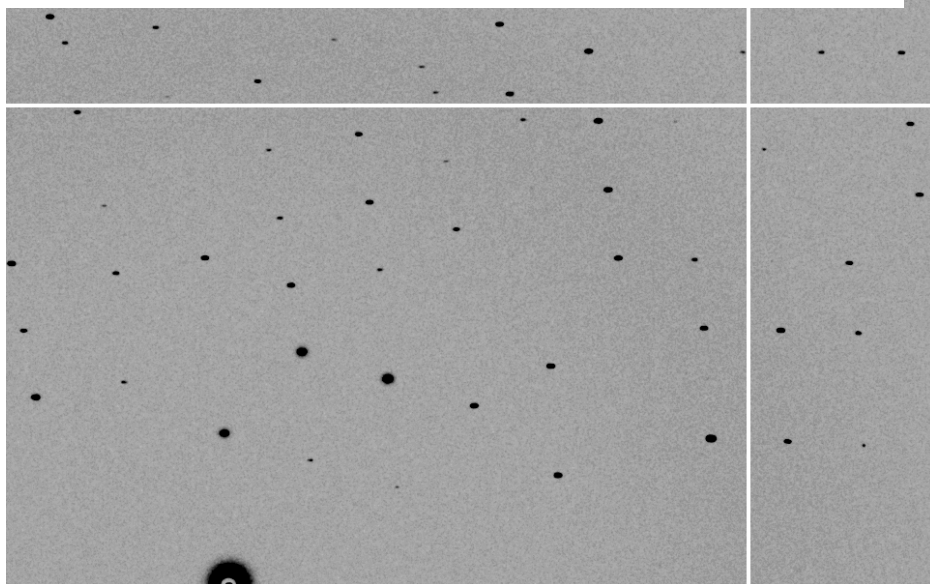
spectral dispersion



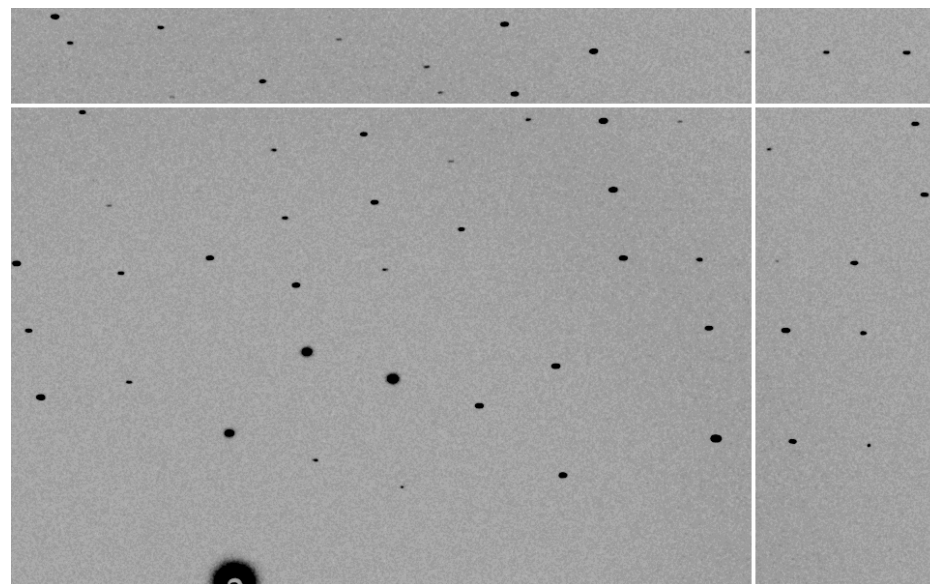
spectral dispersion



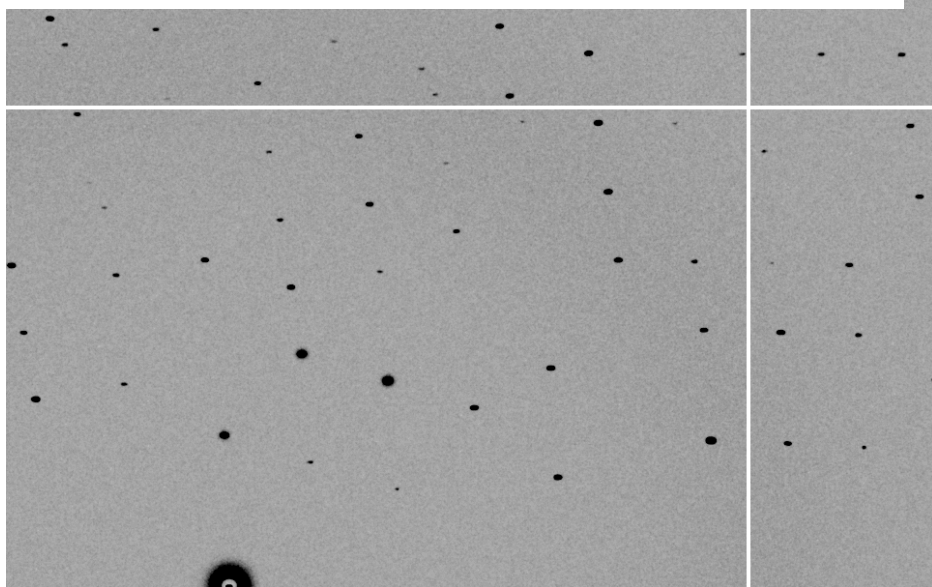
dispersion = 0



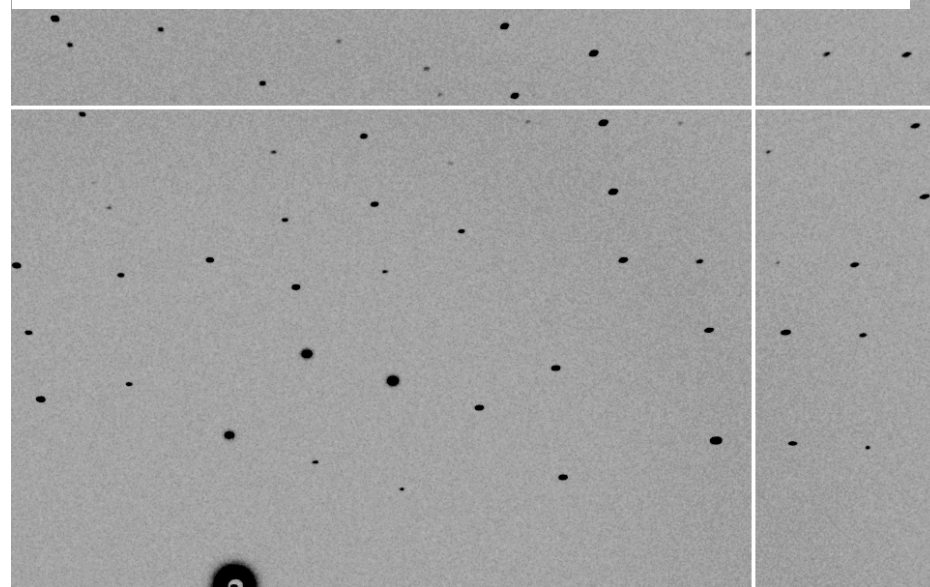
dispersion = 0.014% (Si 111)



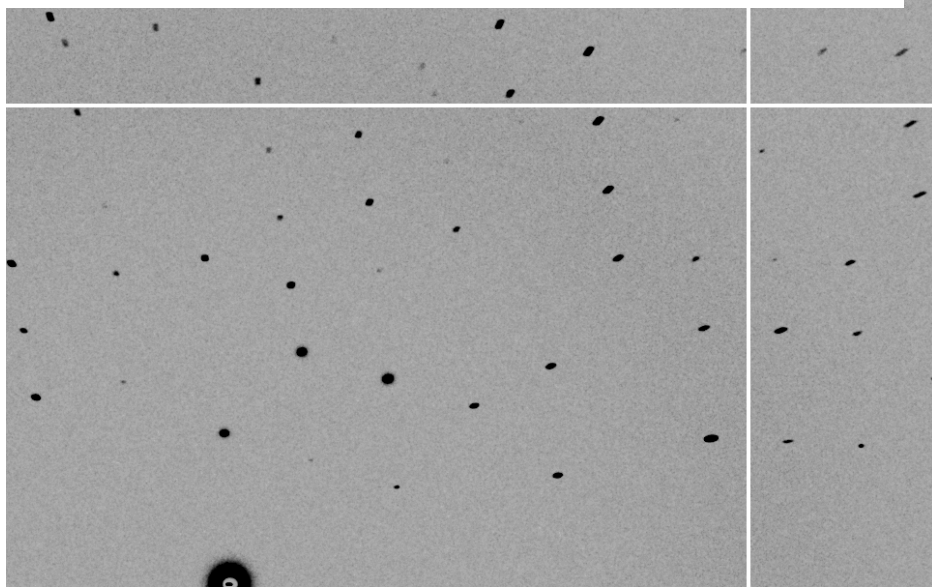
dispersion = 0.25% (CuK $_{\alpha}$)



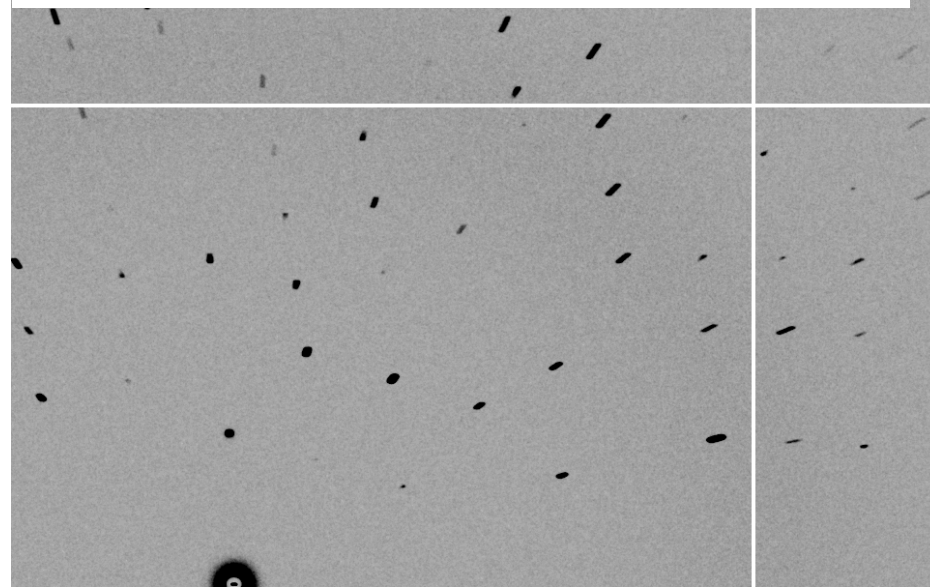
dispersion = 0.6%



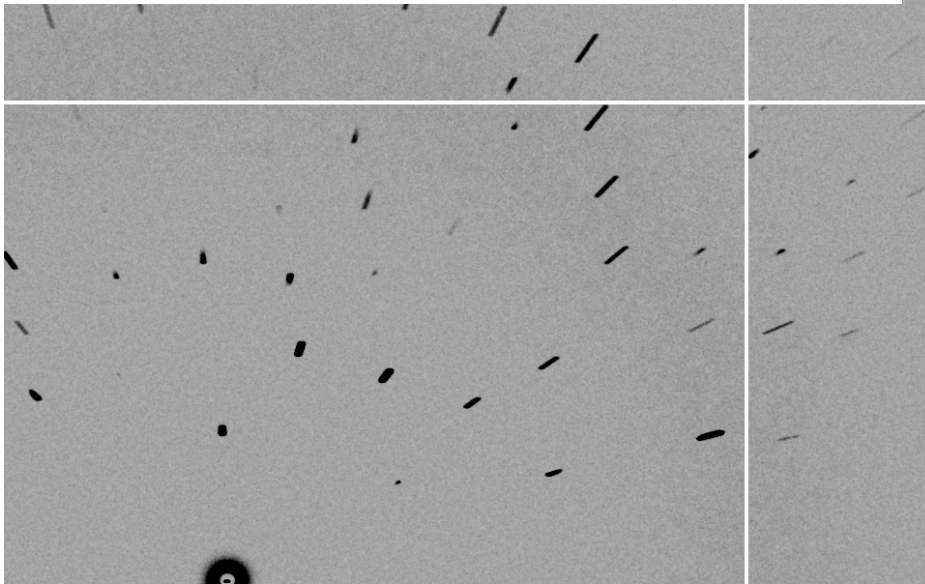
dispersion = 1.3%



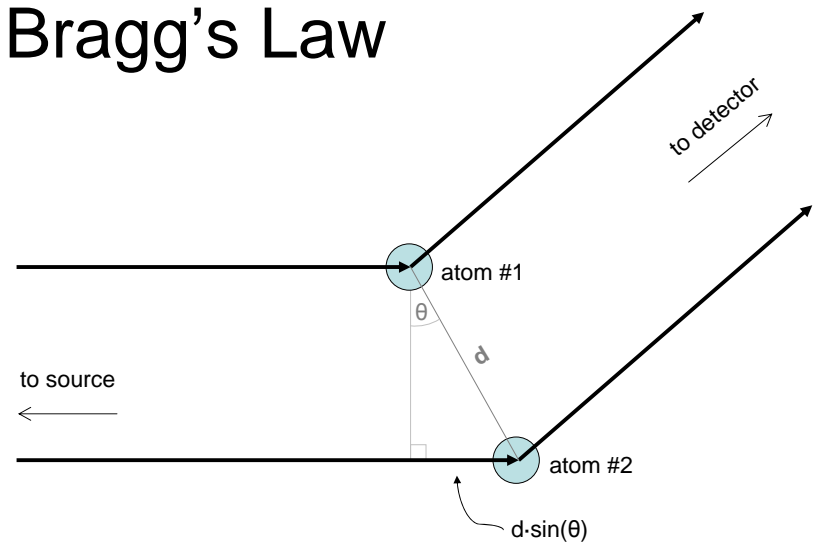
dispersion = 2.6%



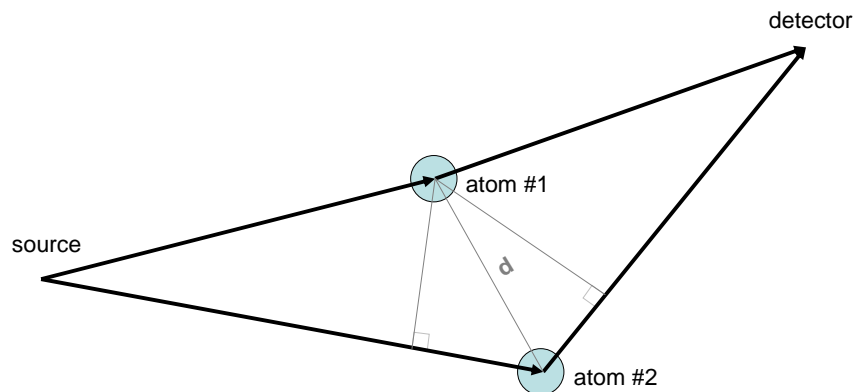
dispersion = 5.1%



Bragg's Law



“near”-ly Bragg's Law



mosaicity vs Wilson B

INTENSITY OF REFLECTION OF X-RAYS BY CRYSTALS 35

(b) *The amplitude reflected by a plane sheet of atoms :* We shall first consider the amplitude of the wave reflected by an infinite plane sheet of atoms, each of which scatters the incident X-rays.

Suppose A, fig. 15, is the source of the radiation, and let the amplitude of the reflected wave be required at B. Let the plane APB be normal to the plane of atoms, and let AP, PB make equal angles θ with this plane. Then P is such that the distance APB is the shortest distance from A to B via the plane. Let M be a point of the plane such that the distance

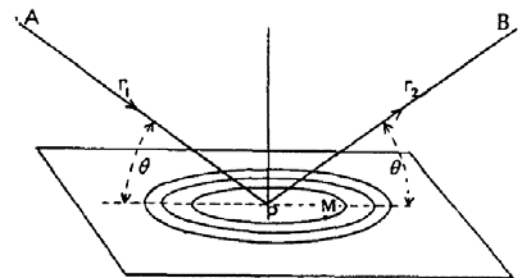
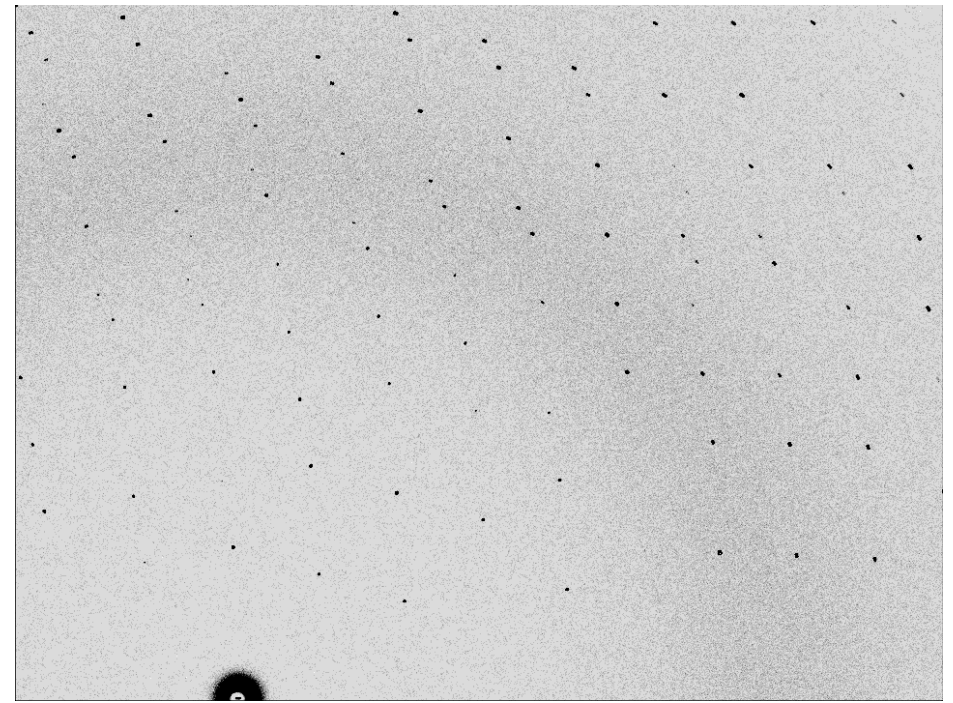
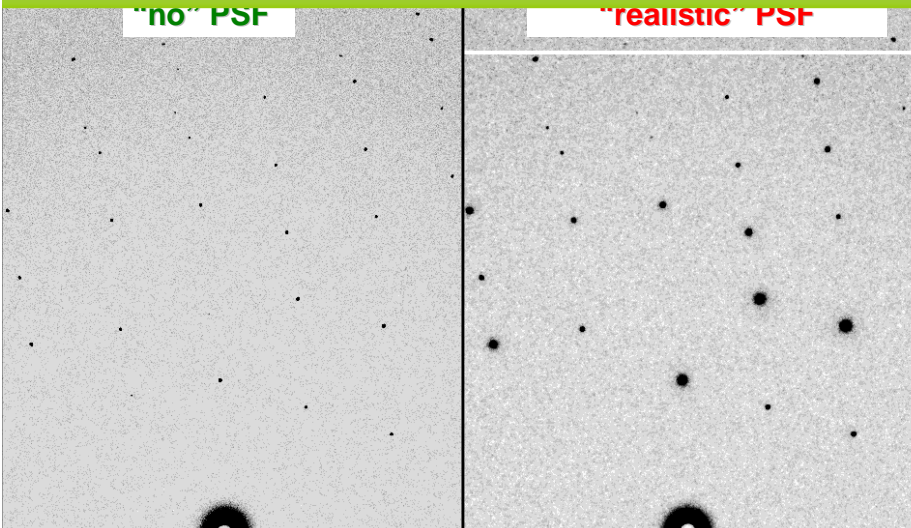


FIG. 15

The diagram illustrates the Ewald sphere and the Φ circle in reciprocal space. The Ewald sphere is shown as a series of concentric arcs. The Φ circle is a green ellipse. The spindle axis is a vertical blue line. The diagram shows the intersection of the Ewald sphere and the Φ circle, with labels for λ^* , λ' , λ'' , λ''' and d^* , d' , d'' , d''' .

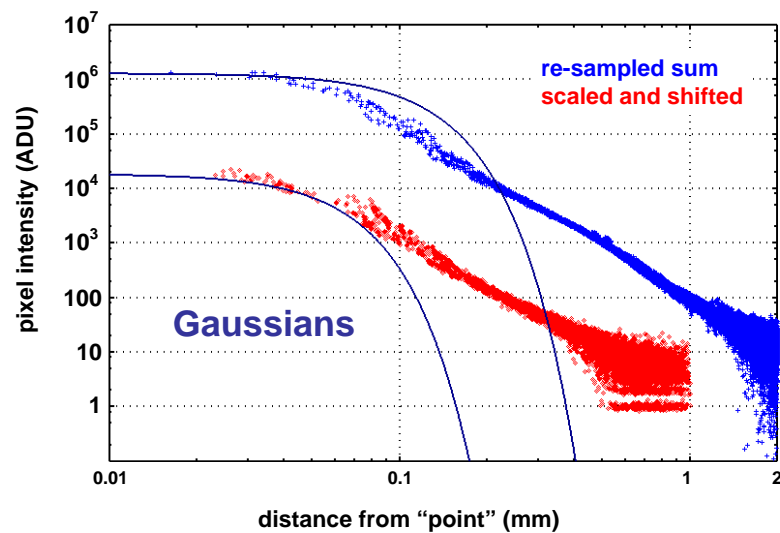


Point Spread Function

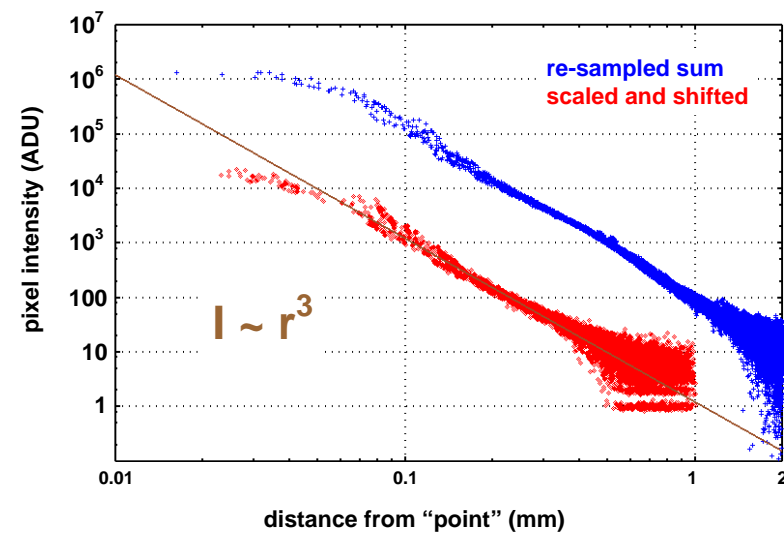


The image is a side-by-side comparison of two grayscale astronomical images. The left image is labeled "no PSF" in green text on a white background. It shows a field of stars where each star is represented by a single, sharp black pixel. The right image is labeled "realistic PSF" in red text on a white background. It shows the same field of stars, but each star is represented by a small, dark, circular disk (the Point Spread Function) instead of a single pixel. This illustrates how a realistic PSF blurs the light from each star.

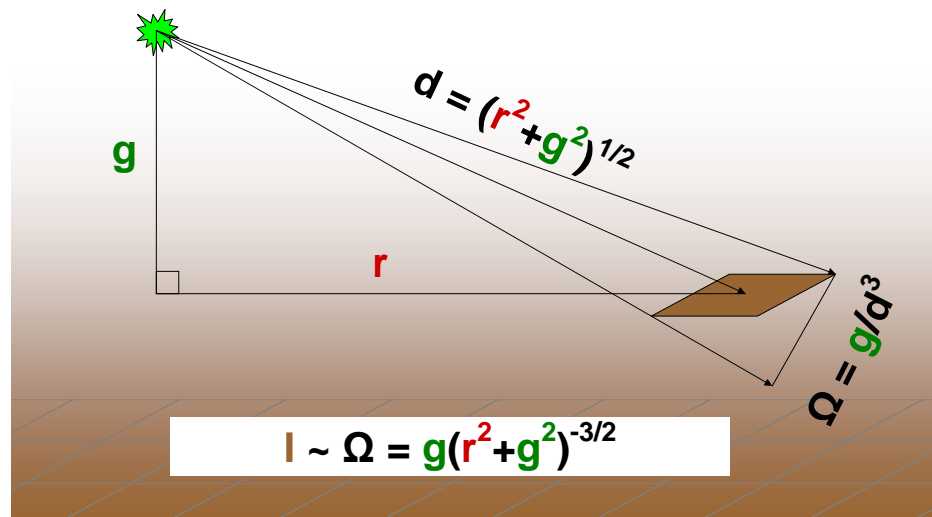
Point Spread Function



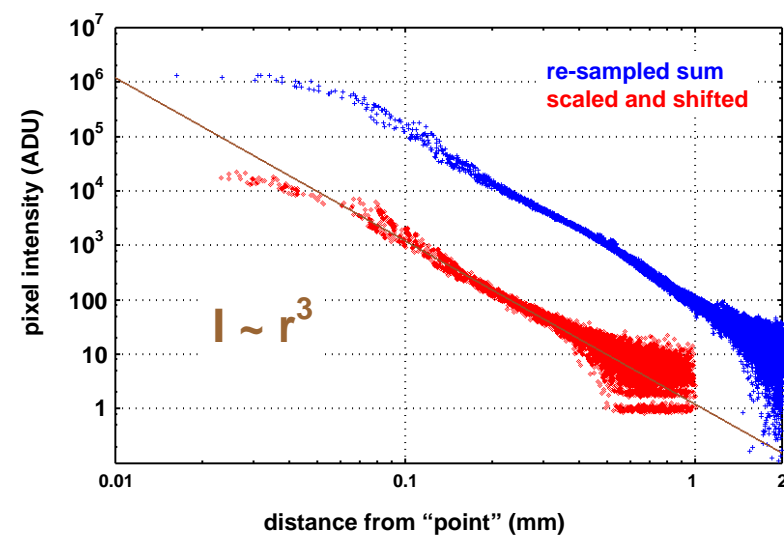
Point Spread Function



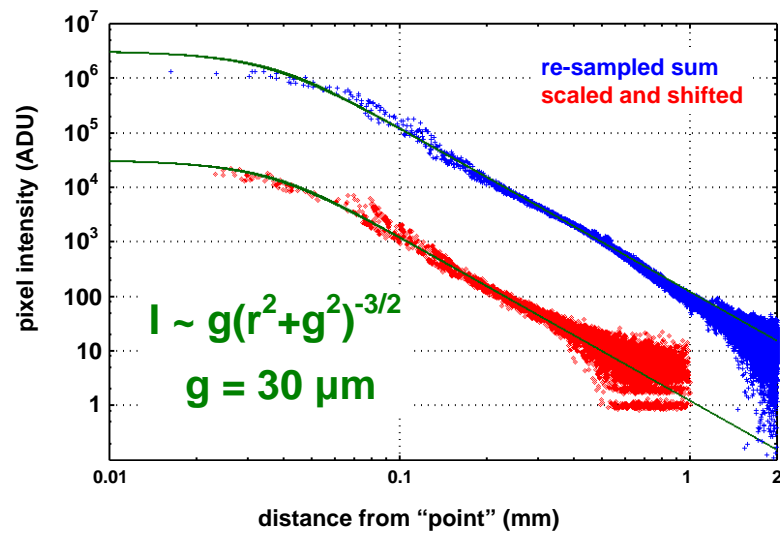
Point Spread Function



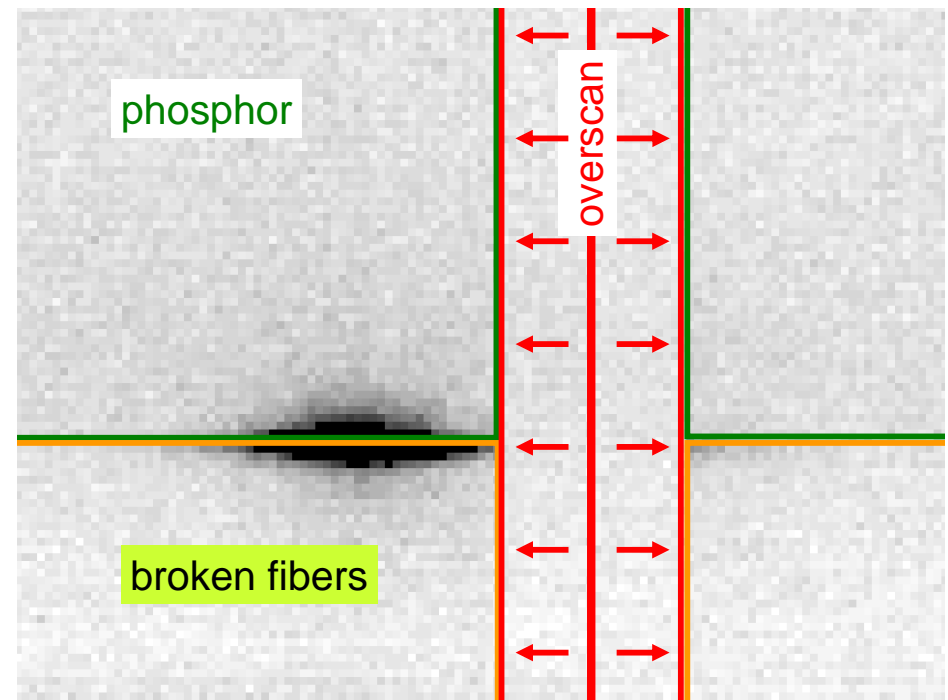
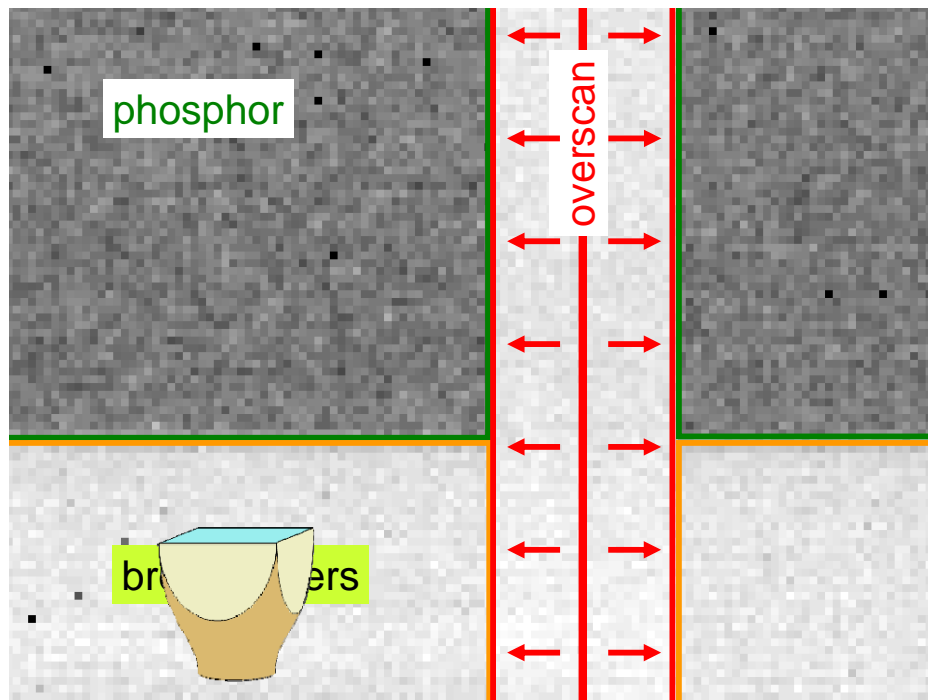
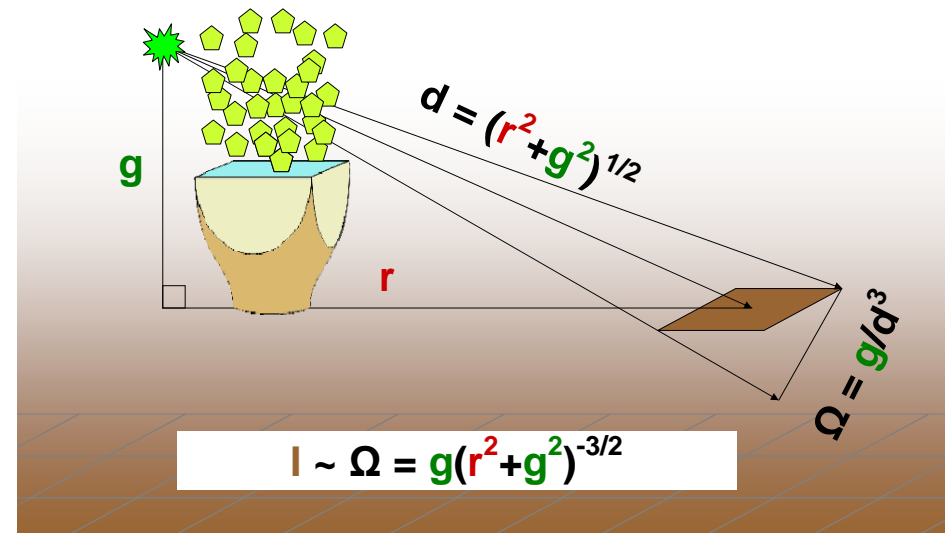
Point Spread Function

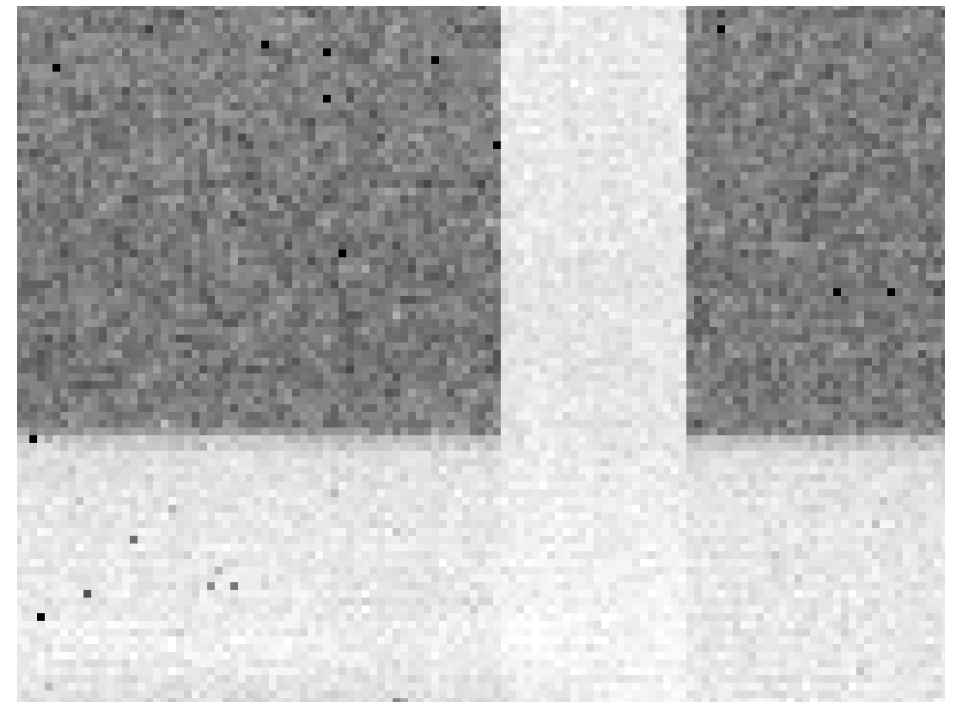
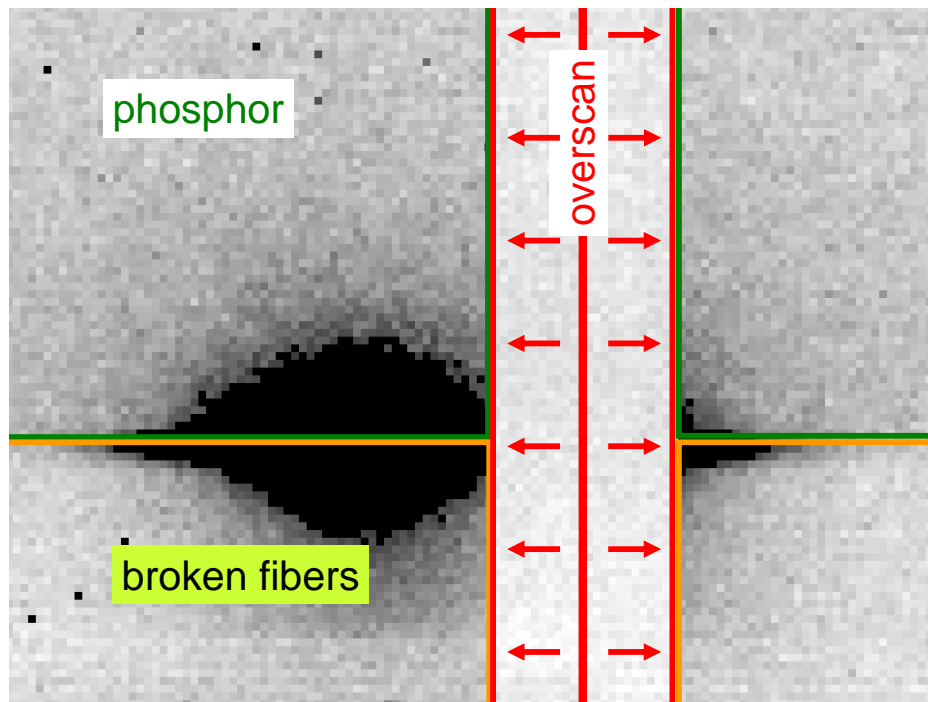


Point Spread Function



Point Spread Function





Intensity of a Bragg spot

$$I_{||} \approx |F(hkl)|^2$$

Darwin's Formula

$$I(hkl) = I_{\text{beam}} r_e^2 \frac{V_{\text{xtal}}}{V_{\text{cell}}} \frac{\lambda^3 L}{\omega V_{\text{cell}}} P A |F(hkl)|^2$$

$I(hkl)$	- photons/spot (fully-recorded)	ω	- rotation speed (radians/s)
I_{beam}	- incident (photons/s/m ²)	L	- Lorentz factor (speed/speed)
r_e	- classical electron radius (2.818x10 ⁻¹⁵ m)	P	- polarization factor $(1+\cos^2(2\theta) - P_{\text{fac}} \cdot \cos(2\Phi) \sin^2(2\theta))/2$
V_{xtal}	- volume of crystal (in m ³)	A	- absorption factor $\exp(-\mu_{\text{xtal}} \cdot l_{\text{path}})$
V_{cell}	- volume of unit cell (in m ³)	$F(hkl)$	- structure amplitude (electrons)
λ	- x-ray wavelength (in meters!)		

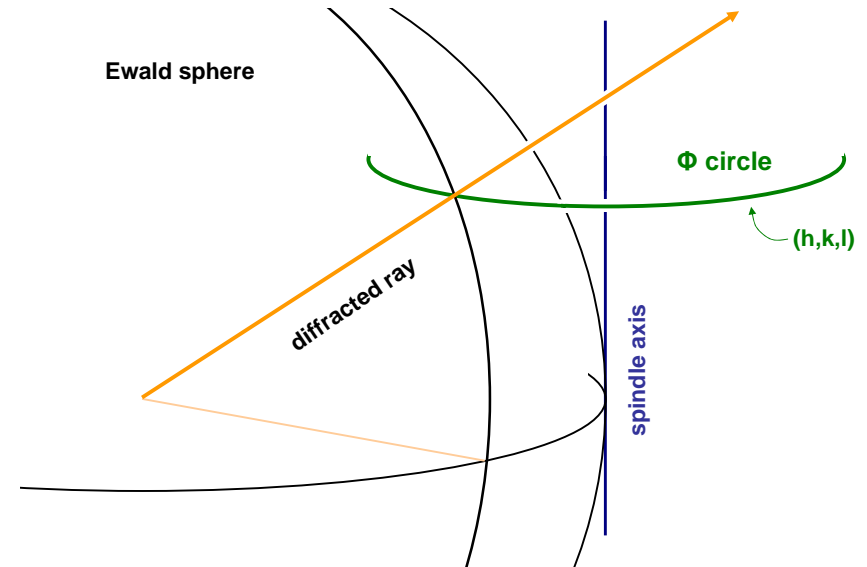
C. G. Darwin (1914)

Darwin's Formula

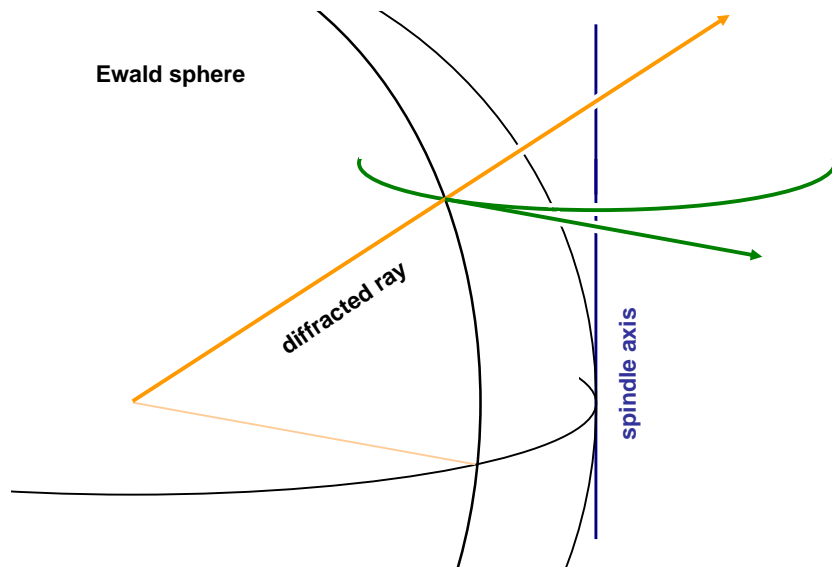
$$I(hkl) = I_{\text{beam}} r_e^2 \frac{V_{\text{xtal}}}{V_{\text{cell}}} \frac{\lambda^3 L}{\omega V_{\text{cell}}} P A |F(hkl)|^2$$

$I(hkl)$	- photons/spot (fully-recorded)	ω	- rotation speed (radians/s)
I_{beam}	- incident (photons/s/m ²)	L	- Lorentz factor (speed/speed)
r_e	- classical electron radius (2.818x10 ⁻¹⁵ m)	P	- polarization factor $(1+\cos^2(2\theta) - P_{\text{fac}} \cdot \cos(2\Phi) \sin^2(2\theta))/2$
V_{xtal}	- volume of crystal (in m ³)	A	- absorption factor $\exp(-\mu_{\text{xtal}} \cdot l_{\text{path}})$
V_{cell}	- volume of unit cell (in m ³)	$F(hkl)$	- structure amplitude (electrons)
λ	- x-ray wavelength (in meters!)		C. G. Darwin (1914)

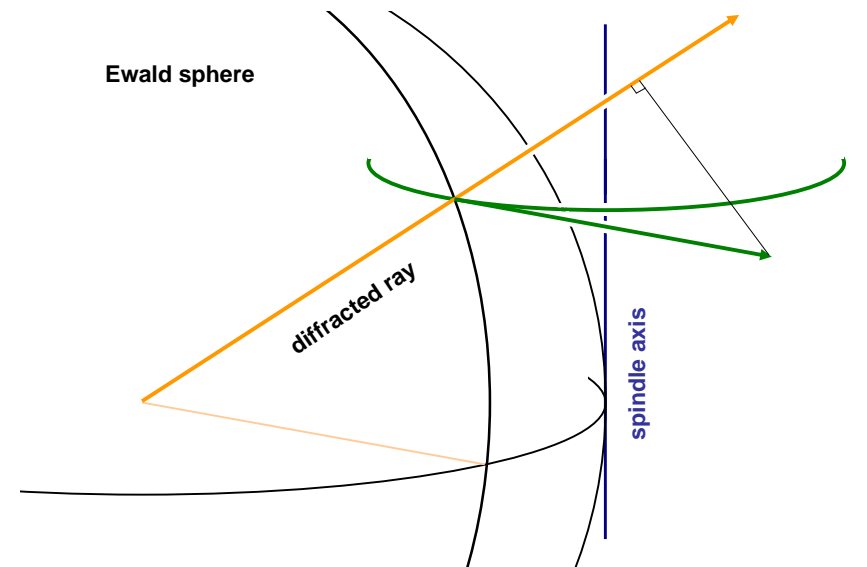
Lorentz Factor



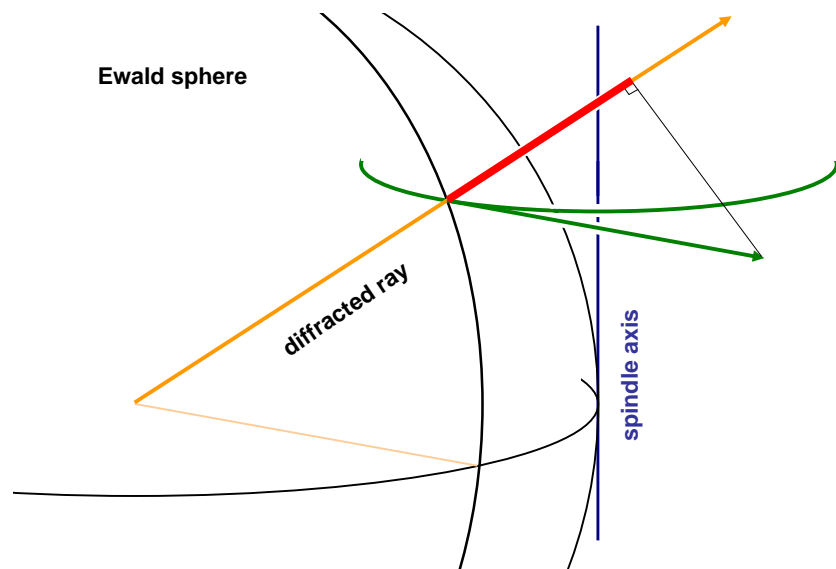
Lorentz Factor



Lorentz Factor



Lorentz Factor



Background scattering

R. W. James (1947)

$$I_{bg} = I_{beam} t r_e^2 \frac{N_A \rho V}{Mr} P A |f(s)|^2$$

I_{bg}	- scattered photons/steradian	V	- volume of material (in m ³)
I_{beam}	- incident (photons/s/m ²)	P	- polarization factor
t	- exposure time (s)	$(1 + \cos^2(2\theta) - P_{fac} \cos(2\Phi) \sin^2(2\theta))/2$	
r_e	- classical electron radius (2.818×10^{-15} m)	A	- absorption factor
N_A	- Avogadro number (6.02×10^{23})	$\exp(-\mu_{xtal} \cdot l_{path})$	
ρ	- density of material (g/m ³)	$f(s)$	- molecular structure amplitude (electrons)
Mr	- molecular weight (g/mol)	s	- scattering length ($\sin(\theta)/\lambda$)

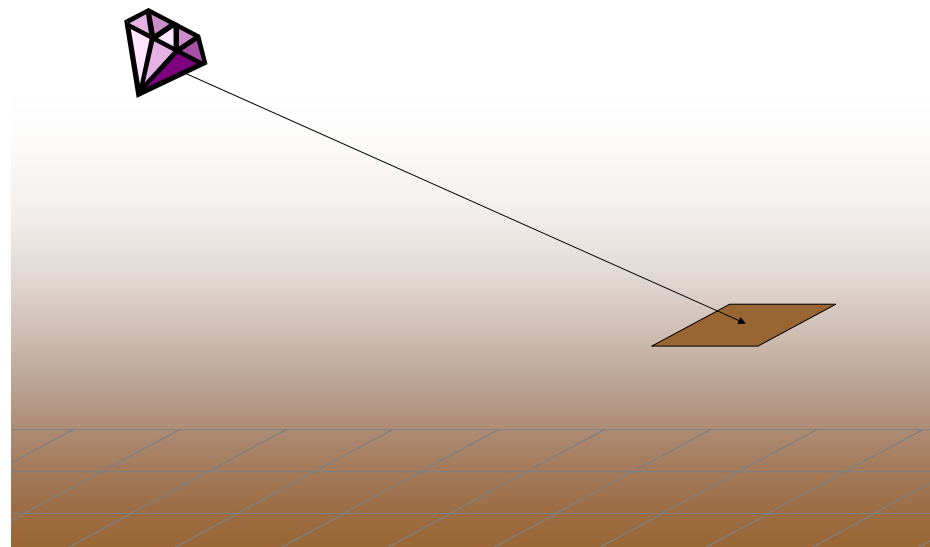
Background scattering

R. W. James (1947)

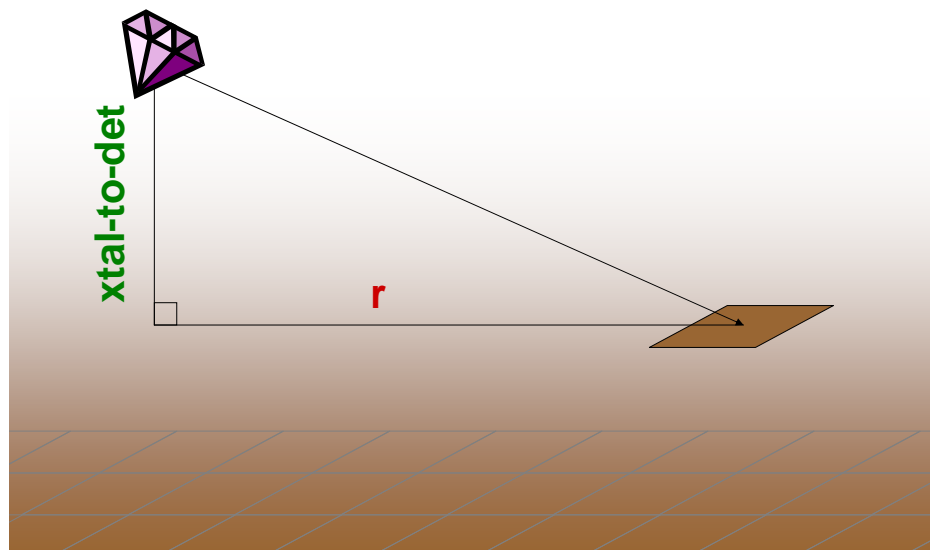
$$I_{bg} = I_{beam} t r_e^2 \frac{N_A \rho V}{Mr} P A |f(s)|^2$$

I_{bg}	- scattered photons/steradian	V	- volume of material (in m ³)
I_{beam}	- incident (photons/s/m ²)	P	- polarization factor
t	- exposure time (s)	$(1 + \cos^2(2\theta) - P_{fac} \cos(2\Phi) \sin^2(2\theta))/2$	
r_e	- classical electron radius (2.818×10^{-15} m)	A	- absorption factor
N_A	- Avogadro number (6.02×10^{23})	$\exp(-\mu_{xtal} \cdot l_{path})$	
ρ	- density of material (g/m ³)	$f(s)$	- molecular structure amplitude (electrons)
Mr	- molecular weight (g/mol)	s	- scattering length ($\sin(\theta)/\lambda$)

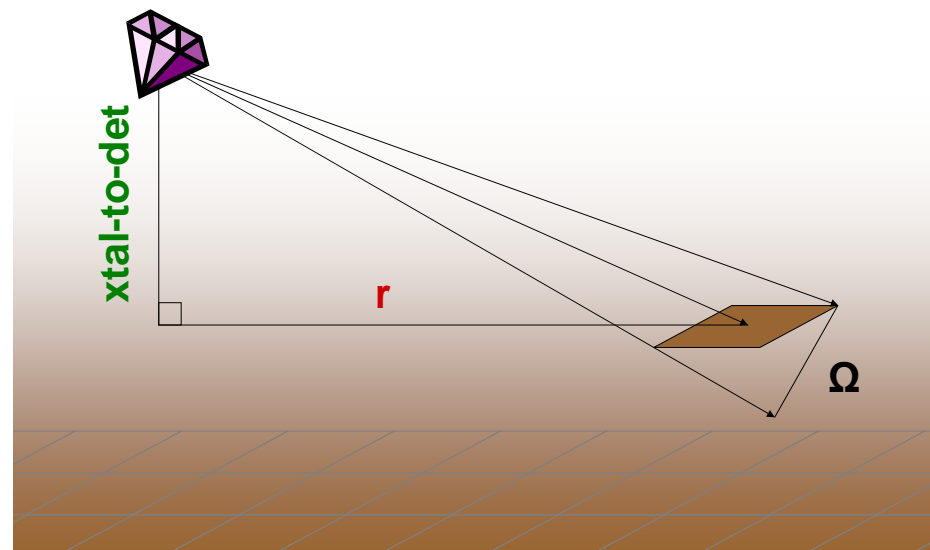
Solid angle of a pixel



Solid angle of a pixel

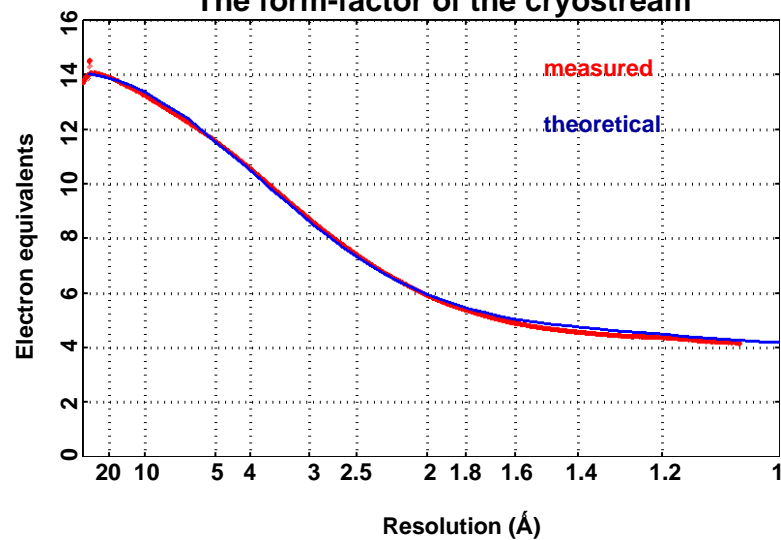


Solid angle of a pixel



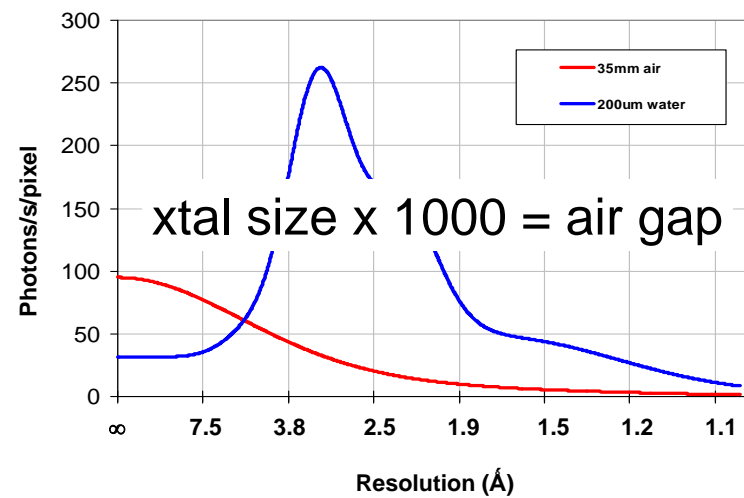
Background scattering

The form-factor of the cryostream



Background scattering

Se edge with detector at 100 mm



Diffuse scattering

R. W. James (1947)

$$I_{ds} = I_{beam} t r_e^2 \frac{V_{xtal}}{V_{ASU}} P A \sum_a |f_a(s)|^2 (1 - \exp(-2B_a \cdot s^2))$$

I_{ds}	- scattered photons/steradian	P	- polarization factor
I_{beam}	- incident (photons/s/m ²)	A	- absorption factor
t	- exposure time (s)	a	- particular atom in the ASU
r_e	- classical electron radius (2.818x10 ⁻¹⁵ m)	$f_a(s)$	- atomic structure amplitude (electrons)
V_{xtal}	- volume of crystal (in m ³)	s	- scattering length (sin(θ)/λ)
V_{ASU}	- asymmetric unit (in m ³)	B_a	- atomic B factor

Diffuse scattering

R. W. James (1947)

$$I_{ds} = I_{beam} t r_e^2 \frac{V_{xtal}}{V_{ASU}} P A \sum_a |f_a(s)|^2 (1 - \exp(-2B_a \cdot s^2))$$

I_{ds}	- scattered photons/steradian	P	- polarization factor
I_{beam}	- incident (photons/s/m ²)	A	- absorption factor
t	- exposure time (s)	a	- particular atom in the ASU
r_e	- classical electron radius (2.818x10 ⁻¹⁵ m)	$f_a(s)$	- atomic structure amplitude (electrons)
V_{xtal}	- volume of crystal (in m ³)	s	- scattering length (sin(θ)/λ)
V_{ASU}	- asymmetric unit (in m ³)	B_a	- atomic B factor

mosaicity vs Wilson B

INTENSITY OF REFLECTION OF X-RAYS BY CRYSTALS 35

(b) *The amplitude reflected by a plane sheet of atoms*: We shall first consider the amplitude of the wave reflected by an infinite plane sheet of atoms, each of which scatters the incident X-rays.

Suppose A, fig. 15, is the source of the radiation, and let the amplitude of the reflected wave be required at B. Let the plane APB be normal to the plane of atoms, and let AP, PB make equal angles θ with this plane. Then P is such that the distance APB is the shortest distance from A to B via the plane. Let M be a point of the plane such that the distance

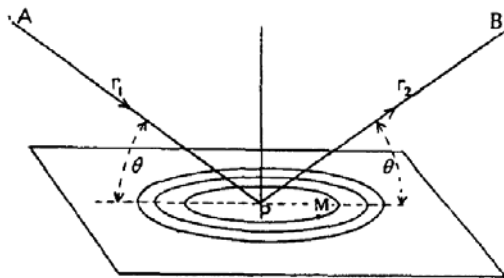


FIG. 15

James R. W. (1962) *Optical Principles of the Diffraction of X rays*. Ox Bow press.

Compton scattering

R. W. James (1947)

$$I_{compton} = I_{beam} t r_e^2 \frac{V_{xtal}}{V_{ASU}} \frac{P A}{1 + \kappa (1 - \cos(\theta))^2} \sum_a Z_a |f_a(s)|$$

$I_{compton}$	- scattered photons/steradian	A	- absorption factor
I_{beam}	- incident (photons/s/m ²)	a	- particular atom in the ASU
t	- exposure time (s)	θ	- Bragg angle
r_e	- classical electron radius (2.818x10 ⁻¹⁵ m)	κ	- photon energy (keV) / 511 keV
V_{xtal}	- volume of crystal (in m ³)	Z_a	- atomic number
V_{ASU}	- asymmetric unit (in m ³)	$f_a(s)$	- atomic structure amplitude (electrons)
P	- polarization factor	s	- scattering length (sin(θ)/λ)

Compton scattering

R. W. James (1947)

$$I_{\text{compton}} = I_{\text{beam}} t r_e^2 \frac{V_{\text{xtal}}}{V_{\text{ASU}}} \frac{P A}{1 + \kappa (1 - \cos(\theta))^2} \sum_a Z_a - |f_a(s)|$$

I_{compton}	- scattered photons/steradian	A	- absorption factor
I_{beam}	- incident (photons/s/m ²)	a	- particular atom in the ASU
t	- exposure time (s)	θ	- Bragg angle
r_e	- classical electron radius (2.818x10 ⁻¹⁵ m)	κ	- photon energy (keV) / 511 keV
V_{xtal}	- volume of crystal (in m ³)	Z_a	- atomic number
V_{ASU}	- asymmetric unit (in m ³)	$f_a(s)$	- atomic structure amplitude (electrons)
P	- polarization factor	s	- scattering length (sin(θ)/λ)

Radiation Damage Model

$$\text{dose} = I_{\text{beam}} t E (1 - \exp(-\mu_{\text{xtal}} T_{\text{xtal}})) / \rho$$

dose	- absorbed dose (Gy)
I_{beam}	- incident (photons/s/m ²)
E	- photon energy (J/photon)
t	- exposure time (s)
μ_{xtal}	- absorption coefficient (in 1/m)
T_{xtal}	- thickness of crystal (in m)
ρ	- crystal density (kg/m ³)

Radiation Damage Model

$$\text{dose} = I_{\text{beam}} t k_{\text{dose}}$$

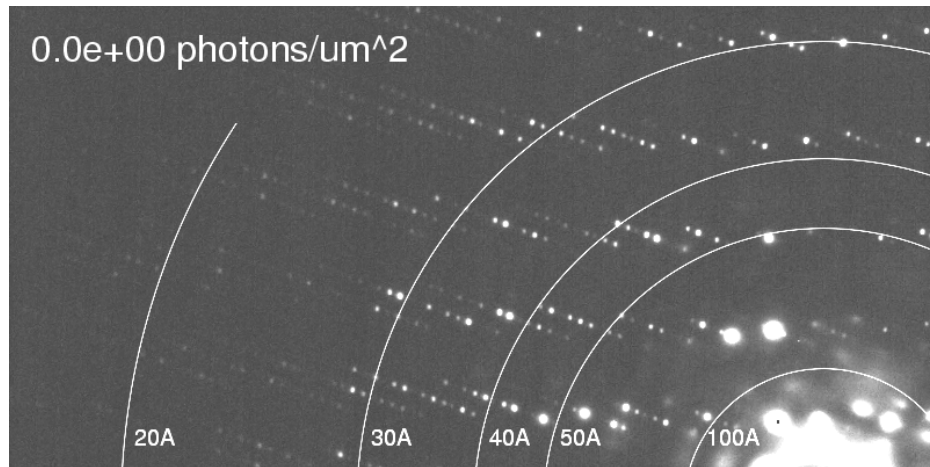
dose	- absorbed dose (MGy)
I_{beam}	- incident (~20 M photons/s/m ²)
t	- exposure time (s)
k_{dose}	- 2000 (photon/μm ² /Gy)

Radiation Damage Model

$$\text{dose} = I_{\text{beam}} t k_{\text{dose}}$$

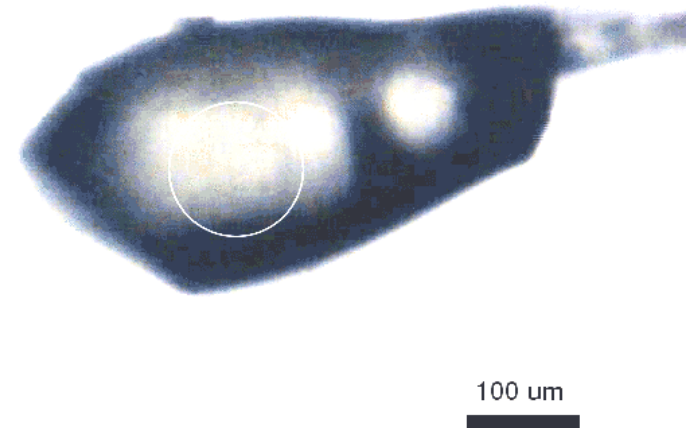
dose	- absorbed dose (MGy)
I_{beam}	- incident (~20 M photons/s/m ²)
t	- exposure time (s)
k_{dose}	- 2000 (photon/μm ² /Gy)

Radiation damage



Macroscopic damage

0.0e+00 photons/ μm^2



crystal expansion

Protein crystal in sucrose, NaWO_4 and oil



crystal expansion

Protein crystal in sucrose, NaWO_4 and oil



crystal expansion

Protein crystal in sucrose, NaWO₄ and oil



crystal expansion

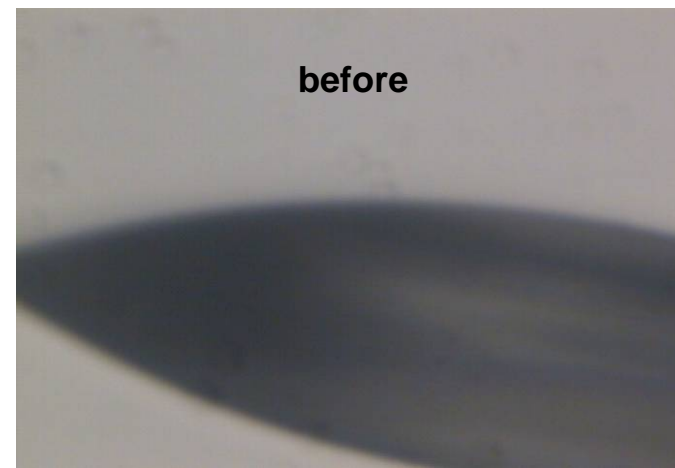
Protein crystal in sucrose, NaWO₄ and oil



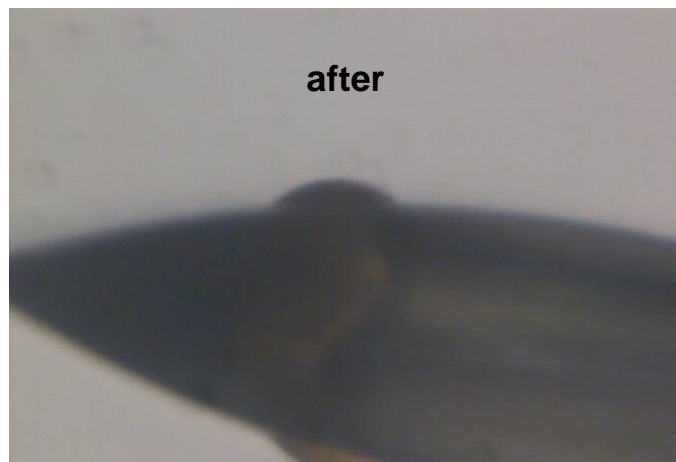
Distention of cryo with dose



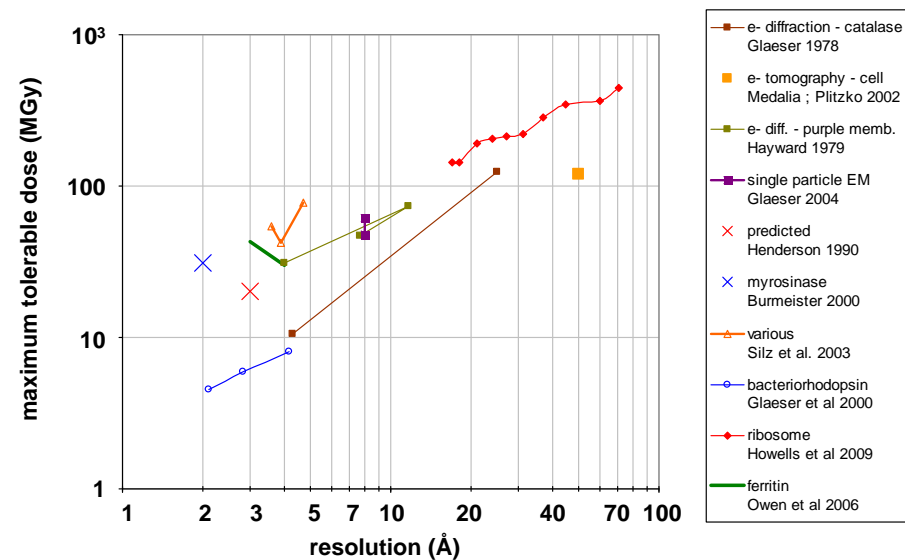
Distention of cryo with dose



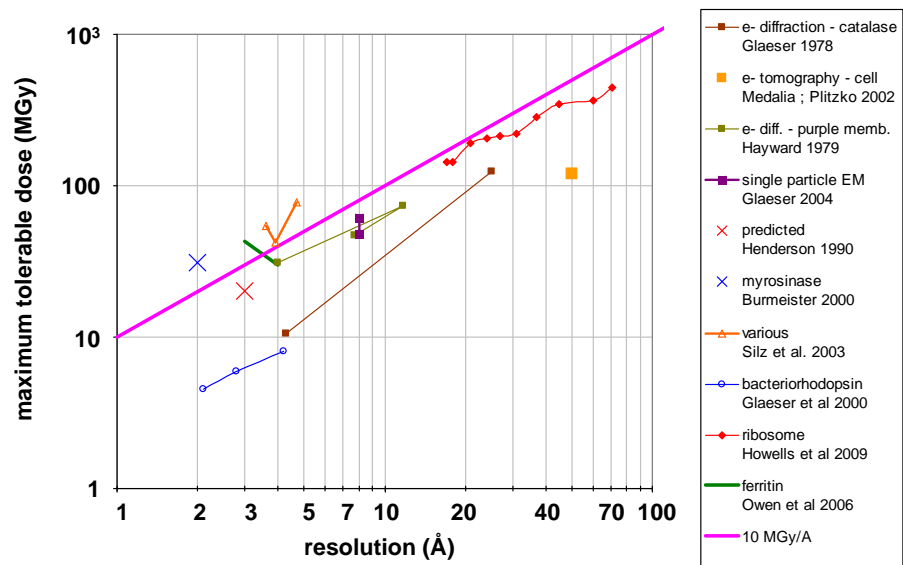
Distention of cryo with dose



Howells *et al.* (2009)
J. Electron. Spectrosc. Relat. Phenom. **170** 4-12

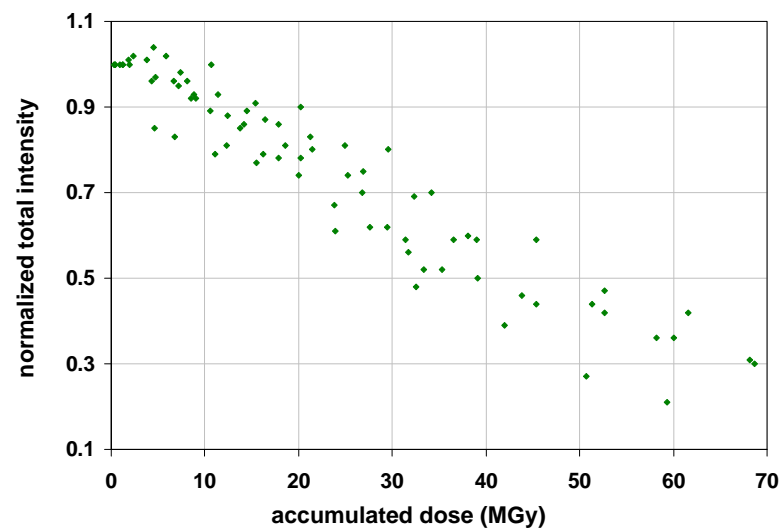


Howells *et al.* (2009)
J. Electron. Spectrosc. Relat. Phenom. **170** 4-12



Radiation Damage Model

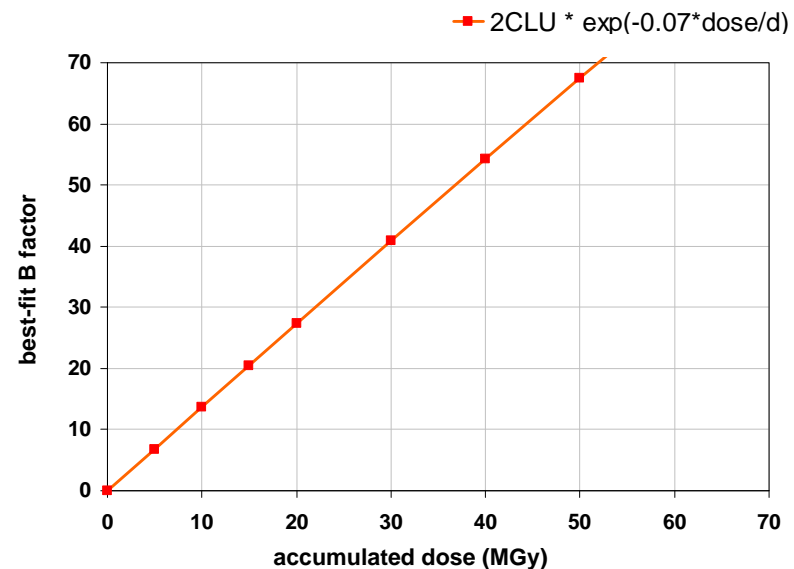
• Owen *et al.* (2006)



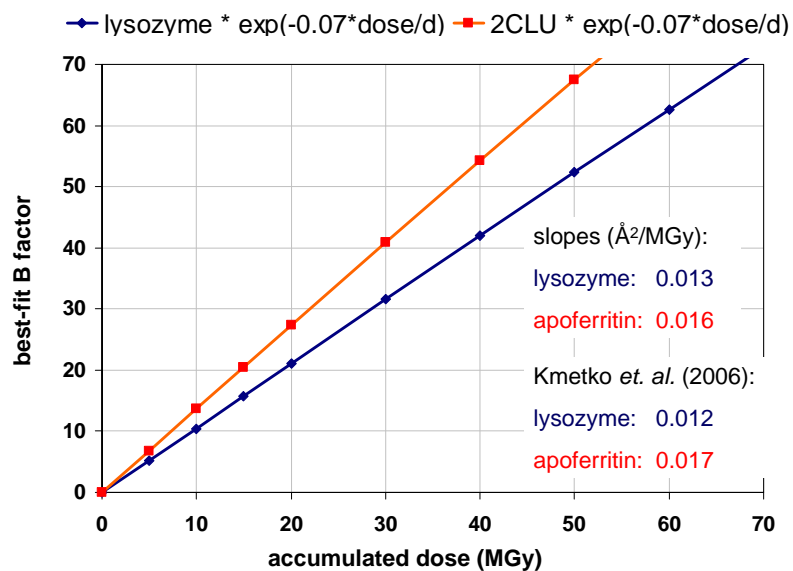
Radiation Damage Model



Radiation Damage Model



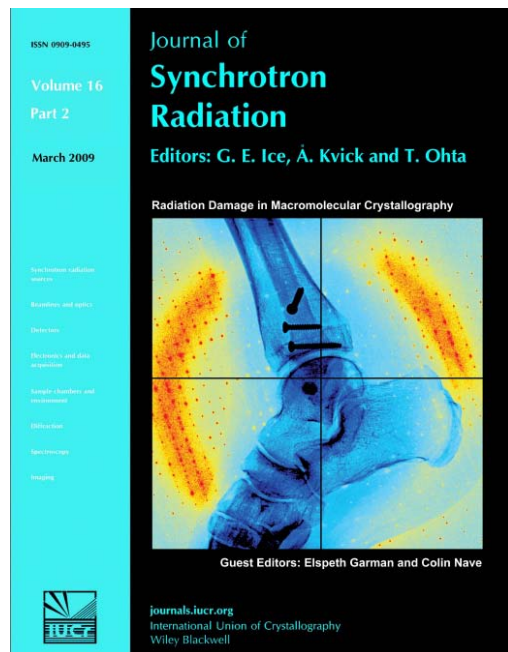
Radiation Damage Model



10 MGy/Å

what the is a MGy?

[http://bl831.als.lbl.gov/
damage_rates.pdf](http://bl831.als.lbl.gov/damage_rates.pdf)



Holton J. M. (2009) *J. Synchrotron Rad.* **16** 133-42

Radiation Damage Model

global (lattice) damage

$$|F_{\text{used}}| = |F_{\text{undam}}| \exp(-\ln(2) \frac{\text{dose}}{d \cdot H})$$

- F_{used} - structure factor used for spot
- F_{undam} - structure factor of undamaged crystal
- dose** - absorbed dose (MGy)
- H - 10 MGy/Å
- d - resolution of spot (Å)

Radiation Damage Model

global (lattice) damage

$$F = F_0 \exp(-A \cdot s)$$

- F - structure factor used for spot
- F_0 - structure factor of undamaged crystal
- A - something Debye said was zero
- s - $0.5/d$
- d - resolution of spot (Å)

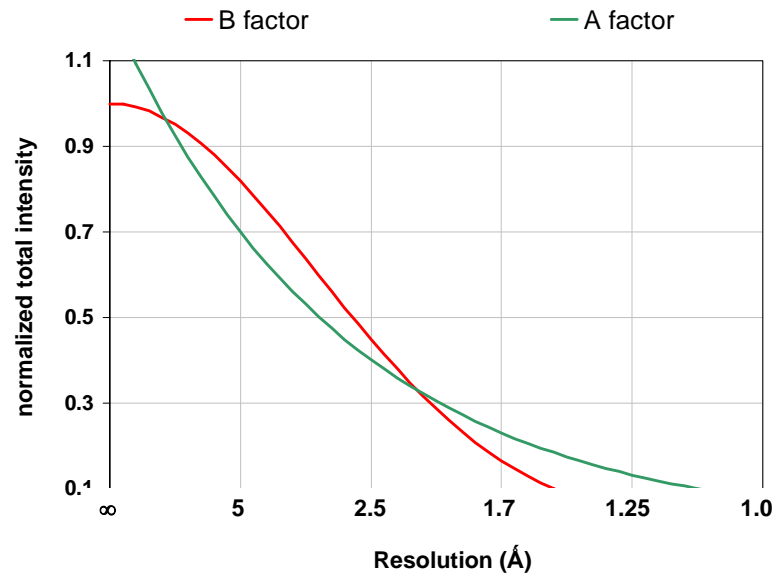
Radiation Damage Model

global (lattice) damage

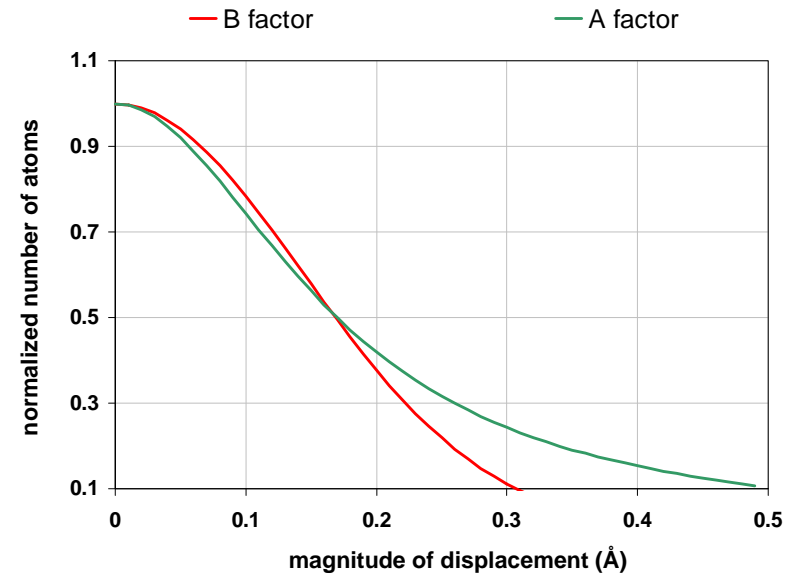
$$F = F_0 \exp(-A \cdot s - B \cdot s^2 - C \cdot s^3 - \dots)$$

- A - something Debye said was zero
- B - canonical Debye-Waller factor
- C - something else Debye said was zero
- s - $0.5/d$
- d - resolution of spot (Å)

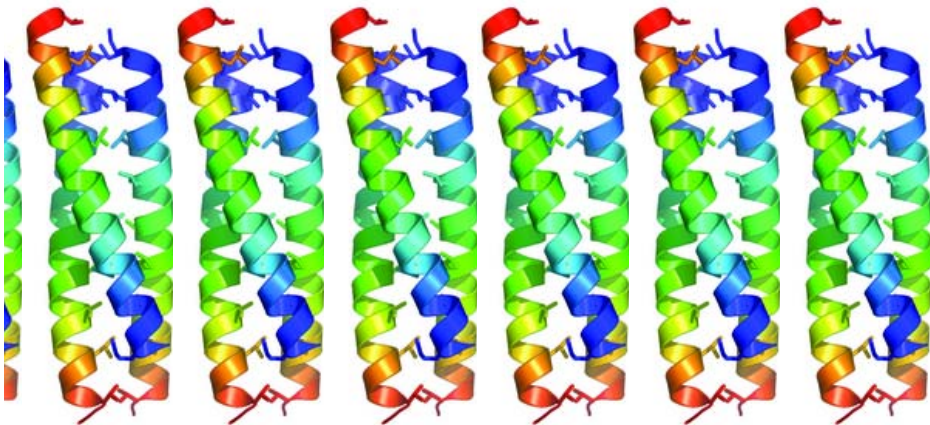
Radiation Damage Model



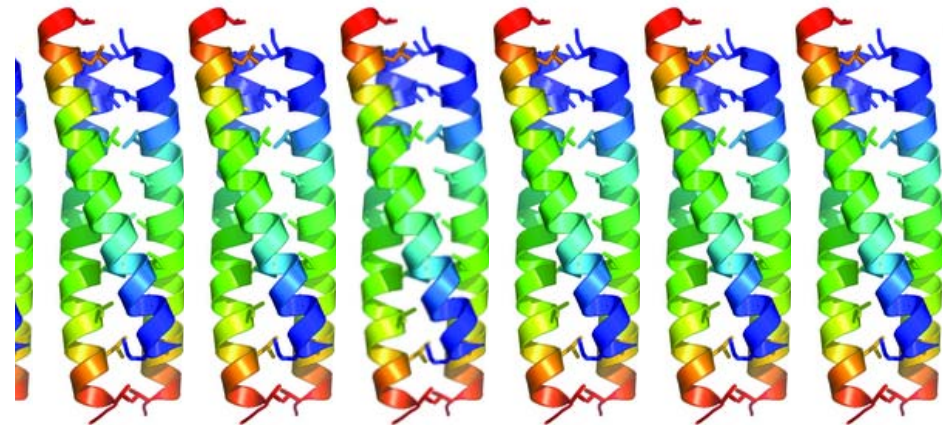
Radiation Damage Model



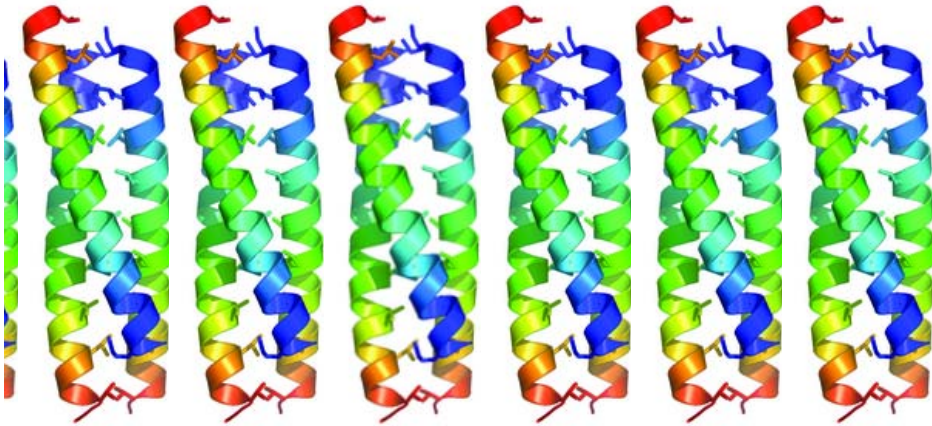
Radiation Damage Model



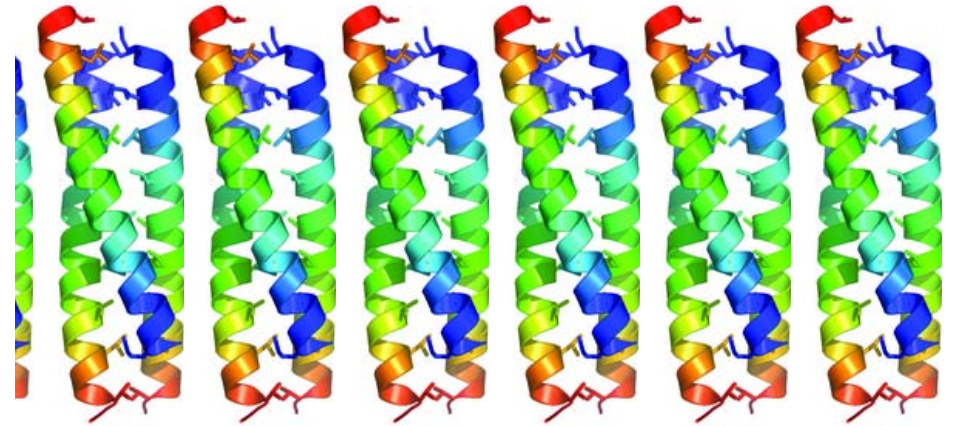
Radiation Damage Model



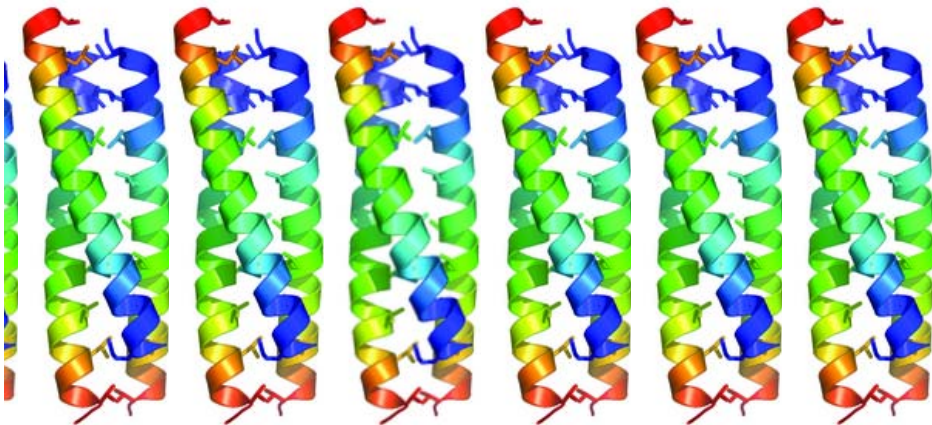
Radiation Damage Model



Radiation Damage Model

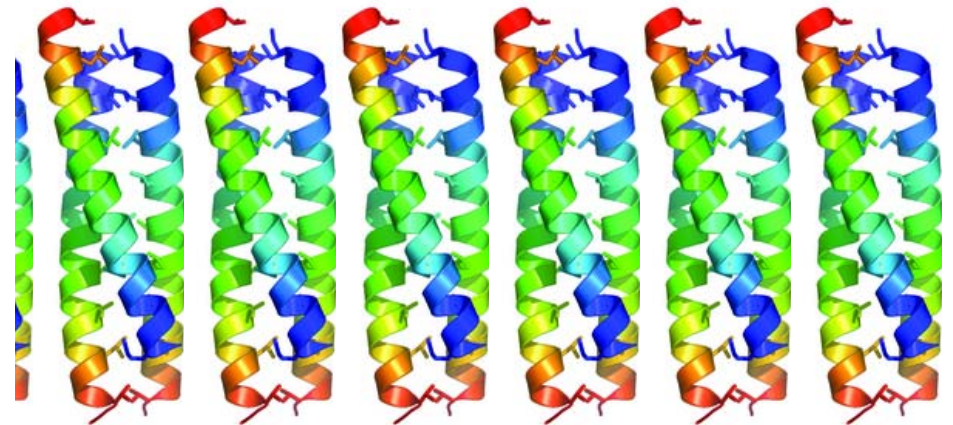


Radiation Damage Model



Kanzaki 1957

Radiation Damage Model



Kanzaki 1957

Radiation Damage Model

global (lattice) damage

$$|F_{\text{used}}| = |F_{\text{undam}}| \exp(-\ln(2) \frac{\text{dose}}{d \cdot H})$$

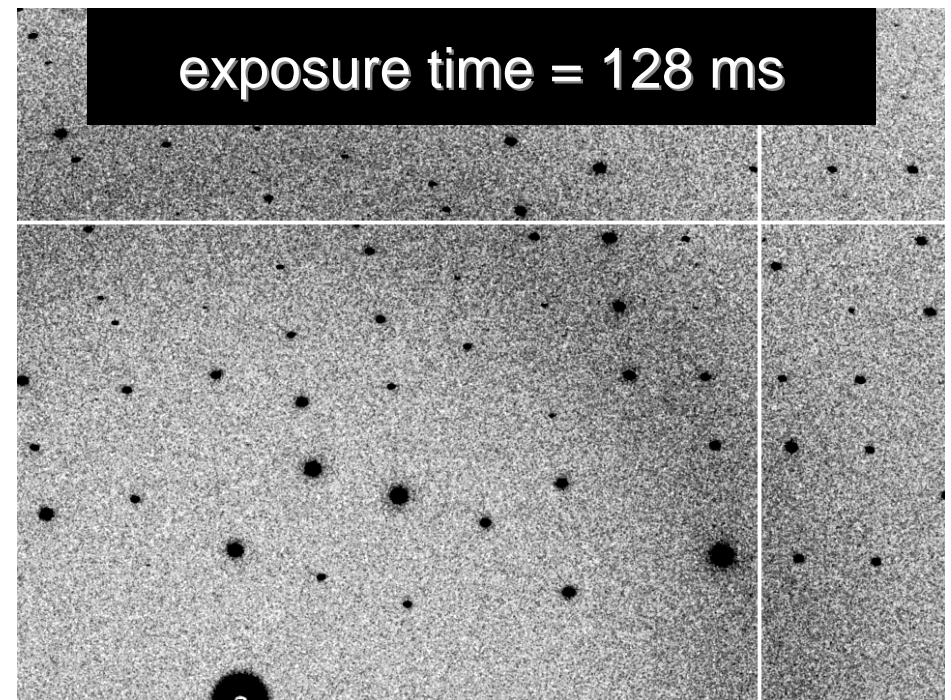
- F_{used} - structure factor used for spot
- F_{undam} - structure factor of undamaged crystal
- dose** - absorbed dose (MGy)
- H - 10 MGy/Å
- d - resolution of spot (Å)

Development Snapshot

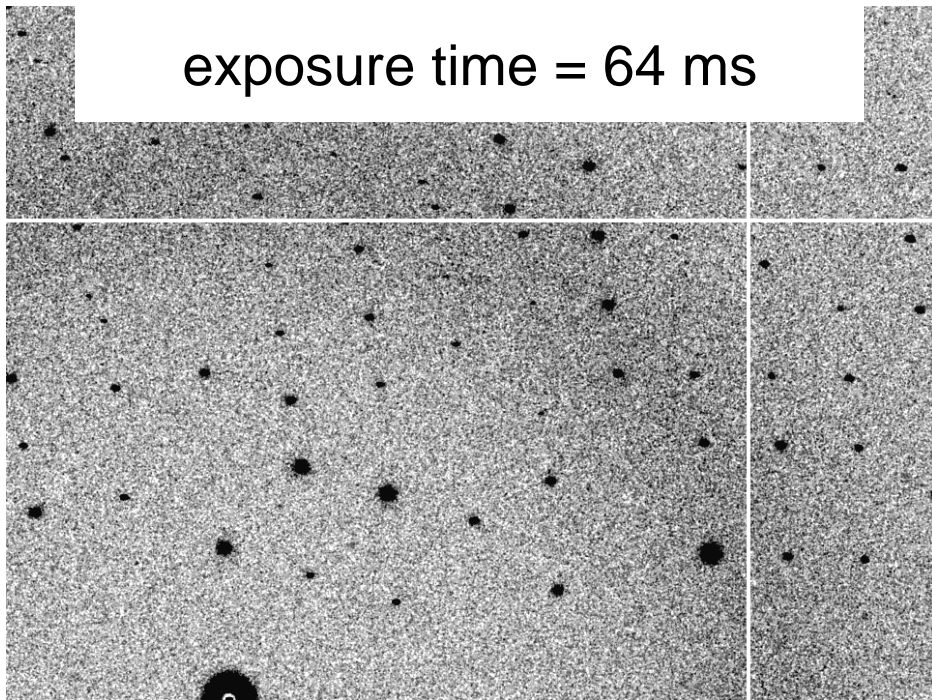
http://bl831.als.lbl.gov/~jamesh/mlfsom/development_snapshot.tar.gz

http://bl831.als.lbl.gov/~jamesh/powerpoint/ACA_2009.ppt

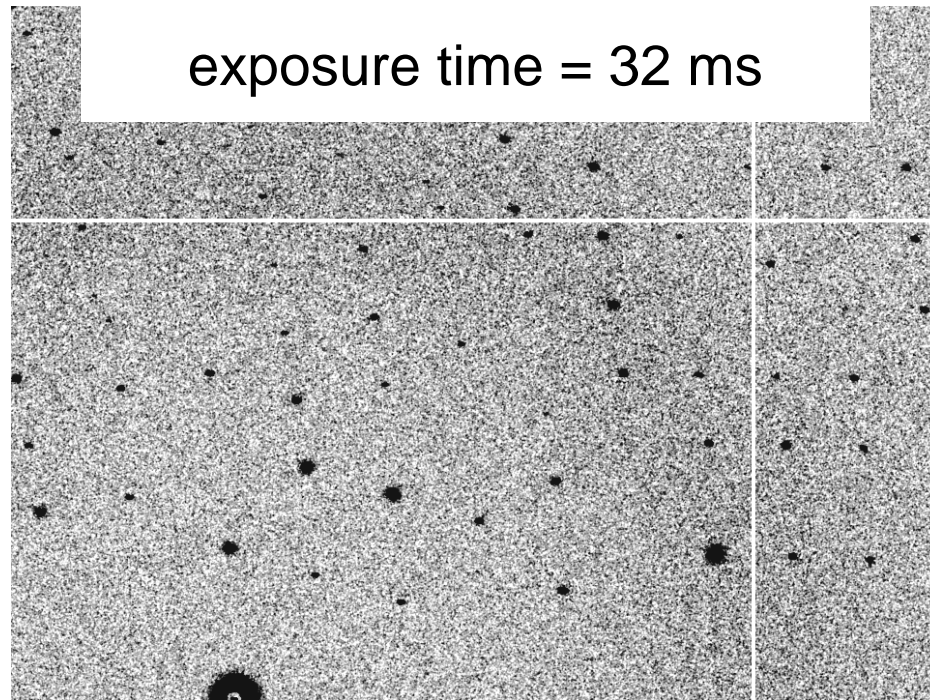
What is the
best exposure time?



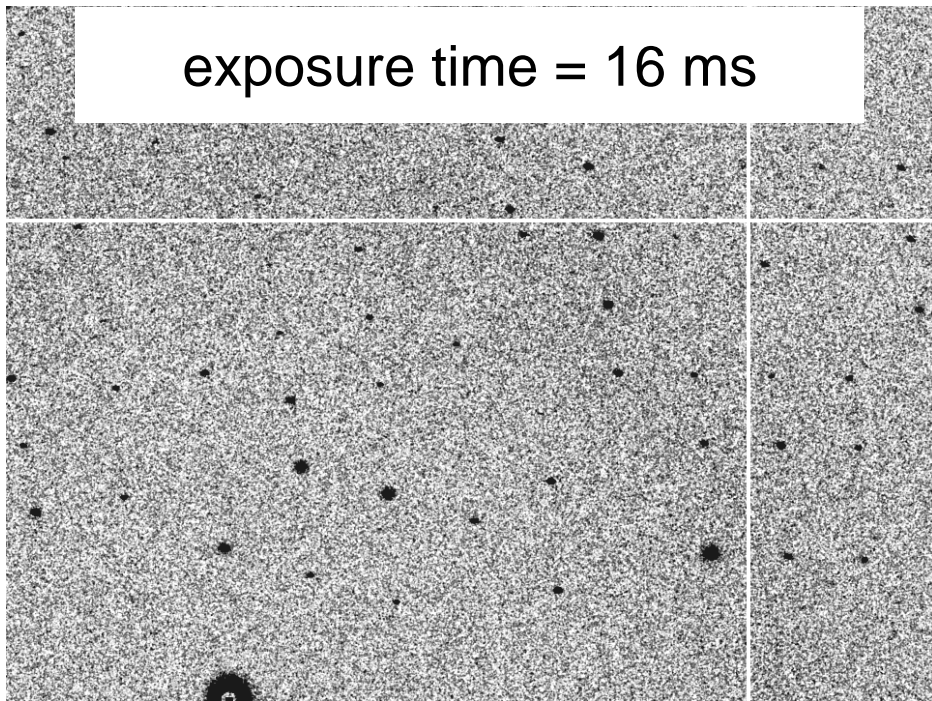
exposure time = 64 ms



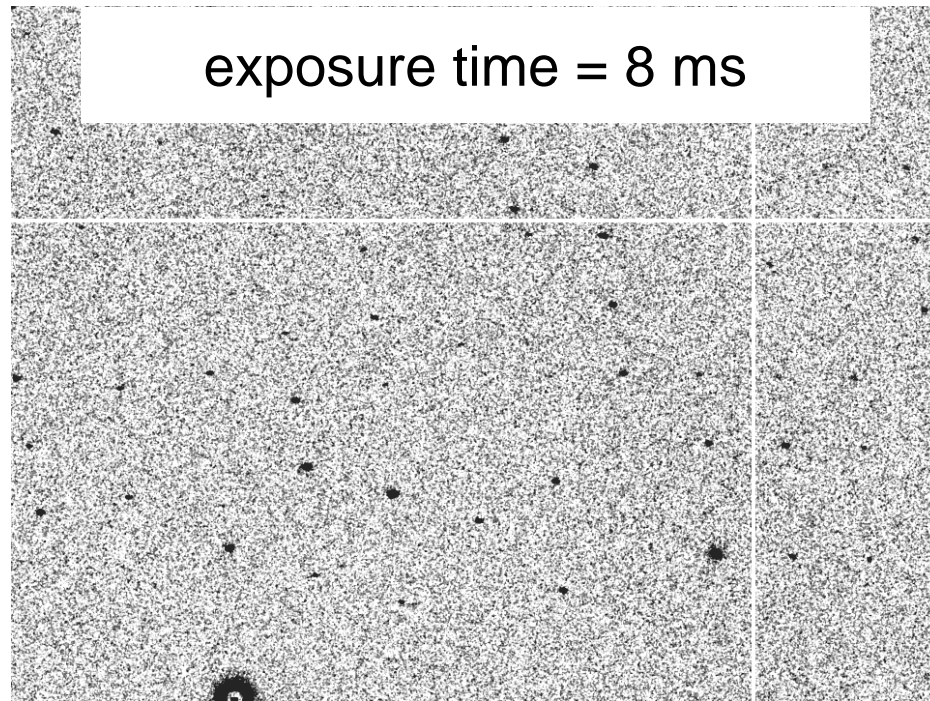
exposure time = 32 ms



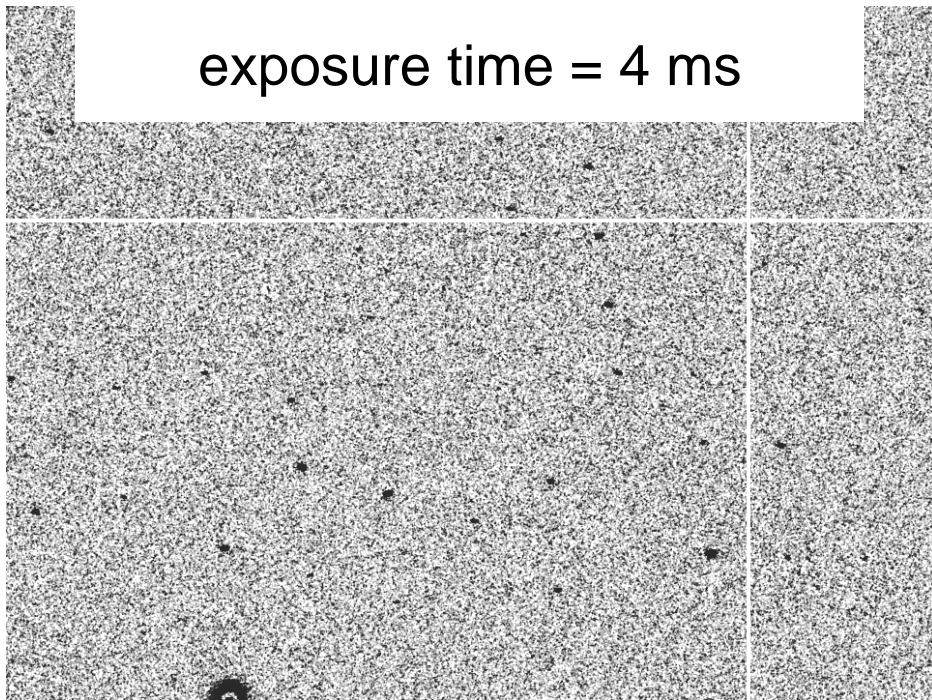
exposure time = 16 ms



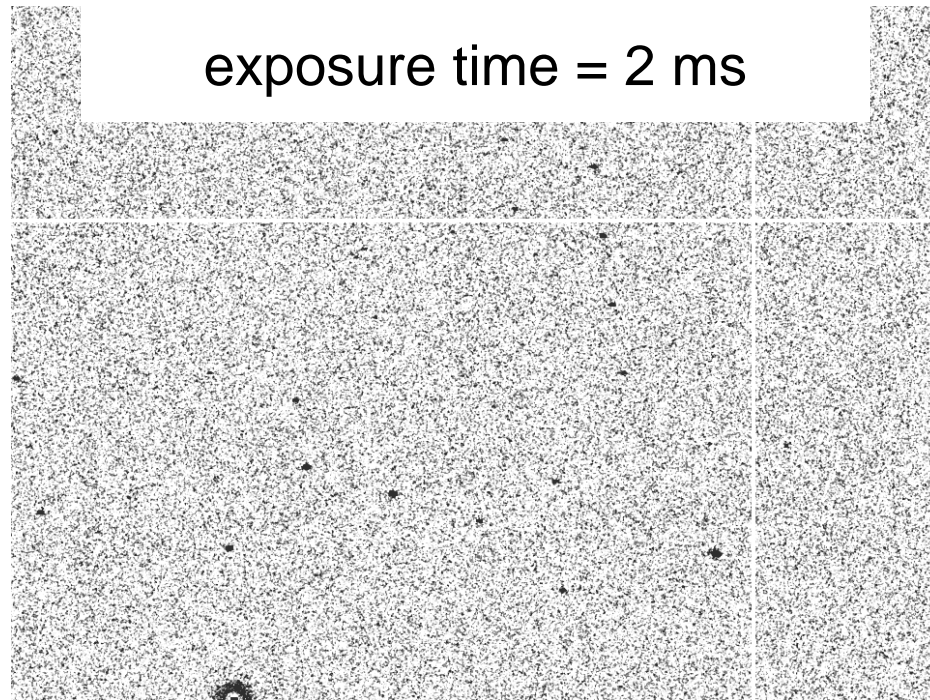
exposure time = 8 ms



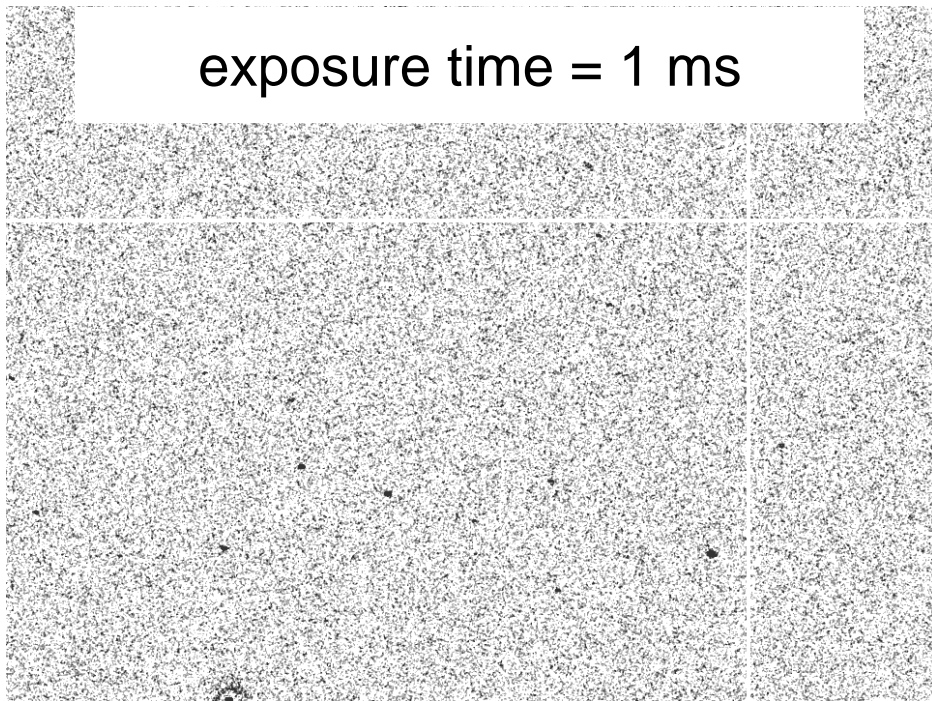
exposure time = 4 ms



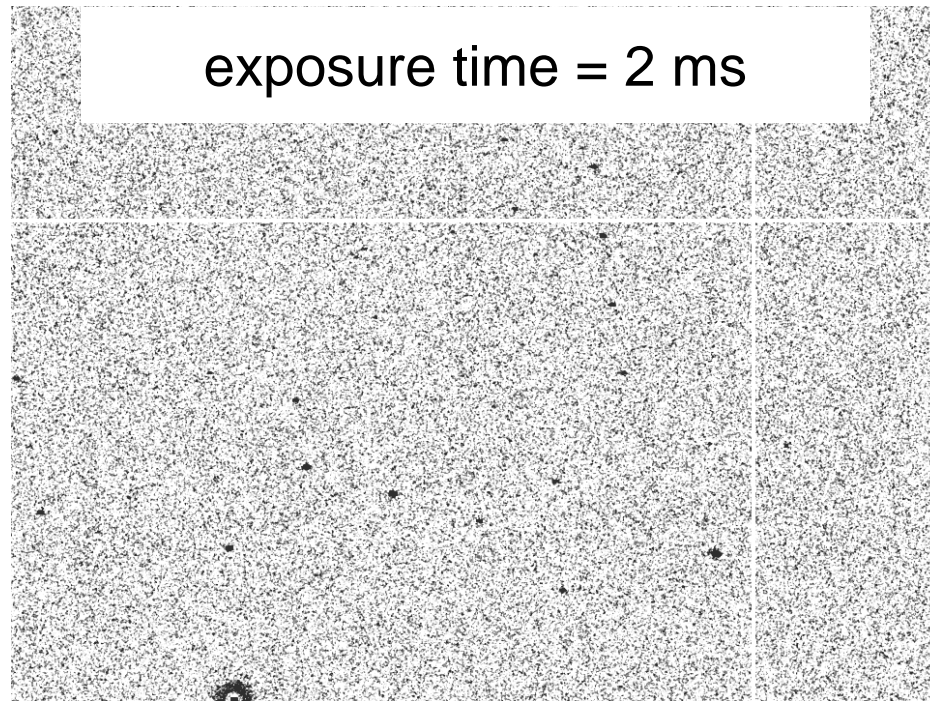
exposure time = 2 ms



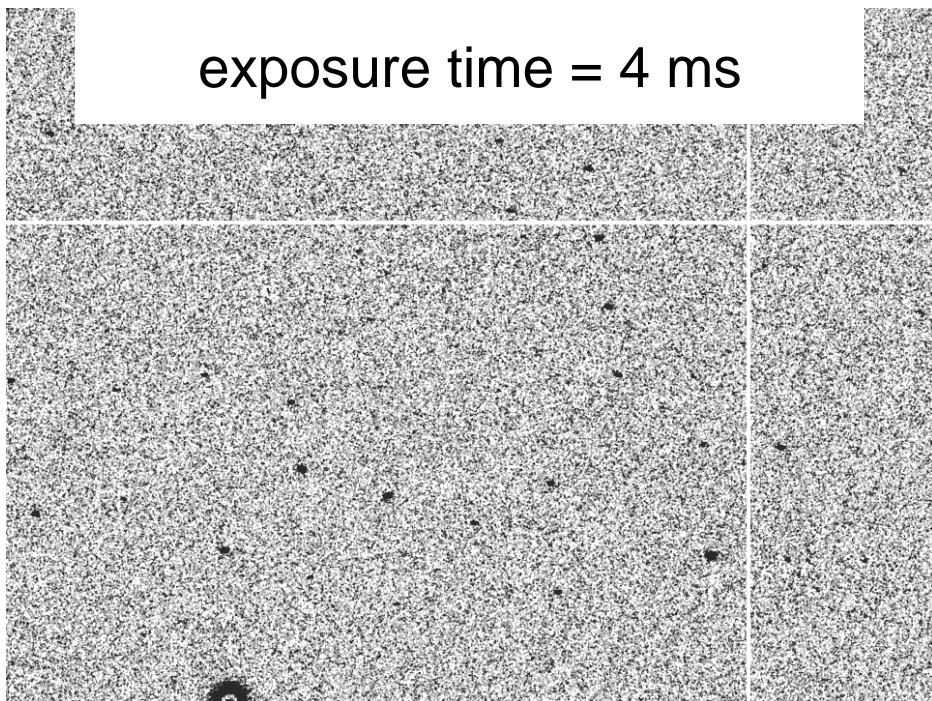
exposure time = 1 ms



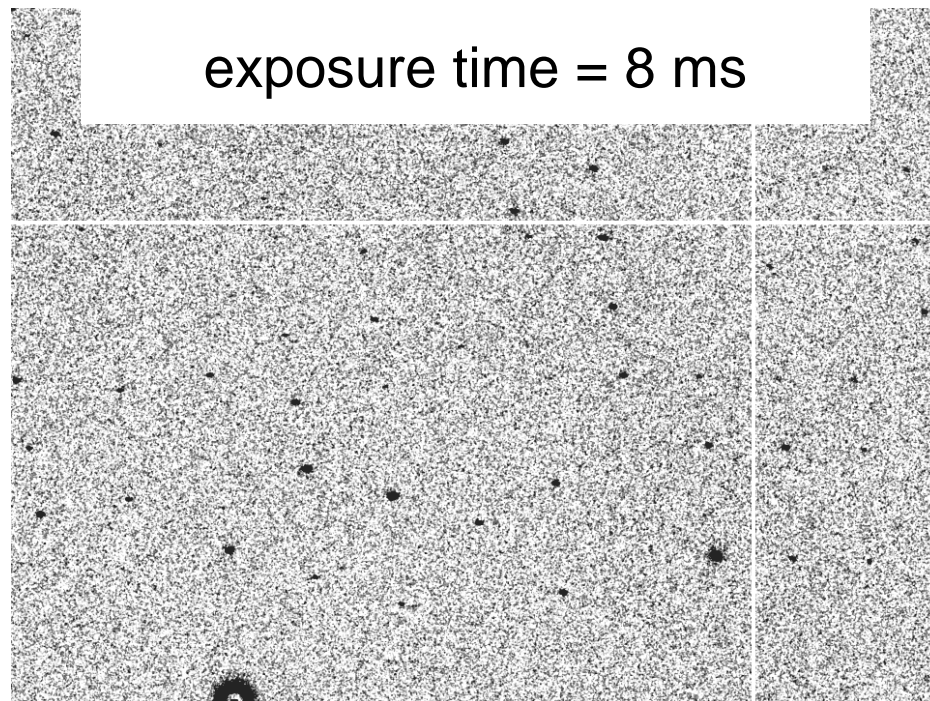
exposure time = 2 ms



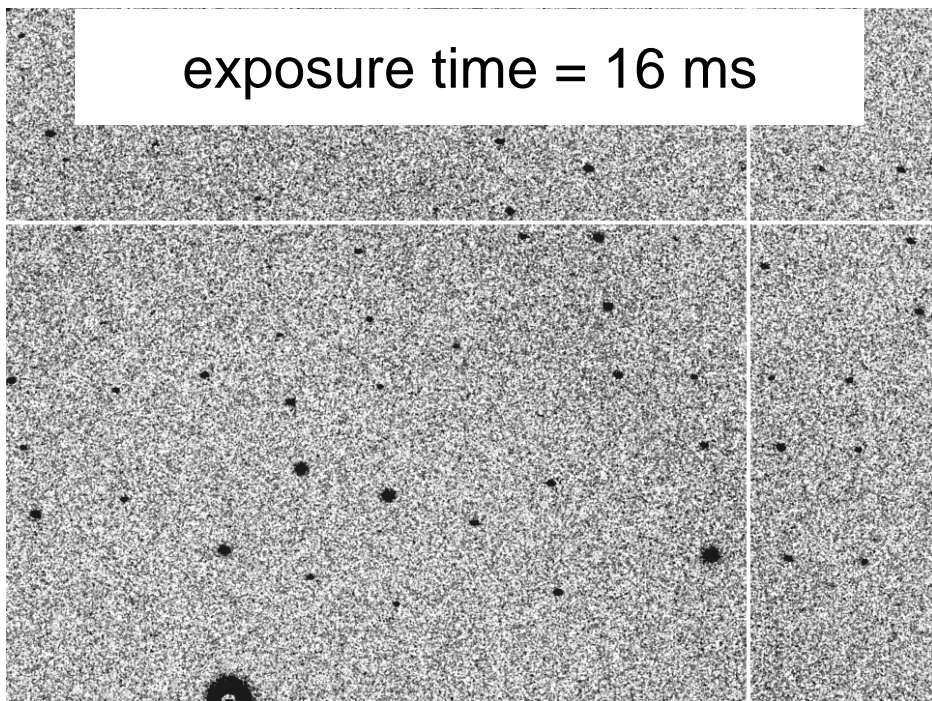
exposure time = 4 ms



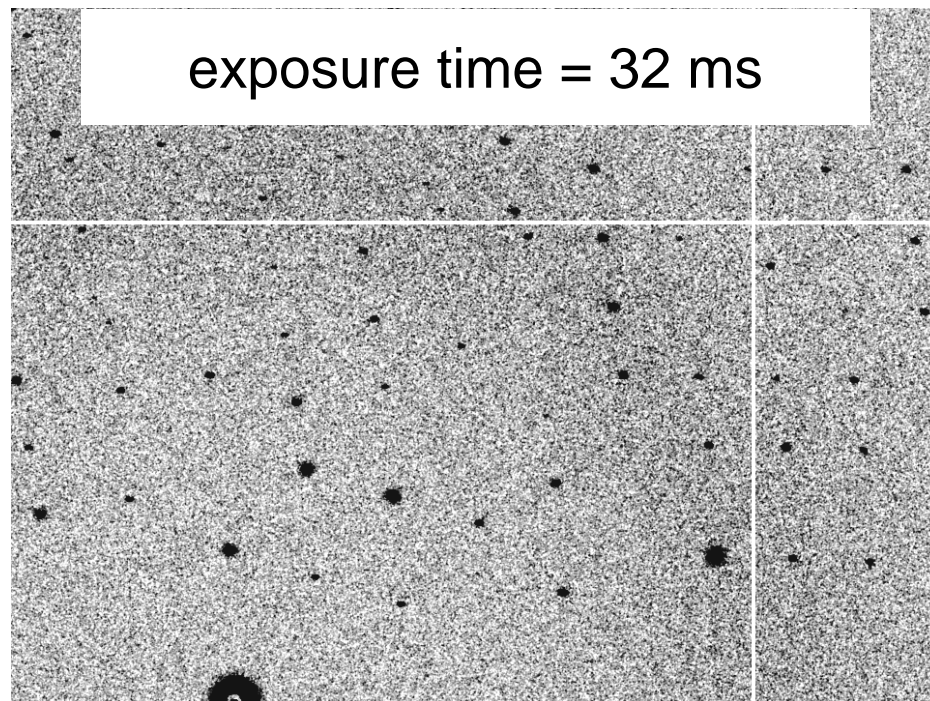
exposure time = 8 ms



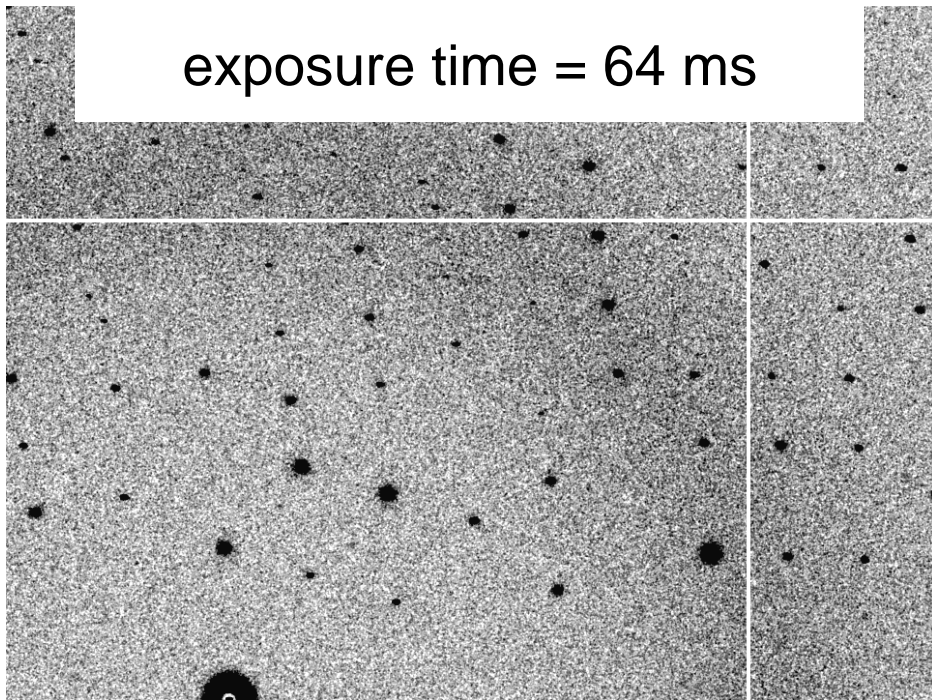
exposure time = 16 ms



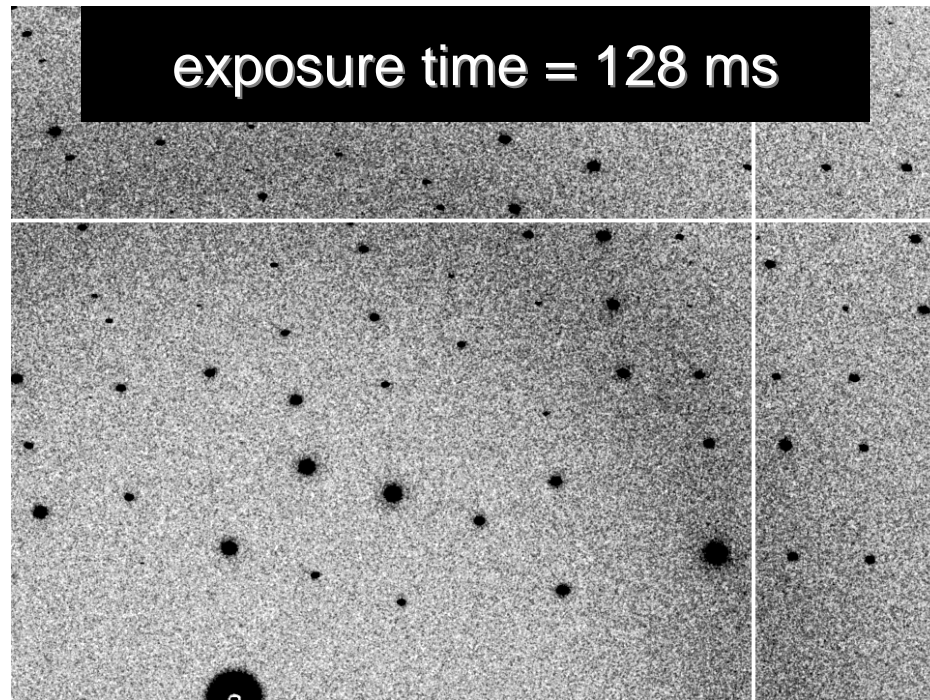
exposure time = 32 ms



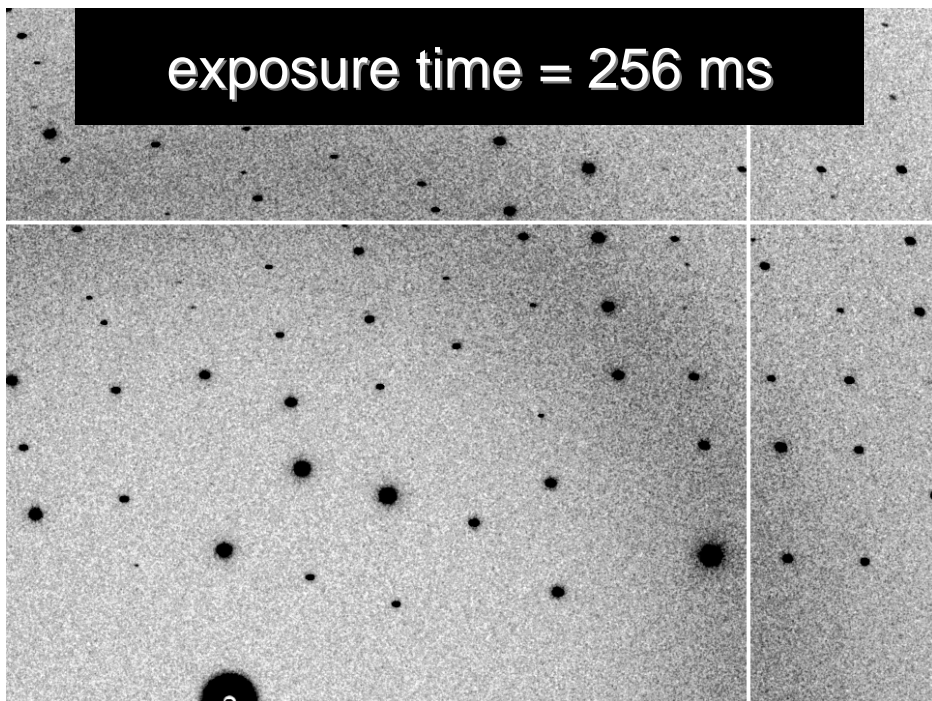
exposure time = 64 ms



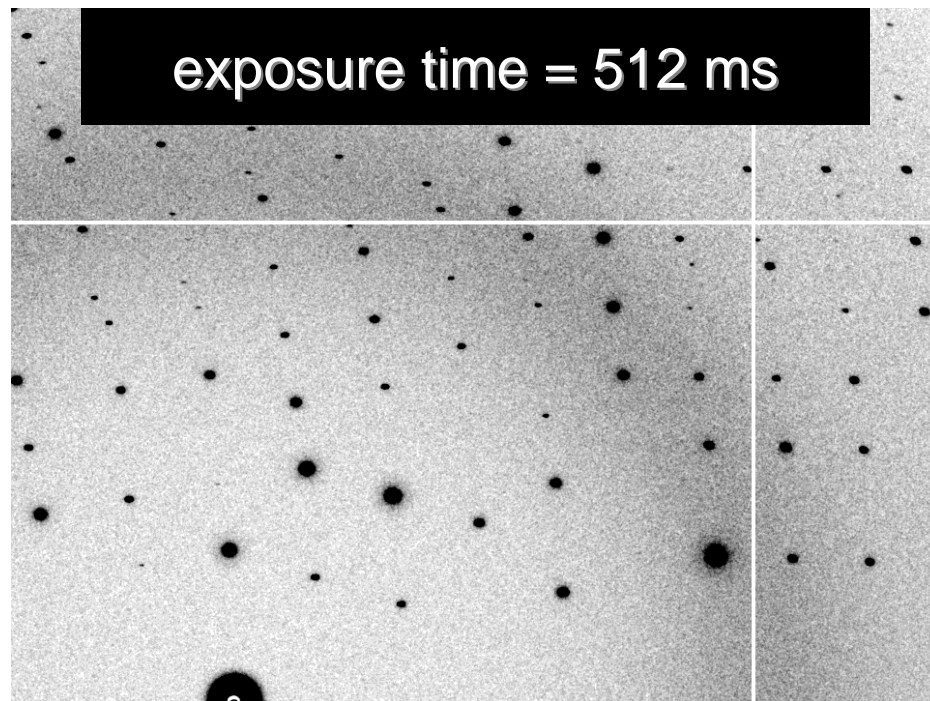
exposure time = 128 ms



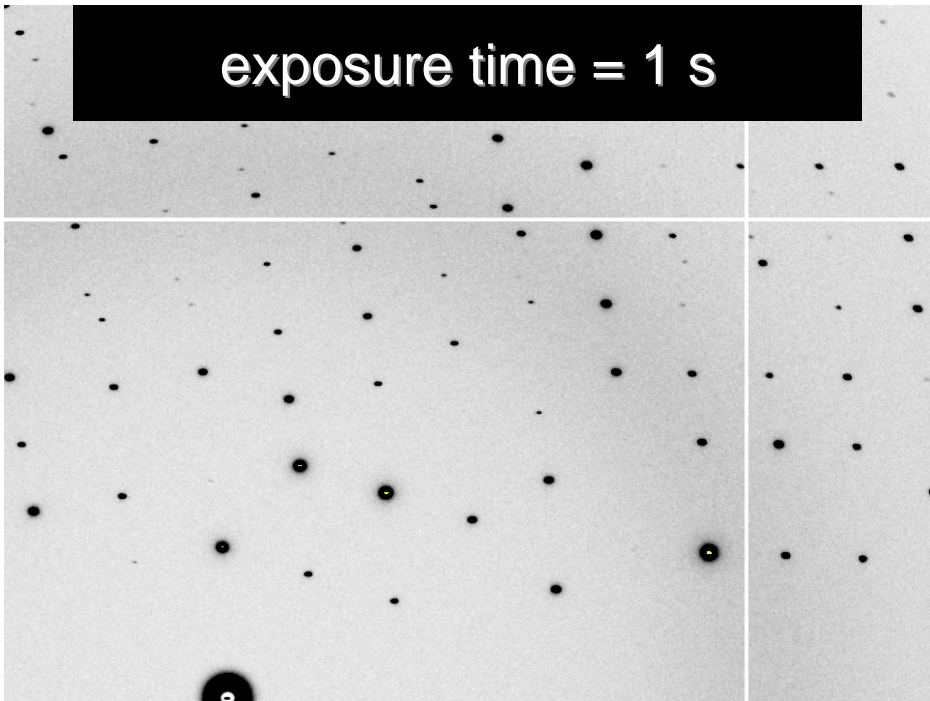
exposure time = 256 ms



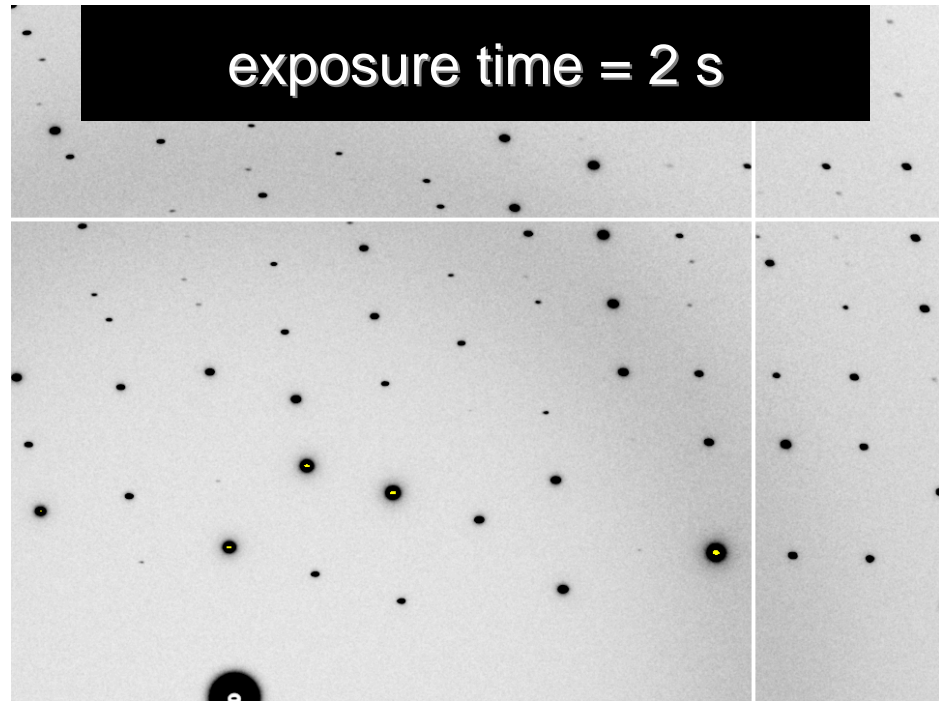
exposure time = 512 ms



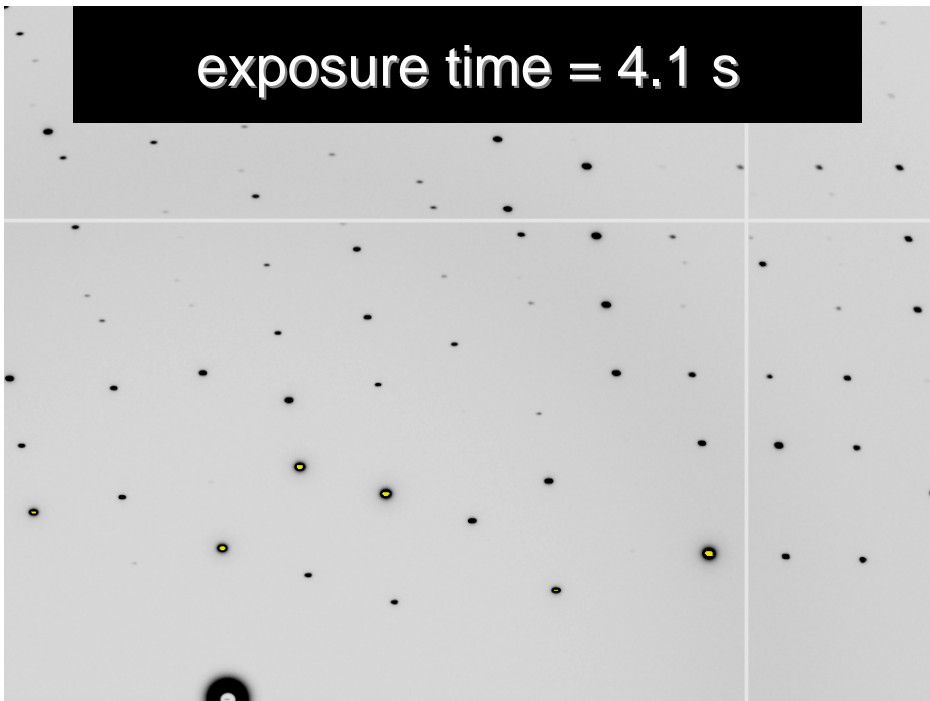
exposure time = 1 s



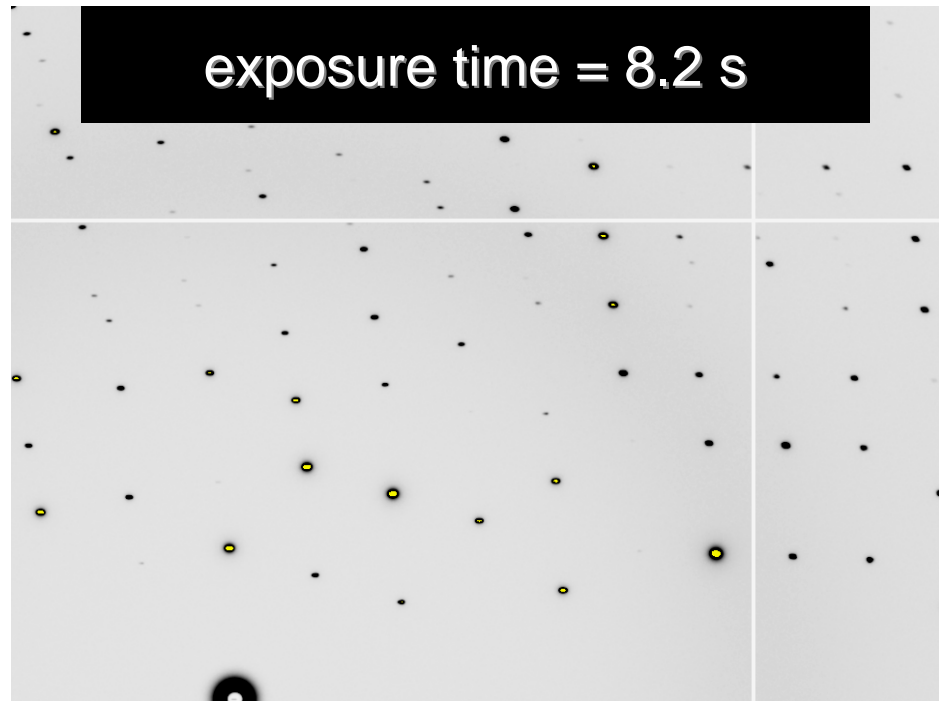
exposure time = 2 s



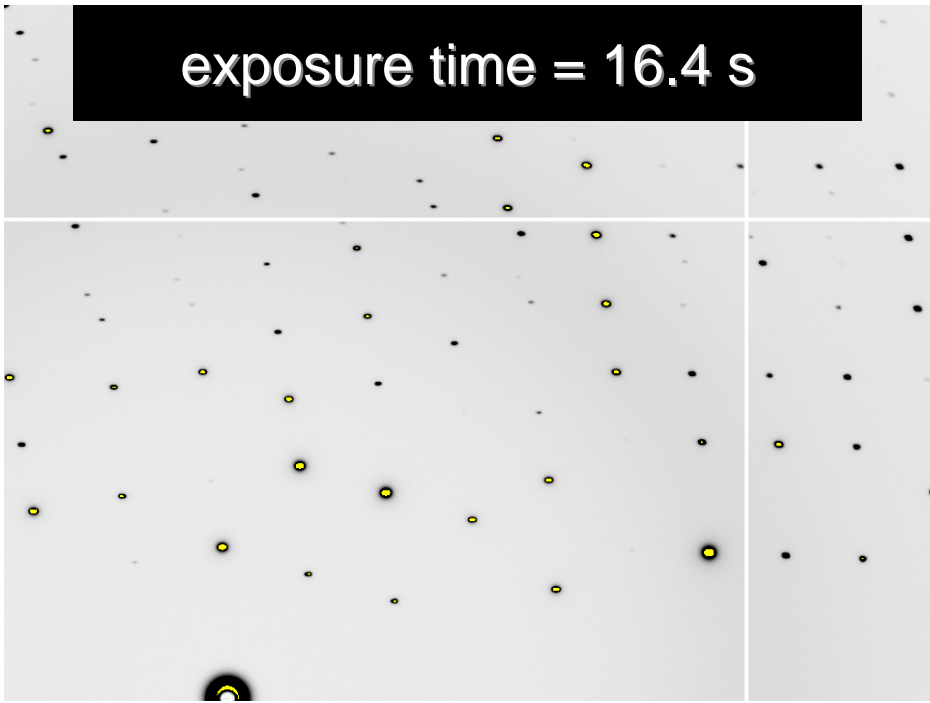
exposure time = 4.1 s



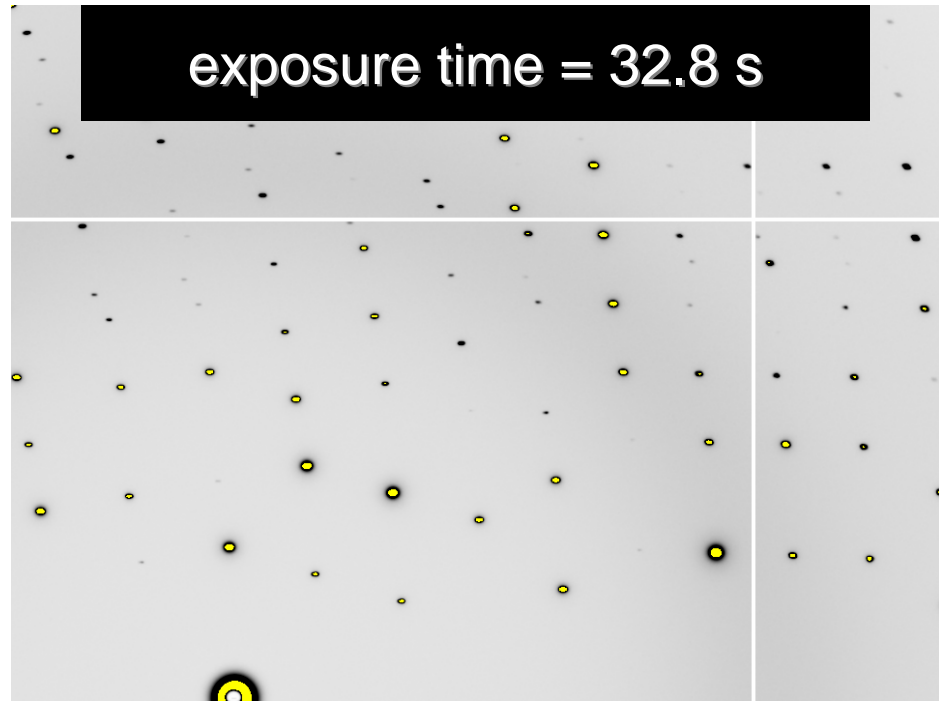
exposure time = 8.2 s



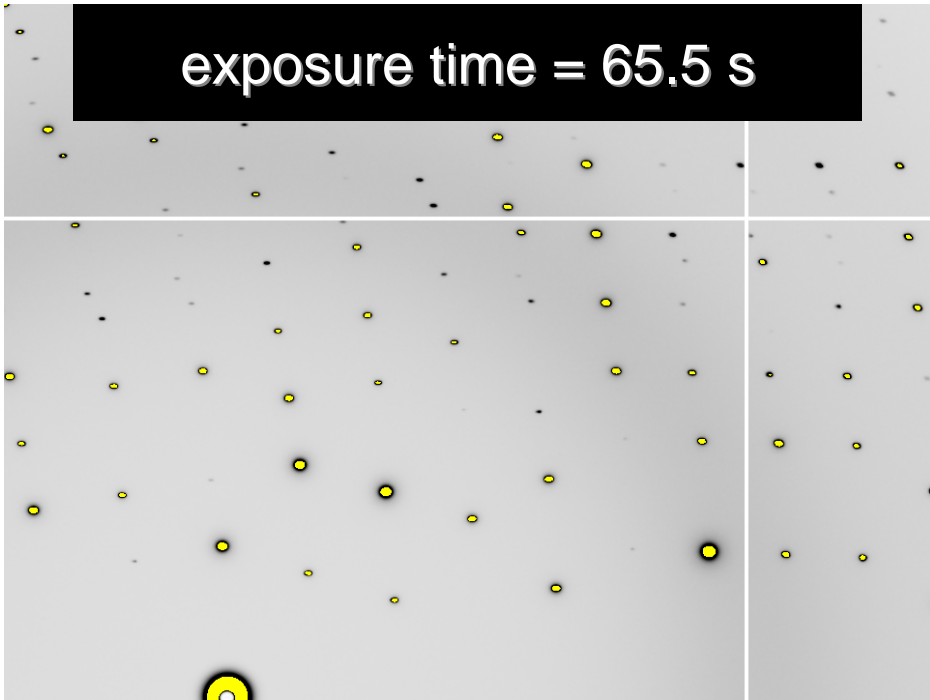
exposure time = 16.4 s



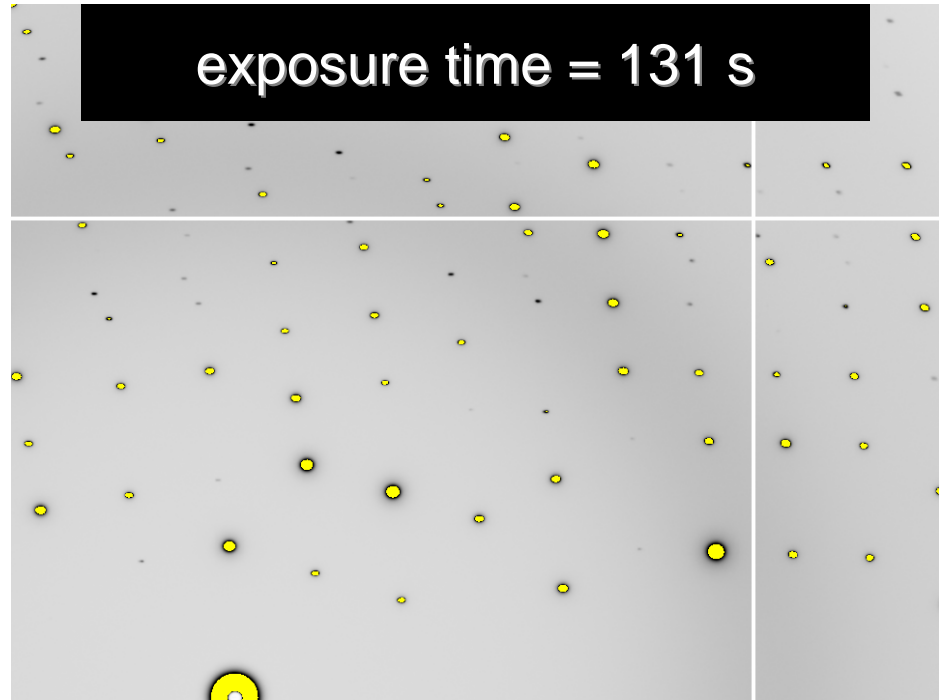
exposure time = 32.8 s



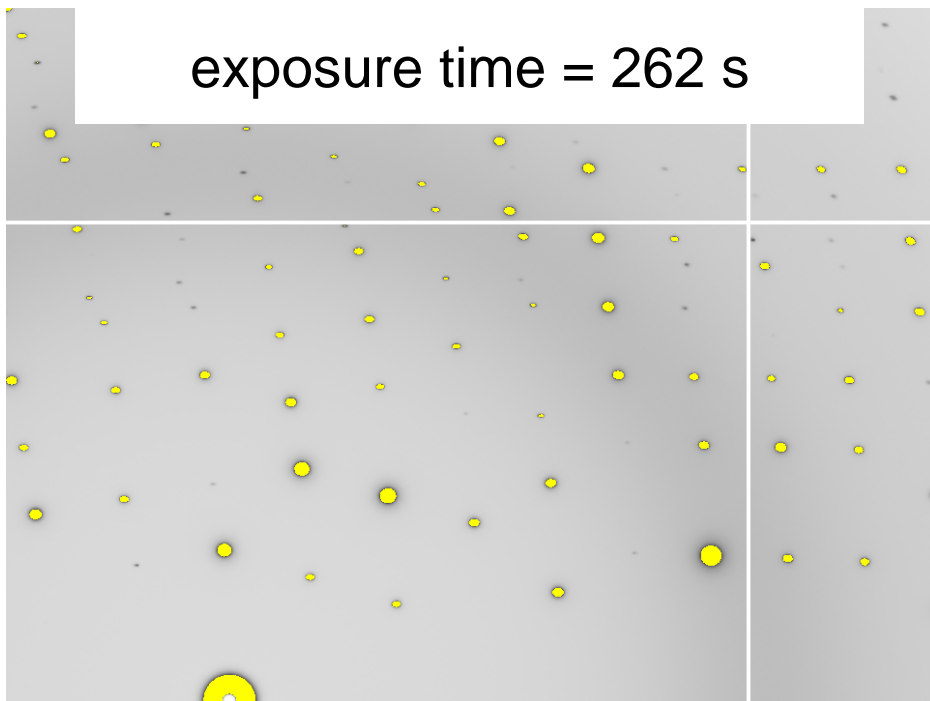
exposure time = 65.5 s



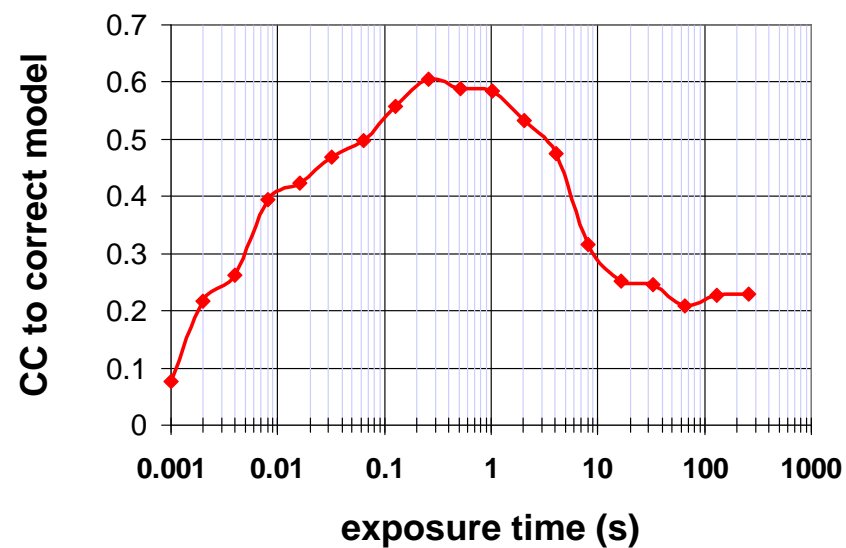
exposure time = 131 s



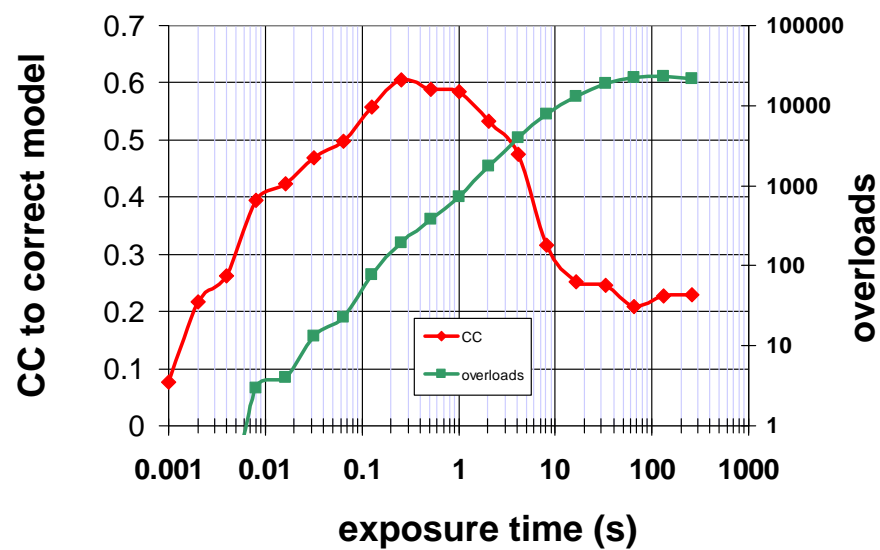
exposure time = 262 s



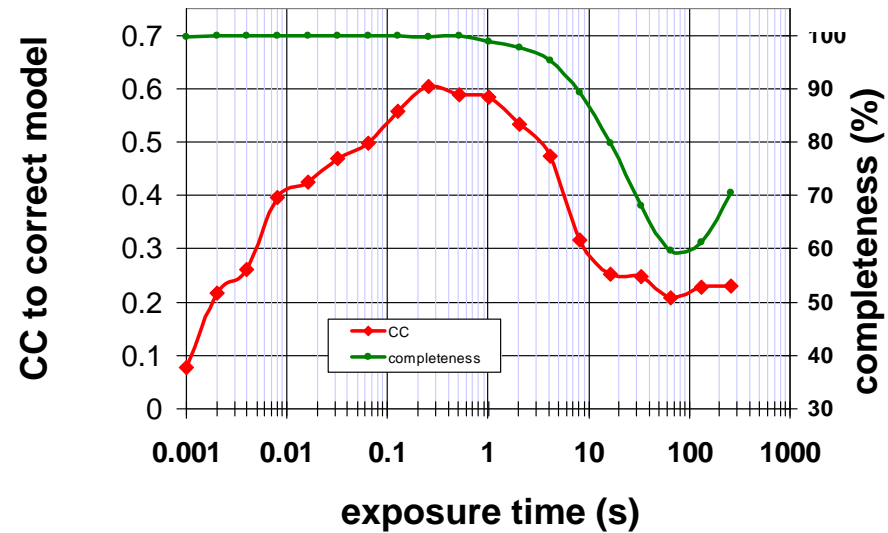
Exposure time



Exposure time



Exposure time



Darwin's Formula

$$I(hkl) = I_{\text{beam}} r_e^2 \frac{V_{\text{xtal}}}{V_{\text{cell}}} \frac{\lambda^3 L}{\omega V_{\text{cell}}} P A | F(hkl) |^2$$

I(hkl)	- photons/spot (fully-recorded)	ω	- rotation speed (radians/s)
I_{beam}	- incident (photons/s/m ²)	L	- Lorentz factor (speed/speed)
r_e	- classical electron radius (2.818x10 ⁻¹⁵ m)	P	- polarization factor (1+cos ² (2θ) -Pfac·cos(2Φ)sin ² (2θ))/2
V_{xtal}	- volume of crystal (in m ³)	A	- absorption factor exp(-μ _{xtal} ·l _{path})
V_{cell}	- volume of unit cell (in m ³)	F(hkl)	- structure amplitude (electrons)
λ	- x-ray wavelength (in meters!)		C. G. Darwin (1914)

Darwin's Formula

$$I(hkl) = I_{\text{beam}} r_e^2 \frac{V_{\text{xtal}}}{V_{\text{cell}}} \frac{\lambda^3 L}{\omega V_{\text{cell}}} P A | F(hkl) |^2$$

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V_{xtal}	- volume of crystal (in m ³)	A	- absorption factor exp(-μ _{xtal} ·l _{path})
V_{cell}	- volume of unit cell (in m ³)	F(hkl)	- structure amplitude (electrons)
λ	- x-ray wavelength (in meters!)		C. G. Darwin (1914)

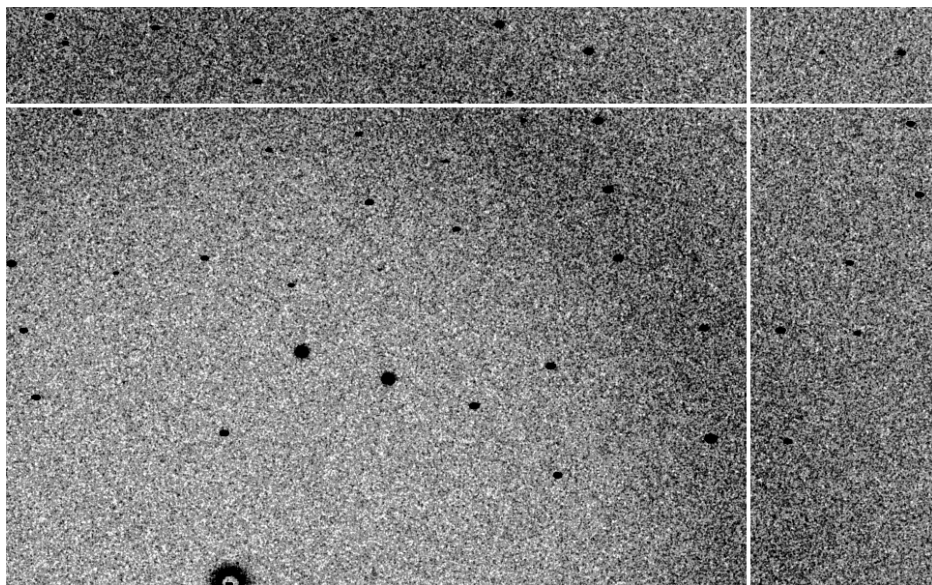
Darwin's Formula

$$I(hkl) = I_{\text{beam}} r_e^2 \frac{V_{\text{xtal}}}{V_{\text{cell}}} \frac{\lambda^3 L}{\omega V_{\text{cell}}} P A | F(hkl) |^2$$

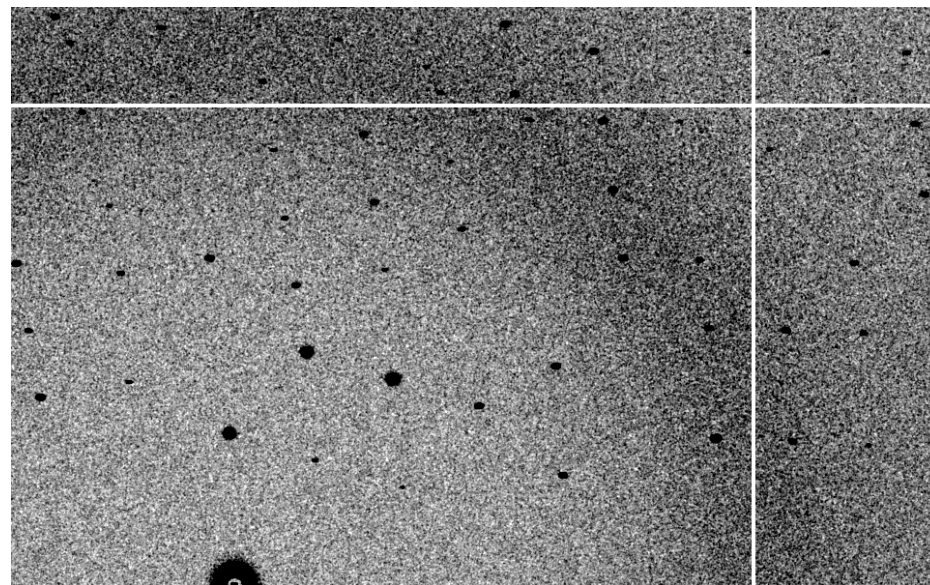
I(hkl)	- photons/spot (fully-recorded)	ω	- rotation speed (radians/s)
I_{beam}	- incident (photons/s/m ²)	L	- Lorentz factor (speed/speed)
r_e	- classical electron radius (2.818x10 ⁻¹⁵ m)	P	- polarization factor (1+cos ² (2θ) -Pfac·cos(2Φ)sin ² (2θ))/2
V_{xtal}	- volume of crystal (in m ³)	A	- absorption factor exp(-μ _{xtal} ·l _{path})
V_{cell}	- volume of unit cell (in m ³)	F(hkl)	- structure amplitude (electrons)
λ	- x-ray wavelength (in meters!)		C. G. Darwin (1914)

How small can the crystal be?

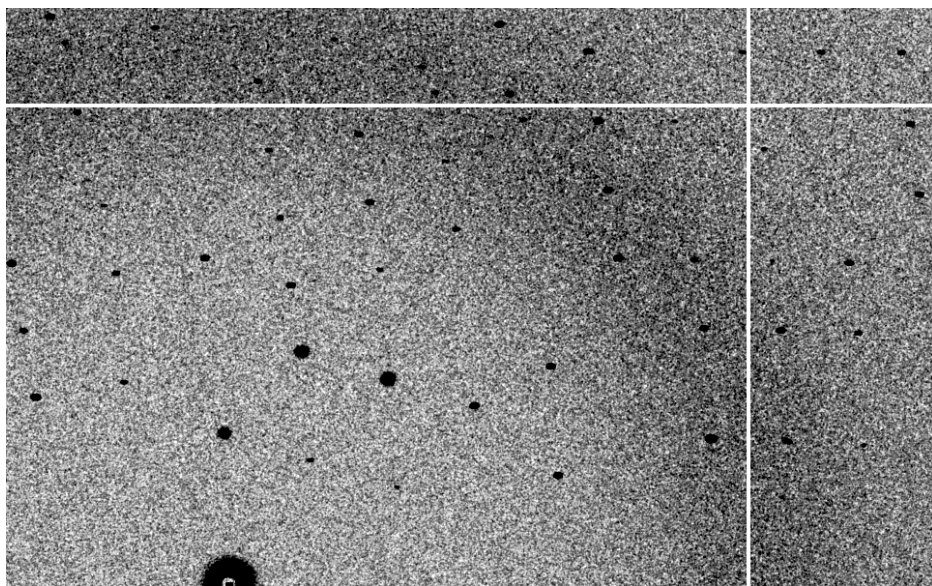
crystal size = 1 mm



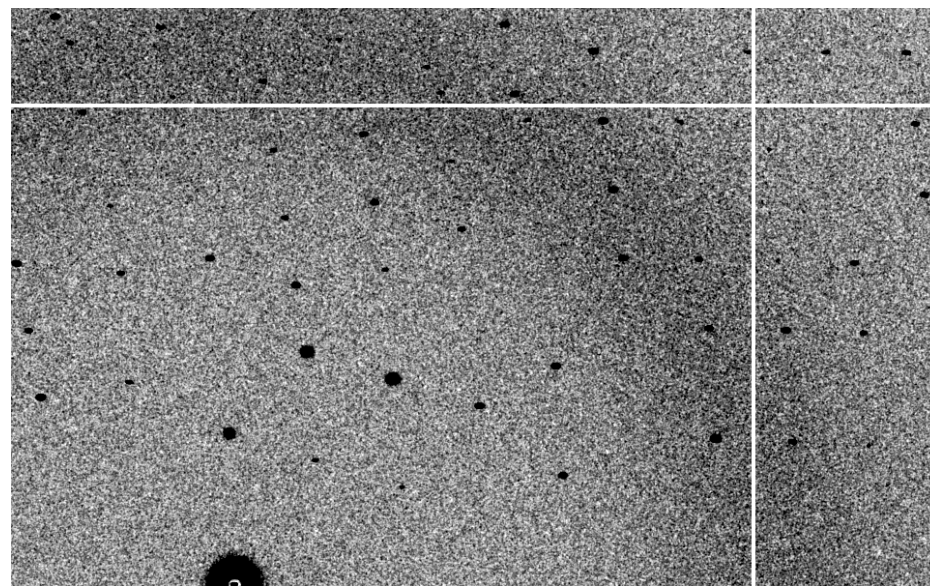
crystal size = 512 μm



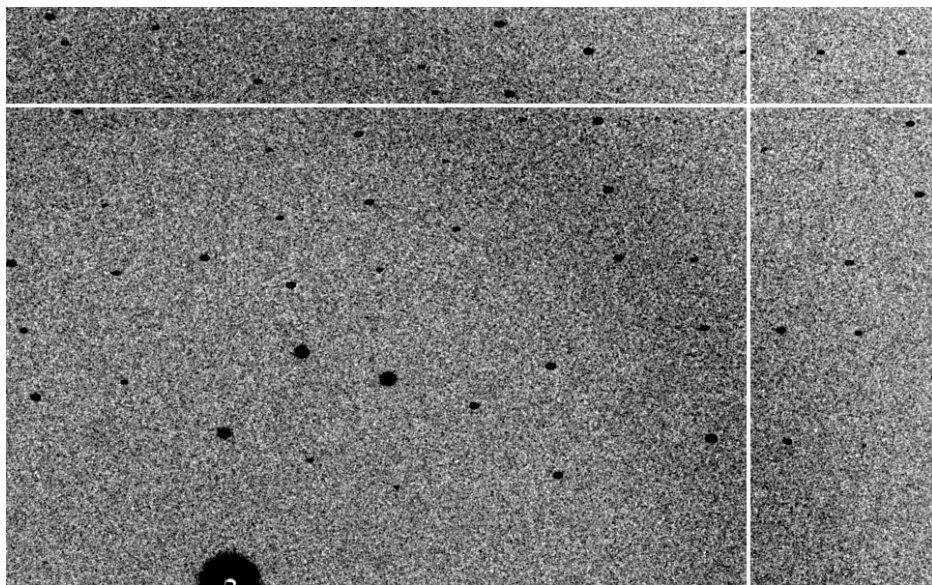
crystal size = 256 μm



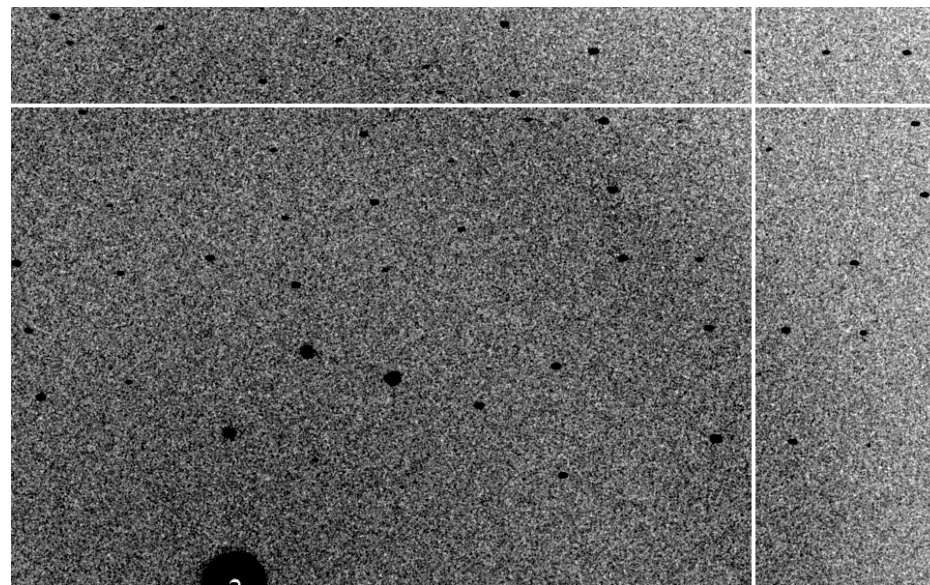
crystal size = 128 μm



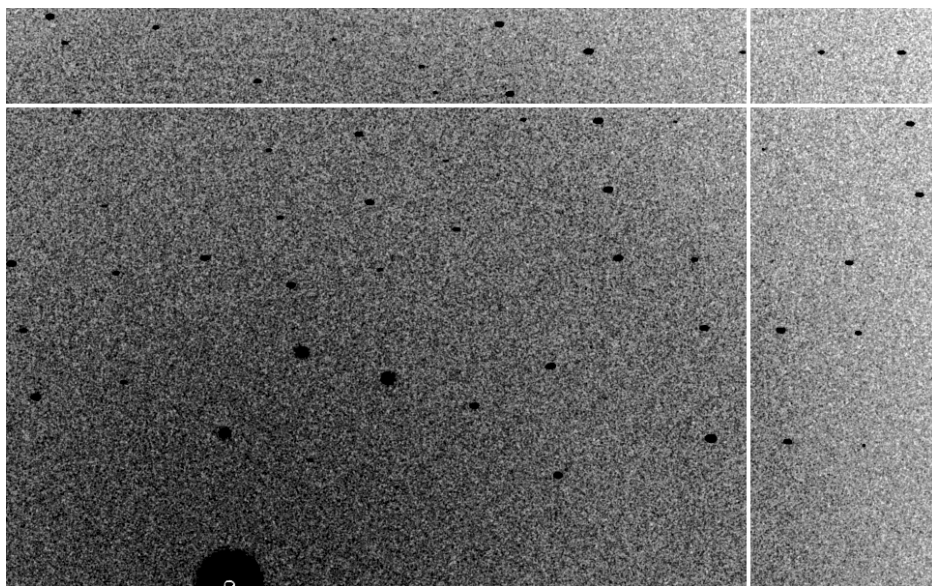
crystal size = 64 μm



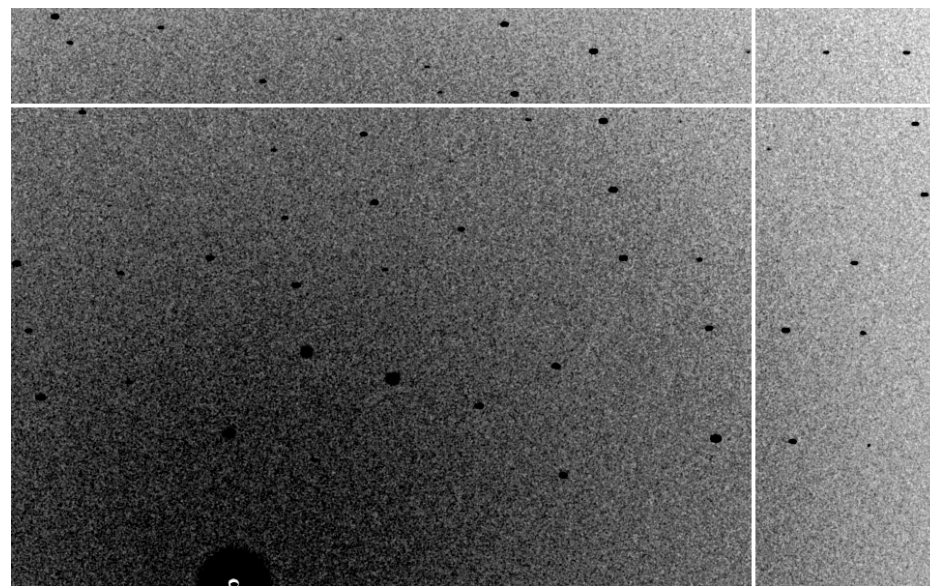
crystal size = 32 μm



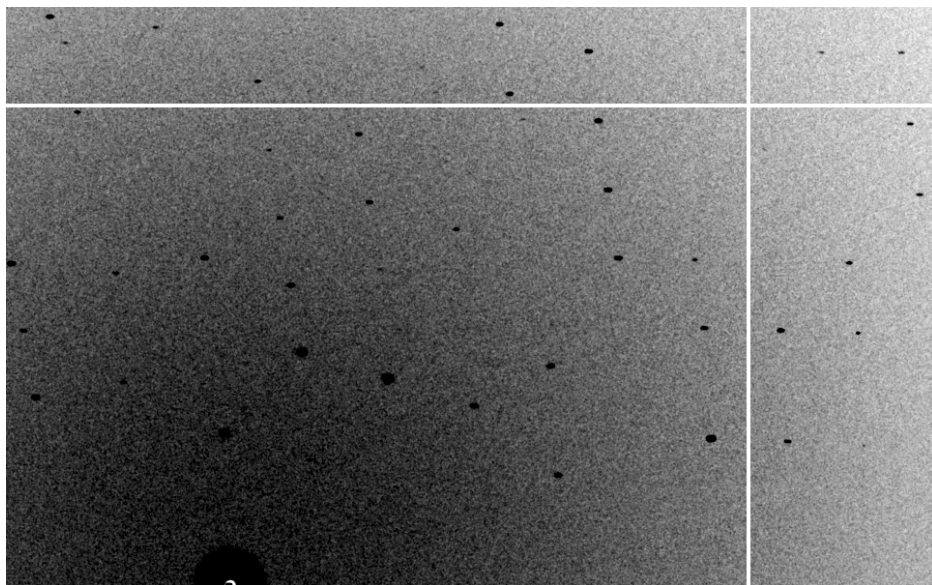
crystal size = 16 μm



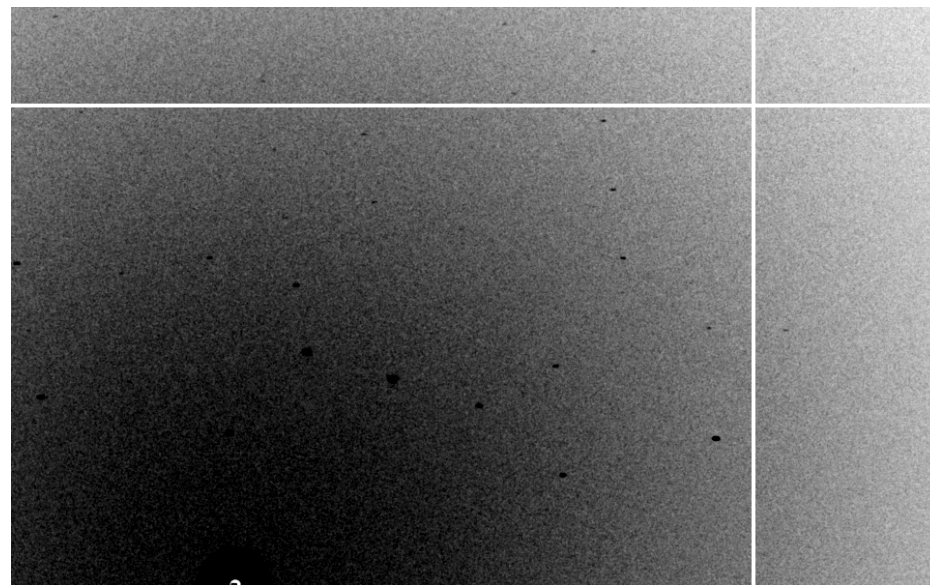
crystal size = 8 μm



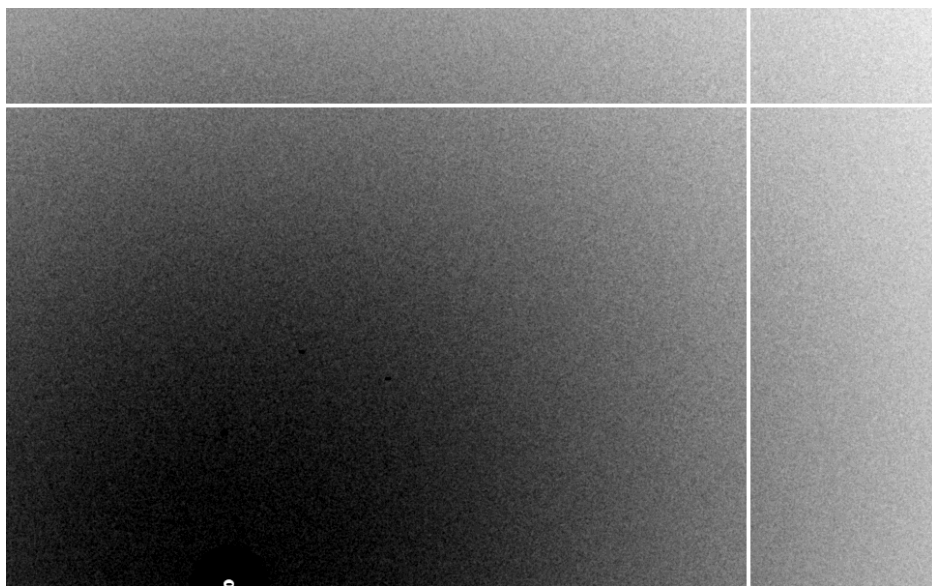
crystal size = 4 μm



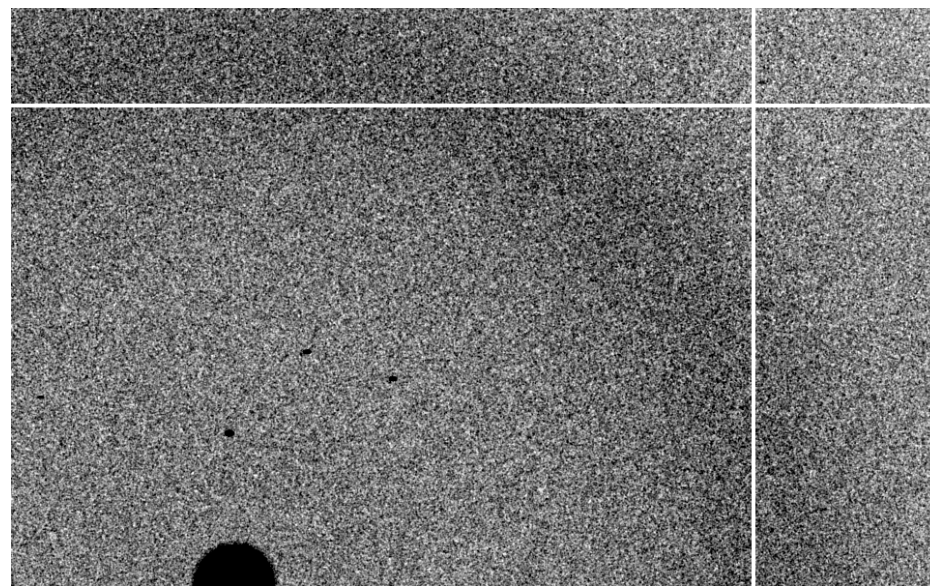
crystal size = 2 μm



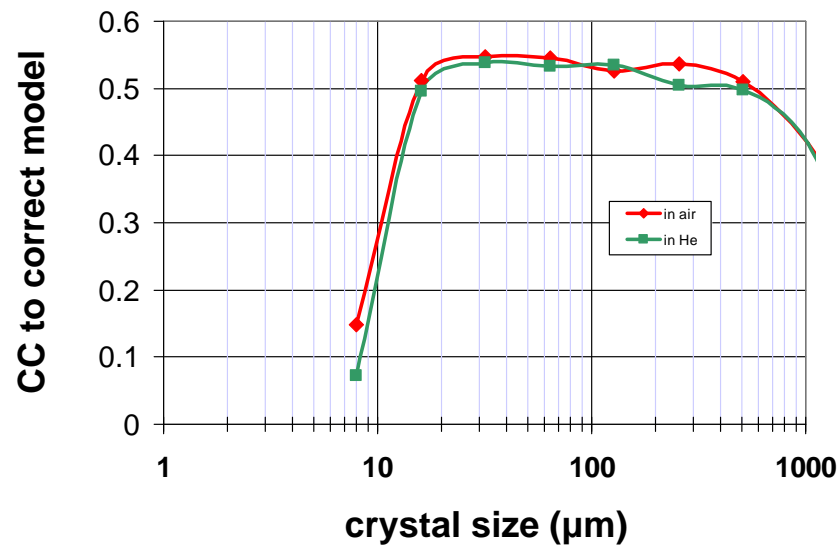
crystal size = 1 μm



crystal size = 1 μm (in He)



Crystal Size



Darwin's Formula

$$I(hkl) = I_{\text{beam}} r_e^2 \frac{V_{\text{xtal}}}{V_{\text{cell}}} \frac{\lambda^3 L}{\omega V_{\text{cell}}} P A |F(hkl)|^2$$

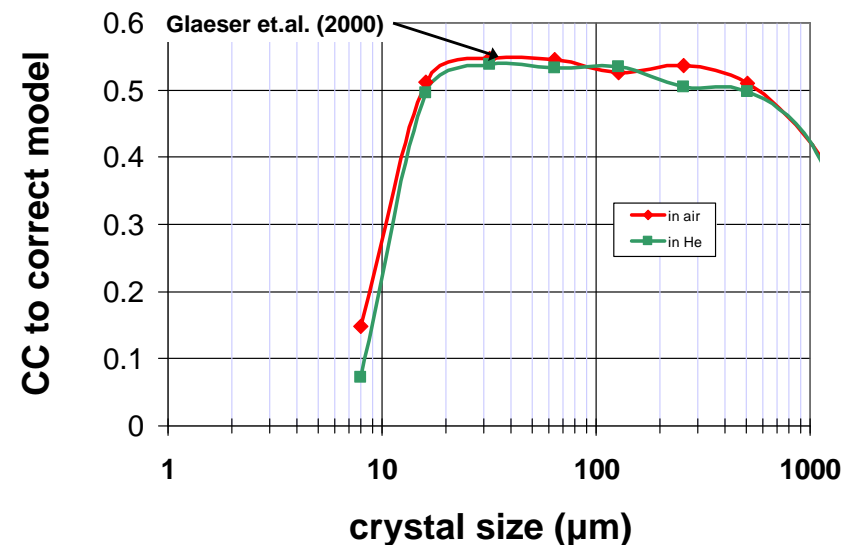
$I(hkl)$	- photons/spot (fully-recorded)	ω	- rotation speed (radians/s)
I_{beam}	- incident (photons/s/m ²)	L	- Lorentz factor (speed/speed)
r_e	- classical electron radius (2.818x10 ⁻¹⁵ m)	P	- polarization factor (1+cos ² (2θ) -Pfac·cos(2Φ)sin ² (2θ))/2
V_{xtal}	- volume of crystal (in m ³)	A	- absorption factor exp(-μ _{xtal} ·l _{path})
V_{cell}	- volume of unit cell (in m ³)	$F(hkl)$	- structure amplitude (electrons)
λ	- x-ray wavelength (in meters!)		C. G. Darwin (1914)

Darwin's Formula

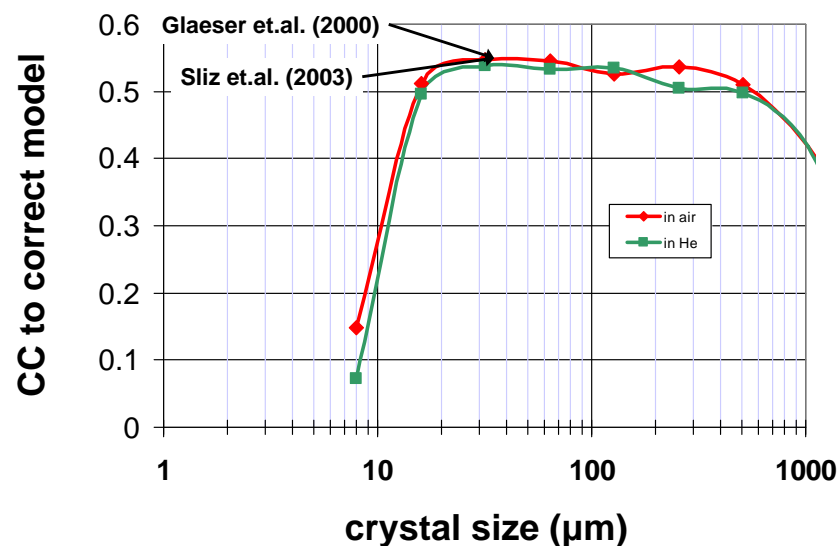
$$I(hkl) = I_{\text{beam}} r_e^2 \frac{V_{\text{xtal}}}{V_{\text{cell}}} \frac{\lambda^3 L}{\omega V_{\text{cell}}} P A |F(hkl)|^2$$

$I(hkl)$	- photons/spot (fully-recorded)	ω	- rotation speed (radians/s)
I_{beam}	- incident (photons/s/m ²)	L	- Lorentz factor (speed/speed)
r_e	- classical electron radius (2.818x10 ⁻¹⁵ m)	P	- polarization factor (1+cos ² (2θ) -Pfac·cos(2Φ)sin ² (2θ))/2
V_{xtal}	- volume of crystal (in m ³)	A	- absorption factor exp(-μ _{xtal} ·l _{path})
V_{cell}	- volume of unit cell (in m ³)	$F(hkl)$	- structure amplitude (electrons)
λ	- x-ray wavelength (in meters!)		C. G. Darwin (1914)

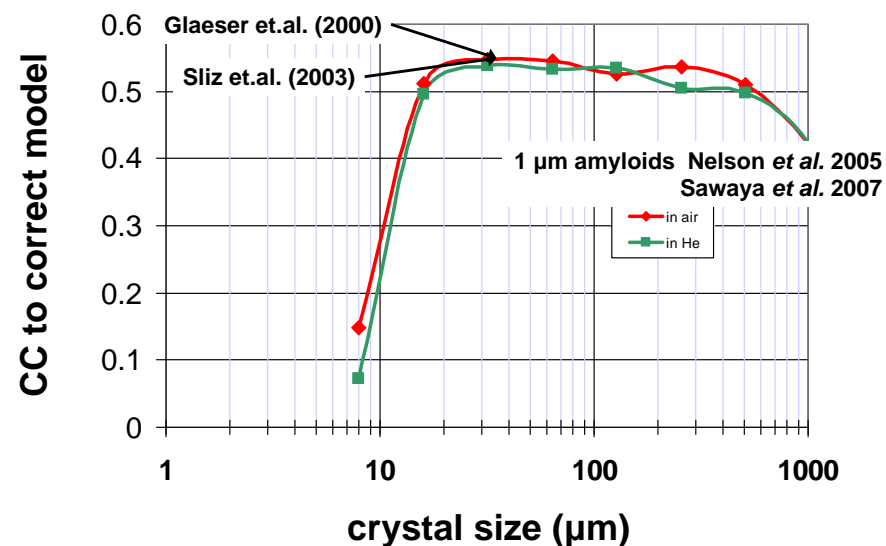
Crystal Size



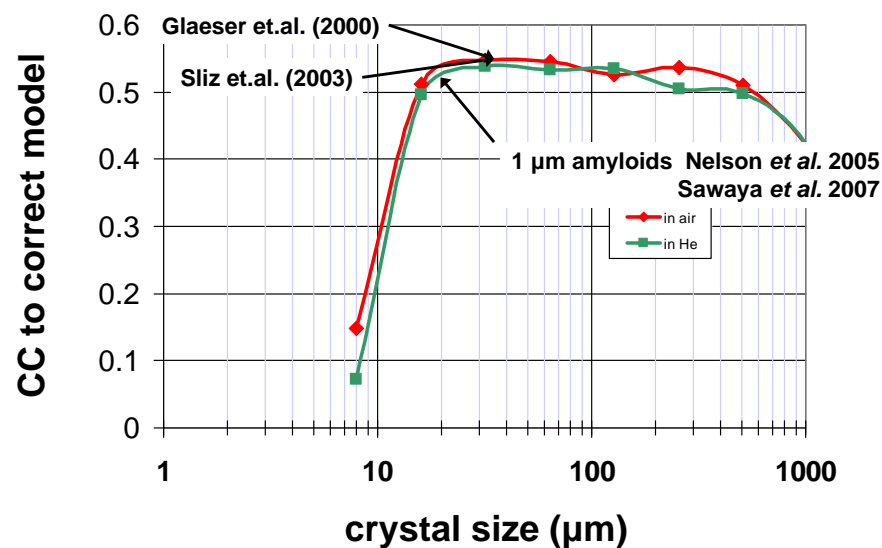
Crystal Size



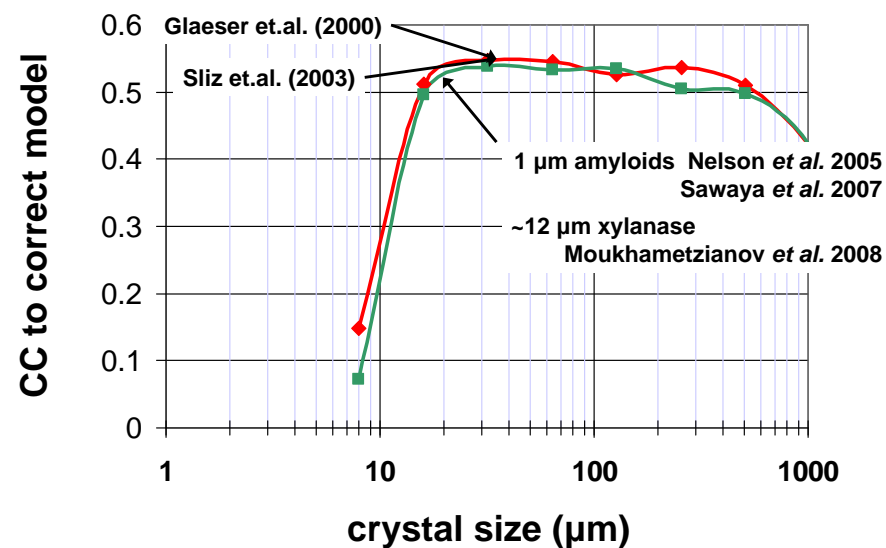
Crystal Size



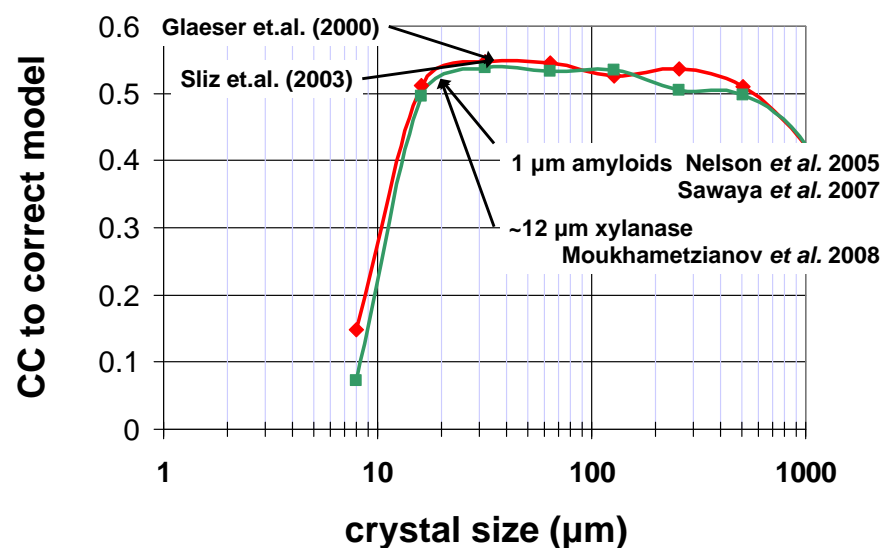
Crystal Size



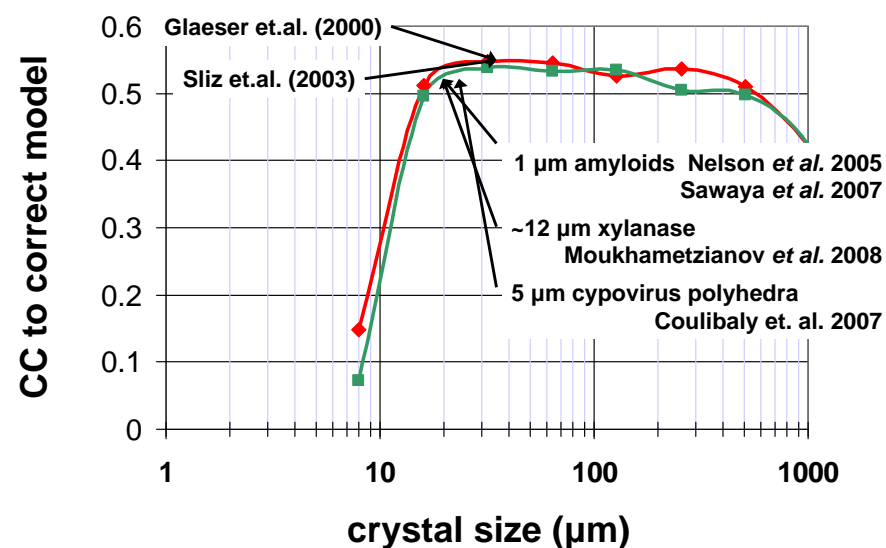
Crystal Size



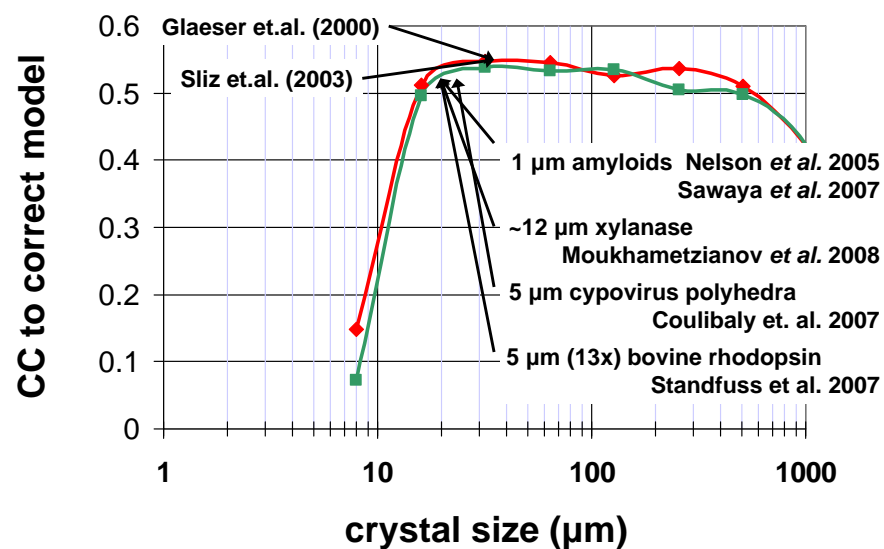
Crystal Size



Crystal Size



Crystal Size



Minimum Crystal Size

$$n_{\text{xtal}} = n_0 \frac{\text{MW } V_M^2}{\ell_x \ell_y \ell_z (d^3 - 1.53) \exp(-0.5 B/d^2)}$$

n_{xtal} - number of crystals needed

n_0 - empirical constant (~ 3)

MW - molecular weight (kDa)

d - d-spacing of interest (Å)

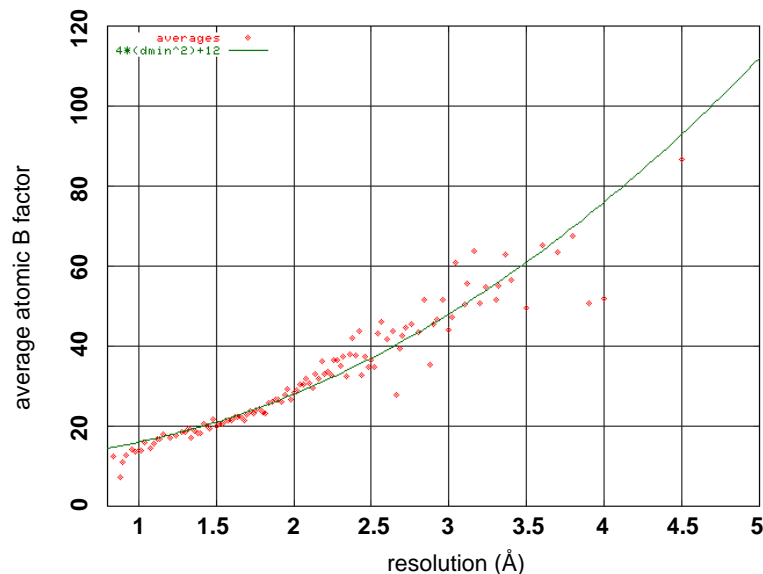
V_M - Matthews number (~2.5 Å³/Da)

B - Wilson B factor (Å²)

ℓ - crystal size (microns)

$$B \approx 4 d^2 + 12$$

$$B \approx 4 d^2 + 12$$



Minimum Crystal Size

<http://bl831.als.lbl.gov/~jamesh/xtalsize.html>

$$n_{\text{xtal}} = n_0 \frac{MW V_M^2}{\ell_x \ell_y \ell_z (d^3 - 1.53) \exp(-0.5 B/d^2)}$$

n_{xtal} - number of crystals needed

n_0 - empirical constant (~ 3)

MW - molecular weight (kDa)

d - d-spacing of interest (Å)

V_M - Matthews number (~2.5 Å³/Da)

B - Wilson B factor (Å²)

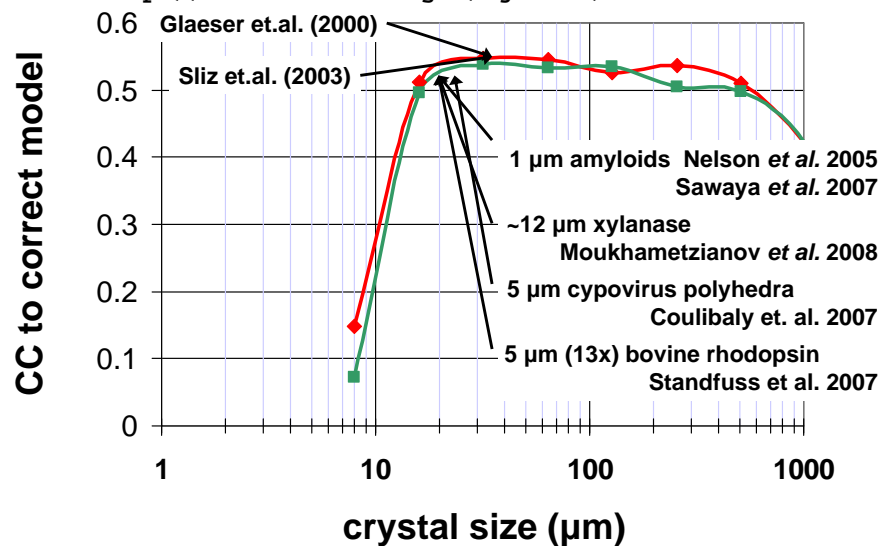
ℓ - crystal size (microns)

$$B \approx 4 d^2 + 12$$

Holton J. M. (2009) *J. Synchrotron Rad.* **16** 133-42

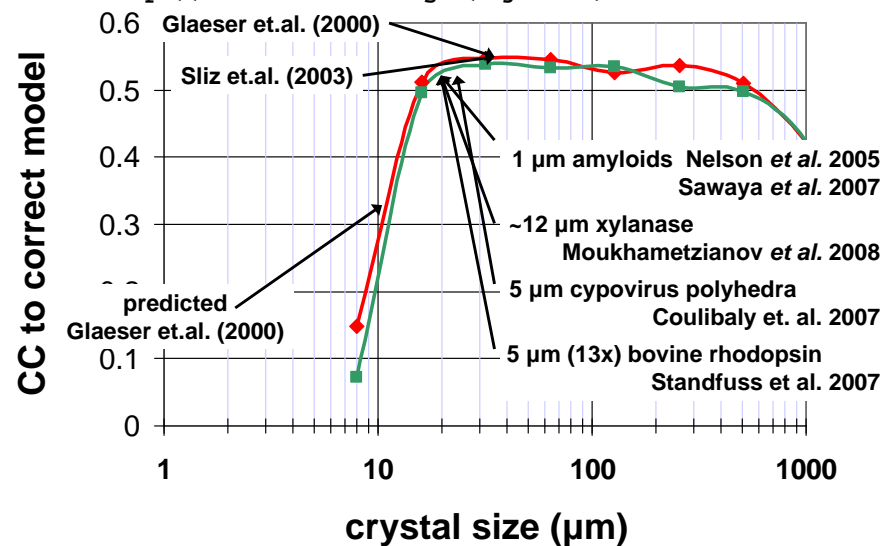
Crystal Size

<http://bl831.als.lbl.gov/~jamesh/xtalsize.html>

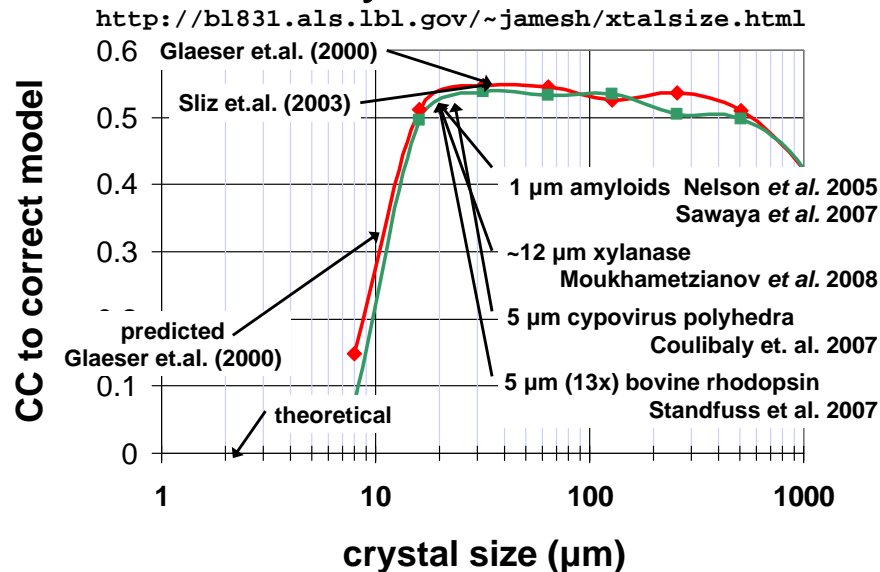


Crystal Size

<http://bl831.als.lbl.gov/~jamesh/xtalsize.html>



Crystal Size



Theoretical limit for lysozyme:

$$\langle I_{\max} \rangle (s) \cong \frac{8 \cdot 10^{29} r_e^2 \rho R^4 \lambda^4}{9 \pi h c H s n_{\text{ASU}} M_r V_M^2} \left(\frac{1.05 \cdot \pi}{4 s \lambda} - \frac{1}{2} \right) \cdot \frac{T_{\text{sphere}}(2\theta, \mu, R)}{1 - T_{\text{sphere}}(0, \mu_{\text{en}}, R)} \cdot \frac{\langle f_a(s) \rangle^2 \exp(-2Bs^2)}{\langle M_a \rangle}$$

Where:

- $\langle I_{\max} \rangle (s)$ - average damage-limited spot intensity (photons) at "s"
- s - $\sin(\theta)/\lambda$ or $0.5/d$ if d is the d-spacing of interest (\AA^{-1})
- θ - Bragg angle
- r_e - classical electron radius ($2.818 \times 10^{-15} \text{ m}$)
- 10^{29} - conversion from m^4 to \AA^4 , g/cm^3 to kg/m^3 and MGy to Gy
- ρ - density of crystal ($\sim 1.2 \text{ g/cm}^3$)
- R - radius of the spherical crystal (m)
- λ - x-ray wavelength (\AA)
- h - Planck's constant ($6.626 \times 10^{-34} \text{ J}\cdot\text{s}$)
- c - speed of light (299792458 m/s)
- H - Howells's coefficient ($11 \times 10^{-3} \text{ \AA/MGy}$)
- n_{ASU} - number of proteins in the asymmetric unit
- M_r - molecular weight of the protein (Daltons or g/mol)
- V_M - Matthews's coefficient ($\sim 2.4 \text{ \AA}^3/\text{Dalton}$)
- $\langle f_a(s) \rangle^2$ - number-averaged squared protein atom structure factor (electron units²)
- $\langle M_a \rangle$ - number-averaged atomic weight of a protein atom ($\sim 7.1 \text{ Daltons}$)
- B - average (Wilson) temperature factor (\AA^2)
- μ - attenuation coefficient of sphere material (m^{-1})
- μ_{en} - mass energy-absorption coefficient of sphere material (m^{-1})

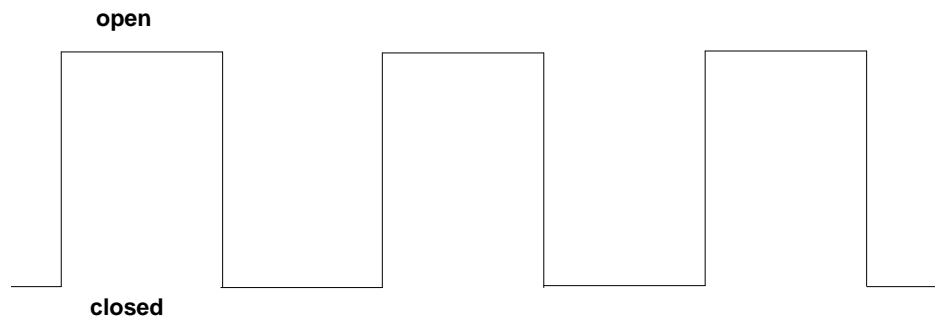
Theoretical limit for lysozyme:

$$1.0 \frac{\text{photon}}{\text{spot } \mu\text{m}^3}$$

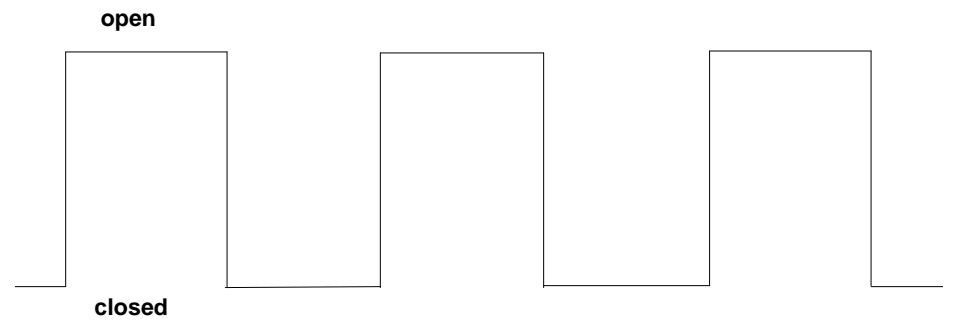
at $\sim 2.4 \text{ \AA}$

What if we had
a "better" beamline?

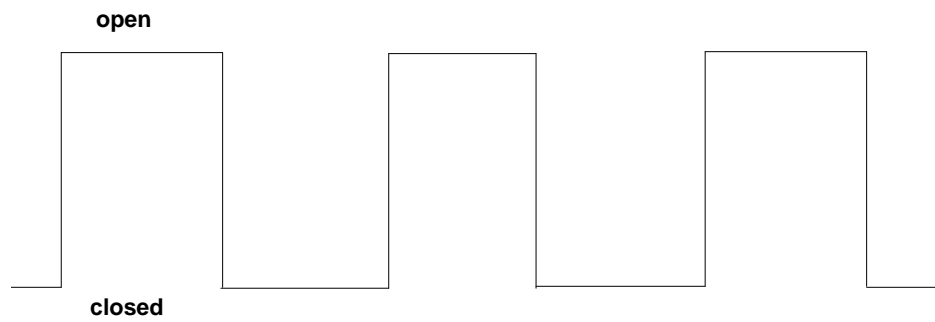
Shutter Jitter



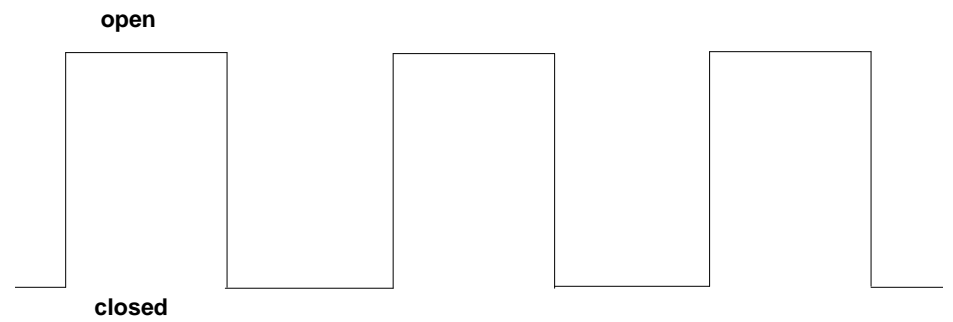
Shutter Jitter



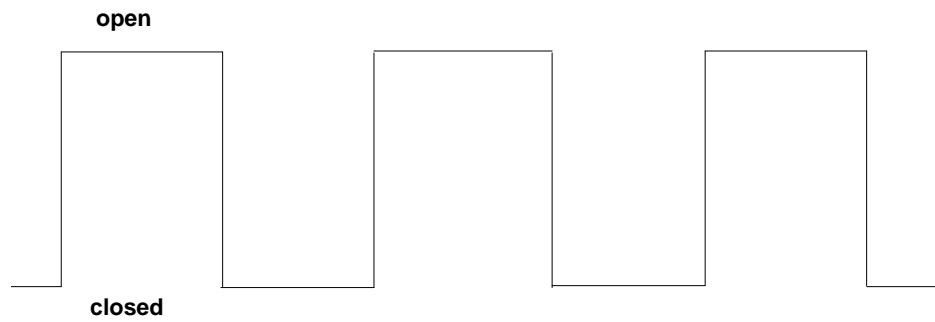
Shutter Jitter



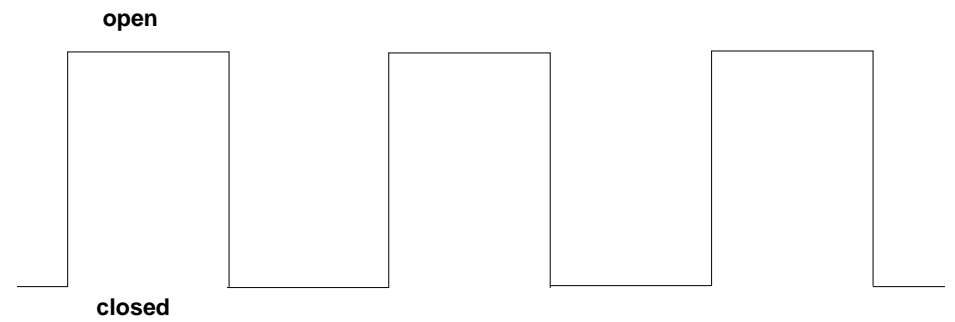
Shutter Jitter



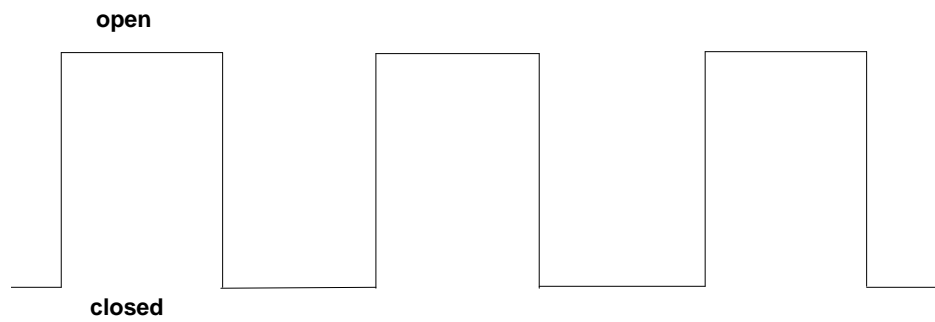
Shutter Jitter



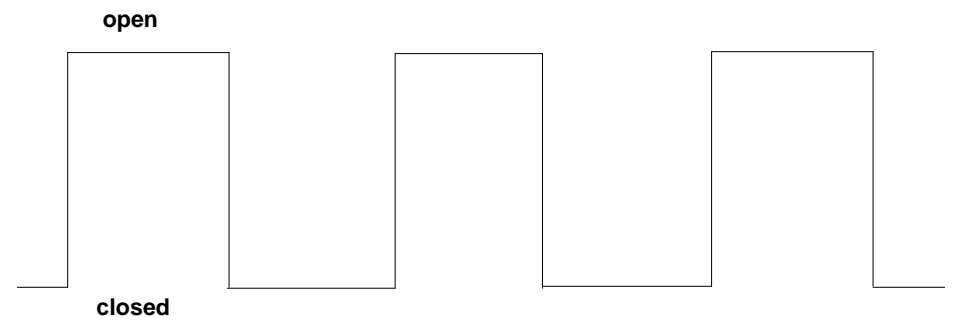
Shutter Jitter



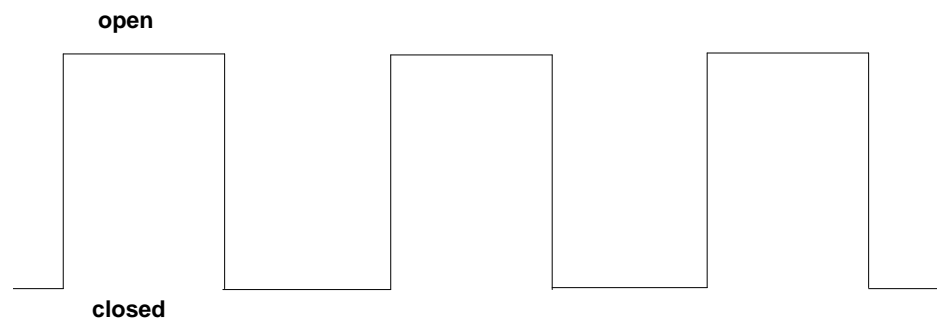
Shutter Jitter



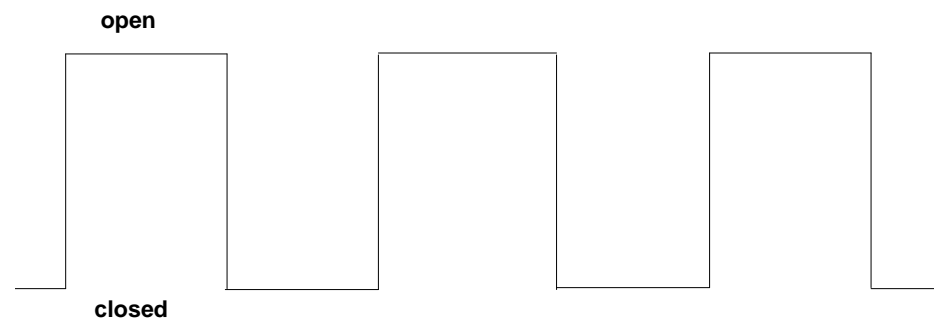
Shutter Jitter



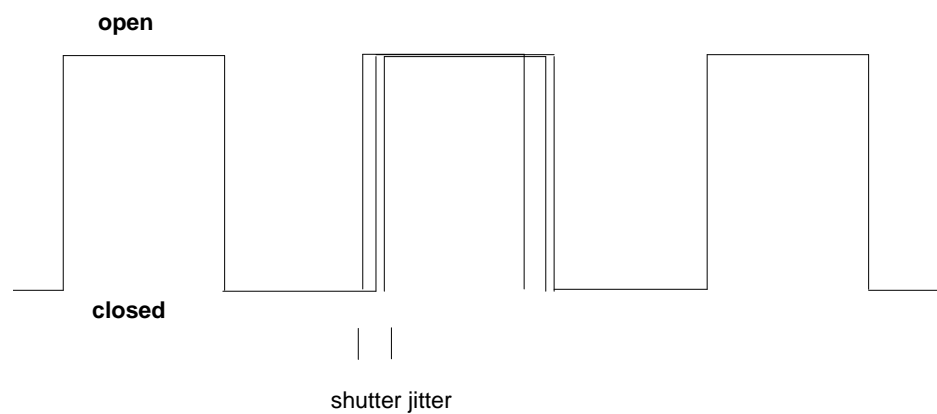
Shutter Jitter



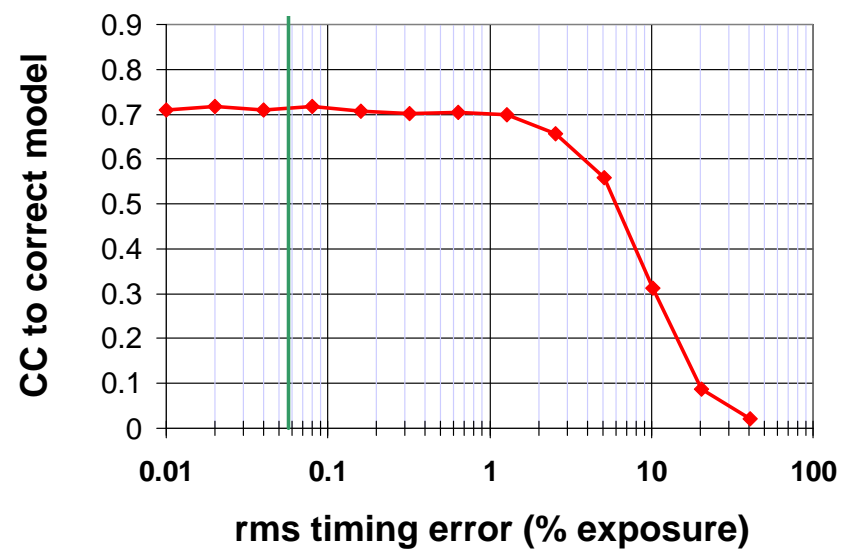
Shutter Jitter



Shutter Jitter



Shutter Jitter



What about
crystal vibration?

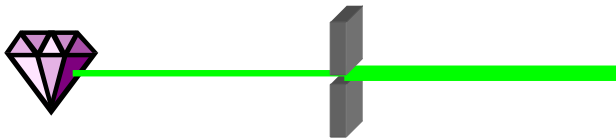
Beam Flicker

1/f noise or “flicker noise”

Beam Flicker

1/f noise or “flicker noise”

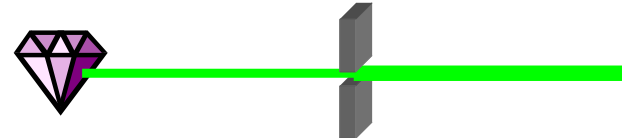
comes from everything



Beam Flicker

1/f noise or “flicker noise”

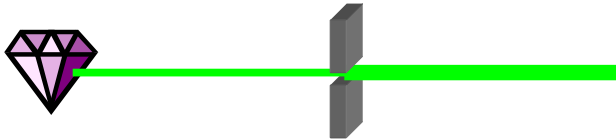
comes from everything



Beam Flicker

1/f noise or “flicker noise”

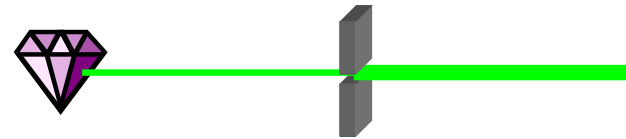
comes from everything



Beam Flicker

1/f noise or “flicker noise”

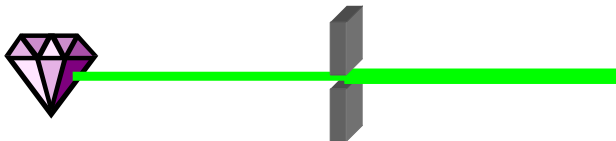
comes from everything



Beam Flicker

1/f noise or “flicker noise”

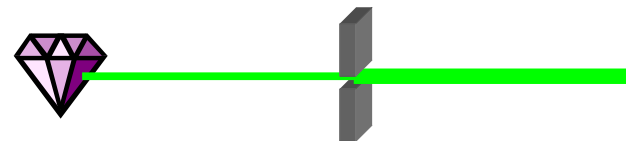
comes from everything



Beam Flicker

1/f noise or “flicker noise”

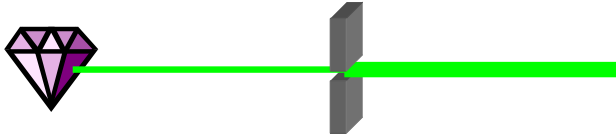
comes from everything



Beam Flicker

1/f noise or “flicker noise”

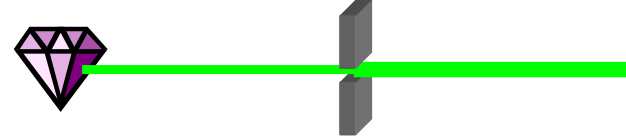
comes from everything



Beam Flicker

1/f noise or “flicker noise”

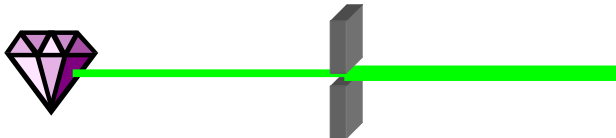
comes from everything



Beam Flicker

1/f noise or “flicker noise”

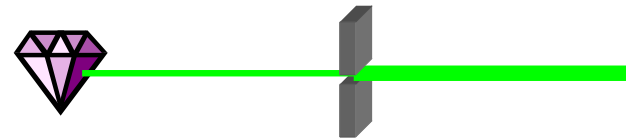
comes from everything



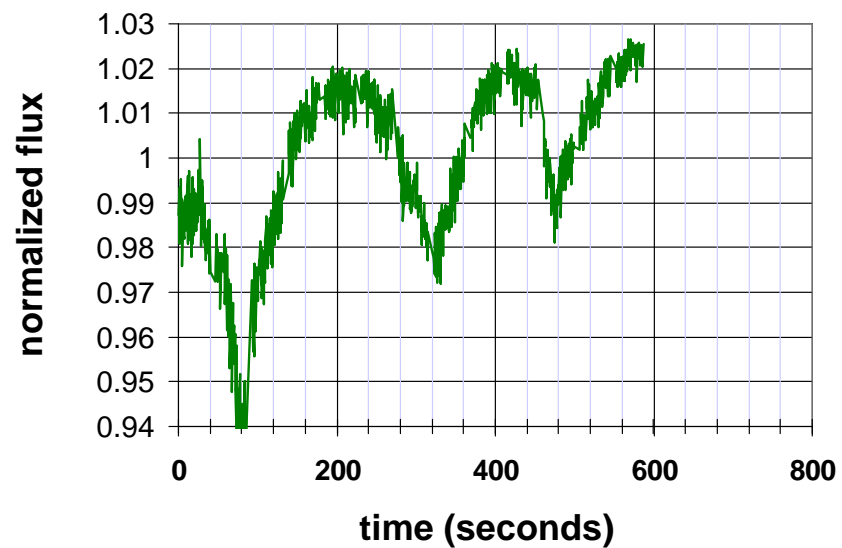
Beam Flicker

1/f noise or “flicker noise”

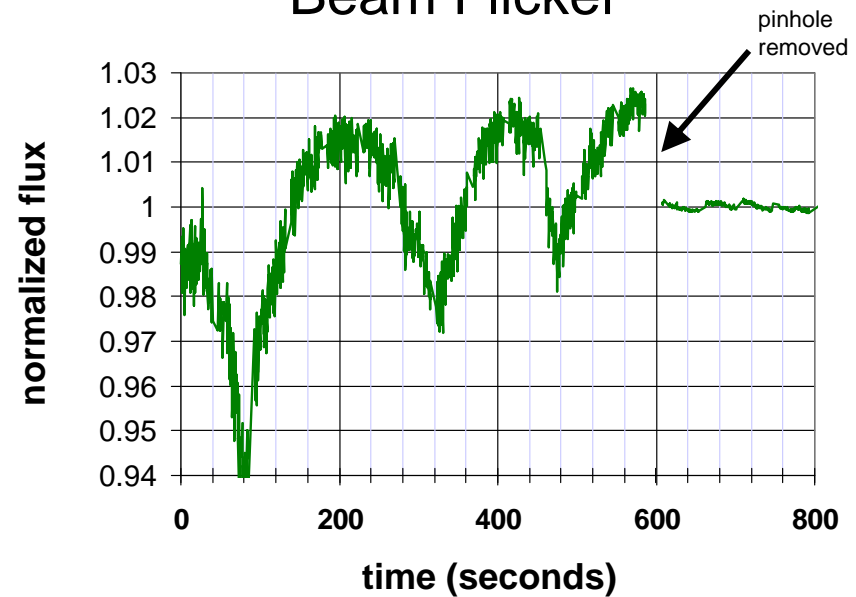
comes from everything



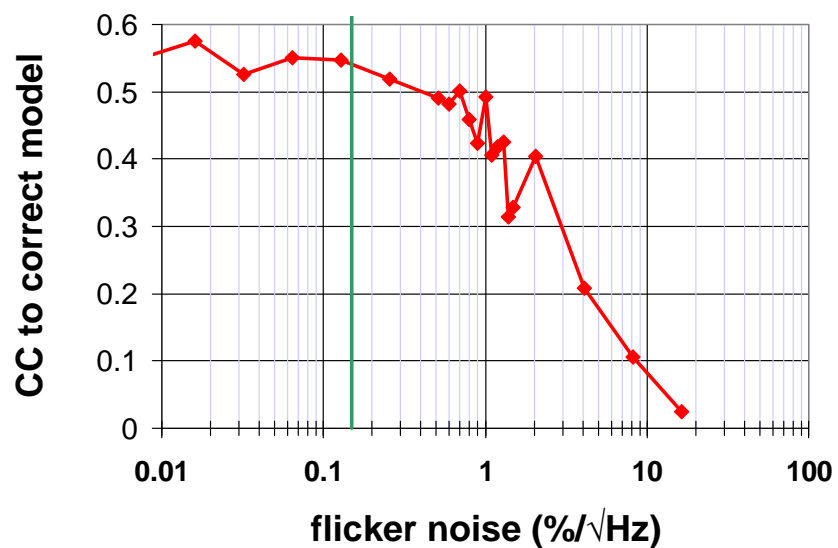
Beam Flicker



Beam Flicker



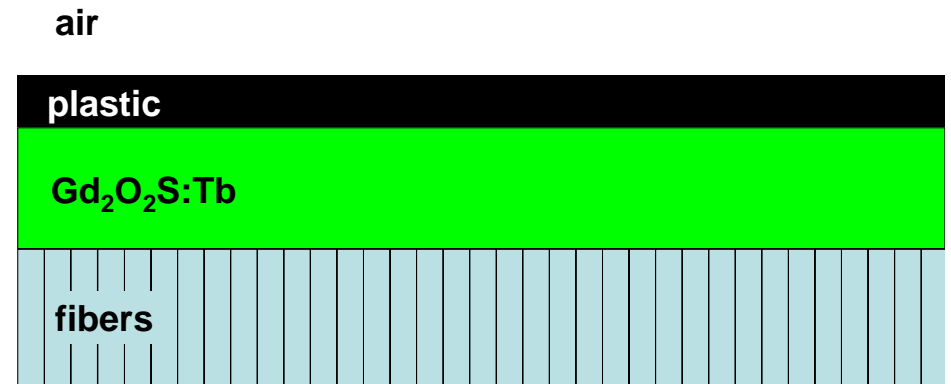
Beam Flicker



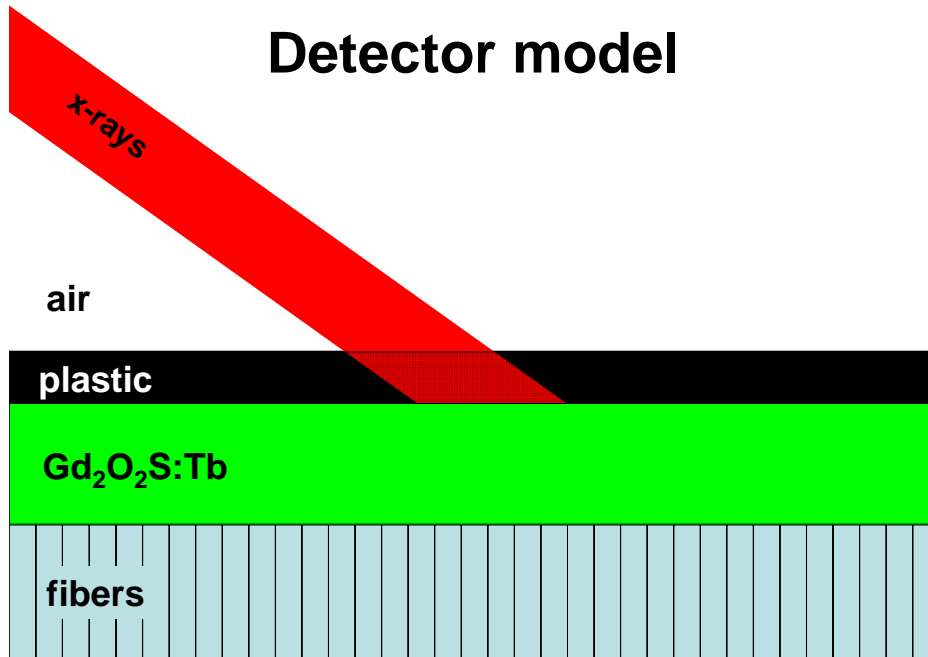
What about a better detector?

What if we had a perfect detector?

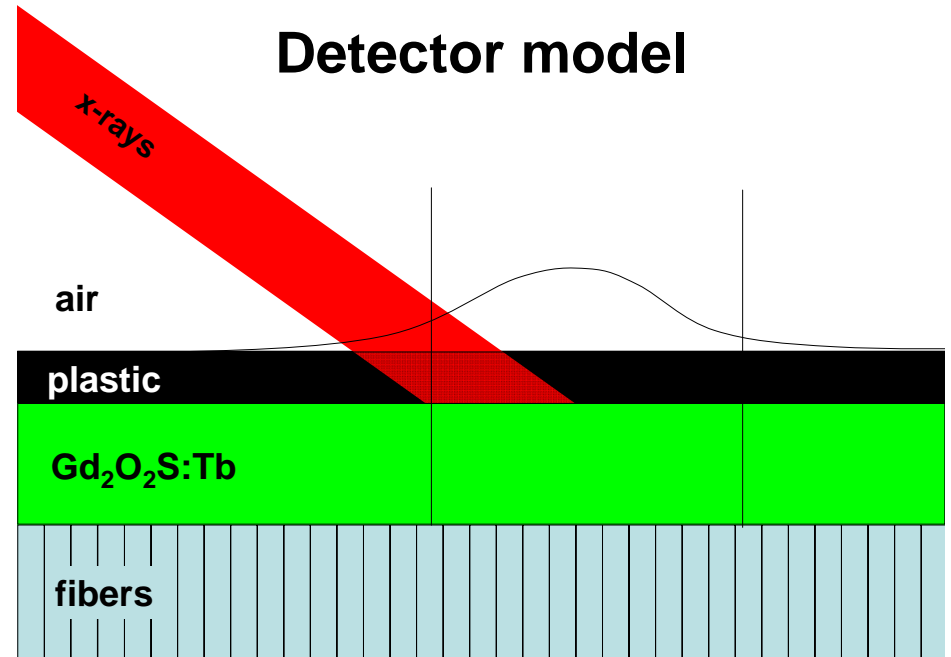
Detector model

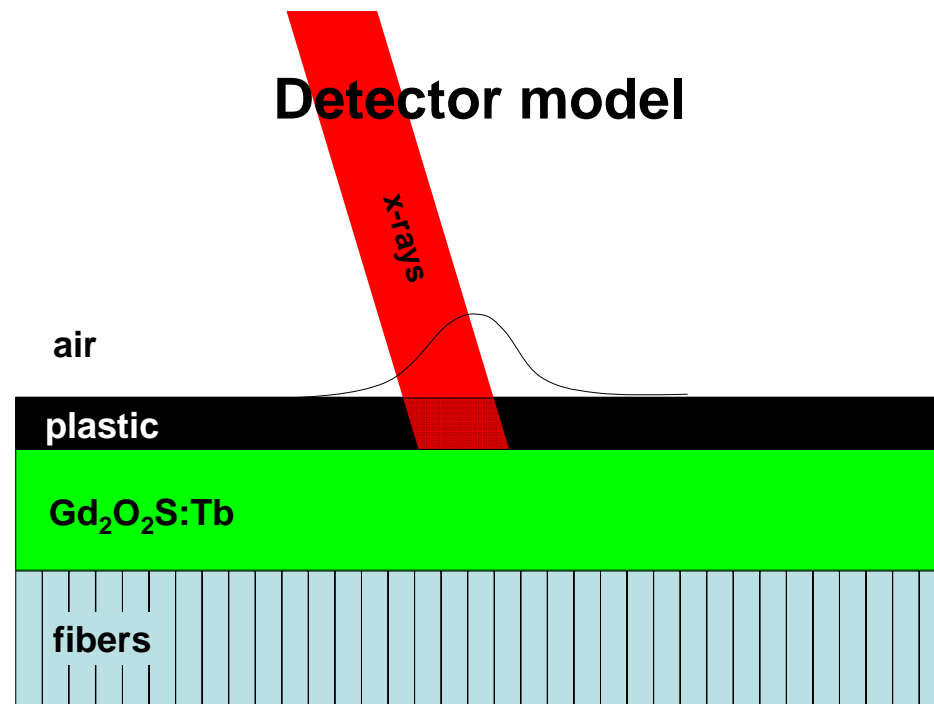
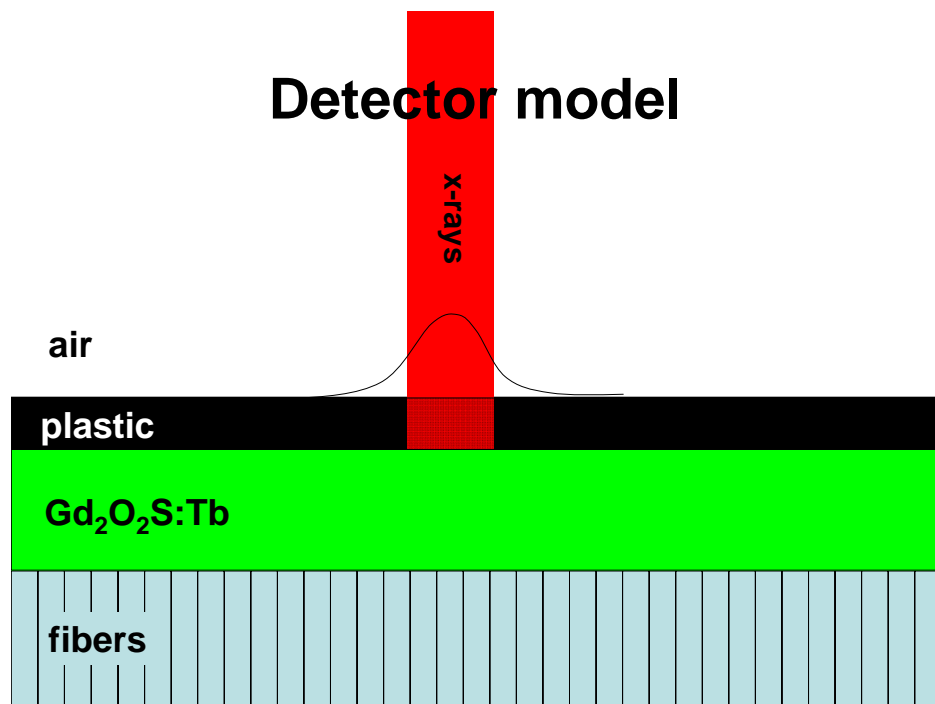
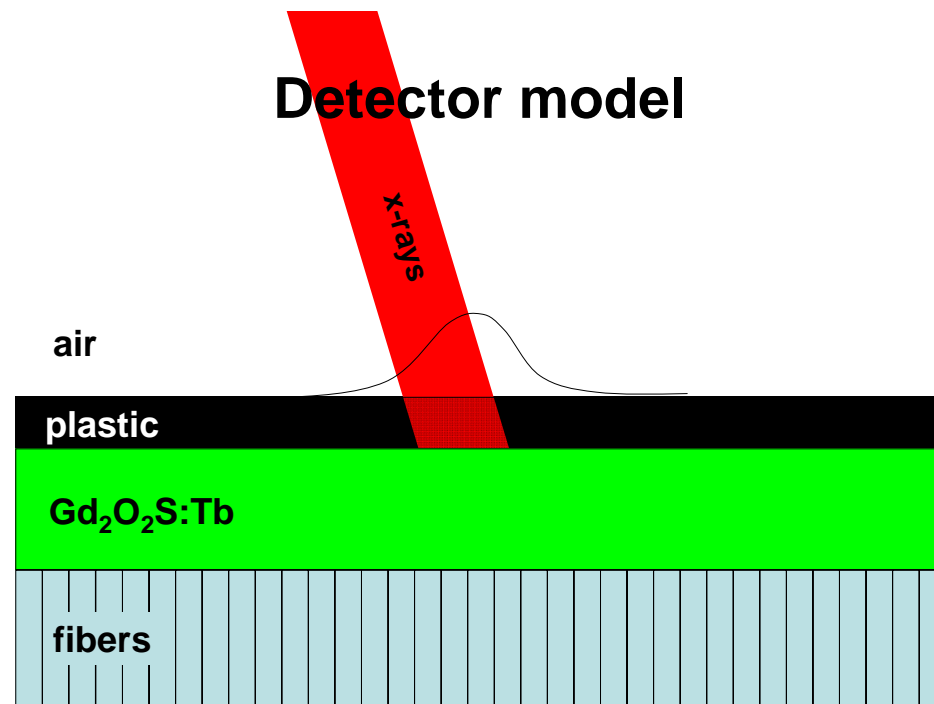
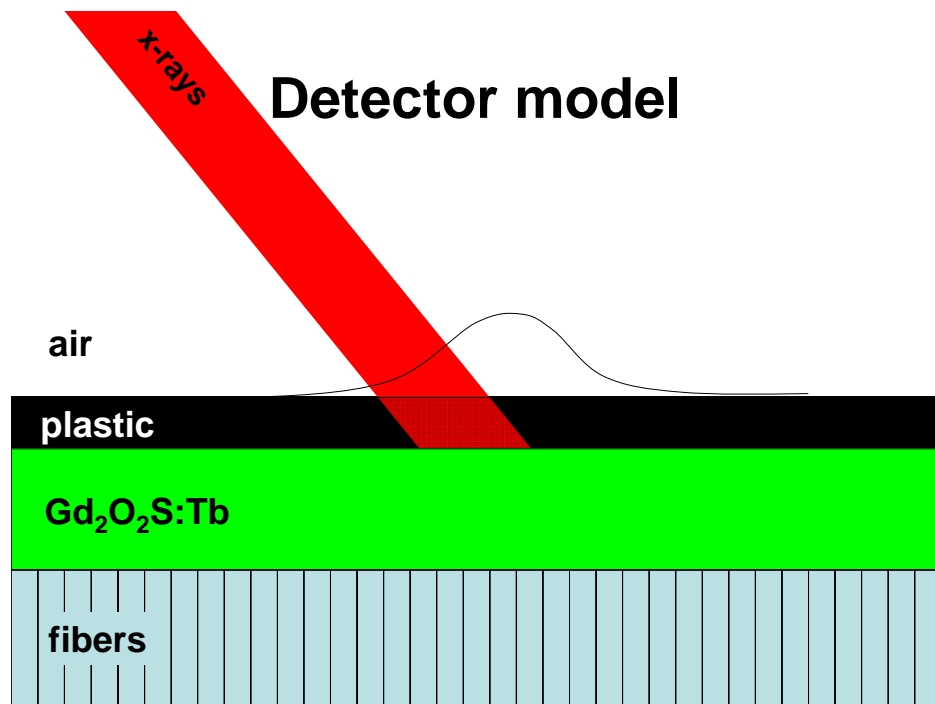


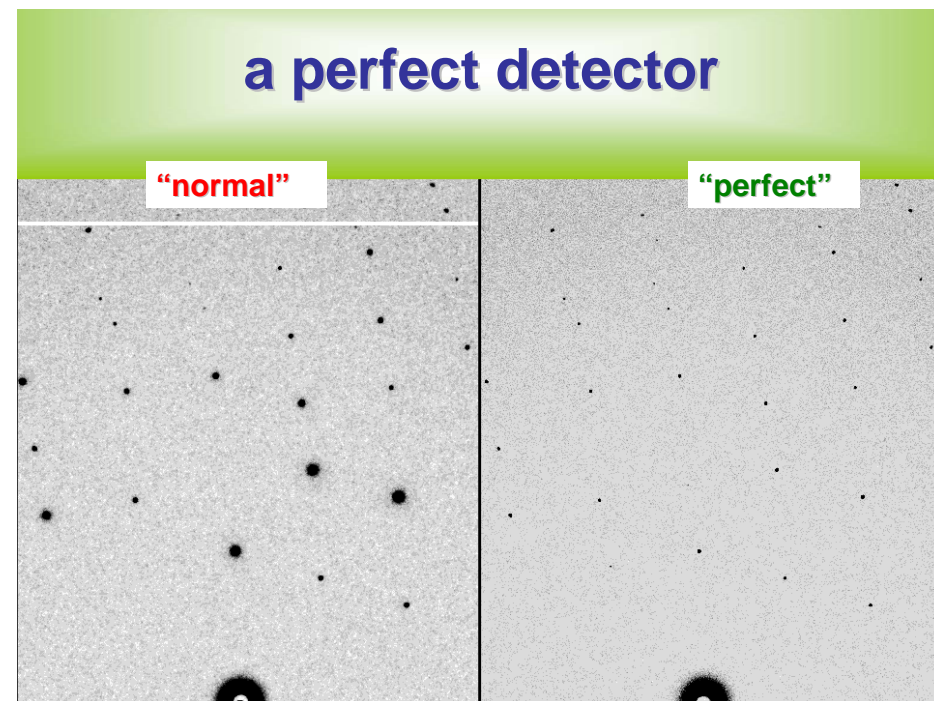
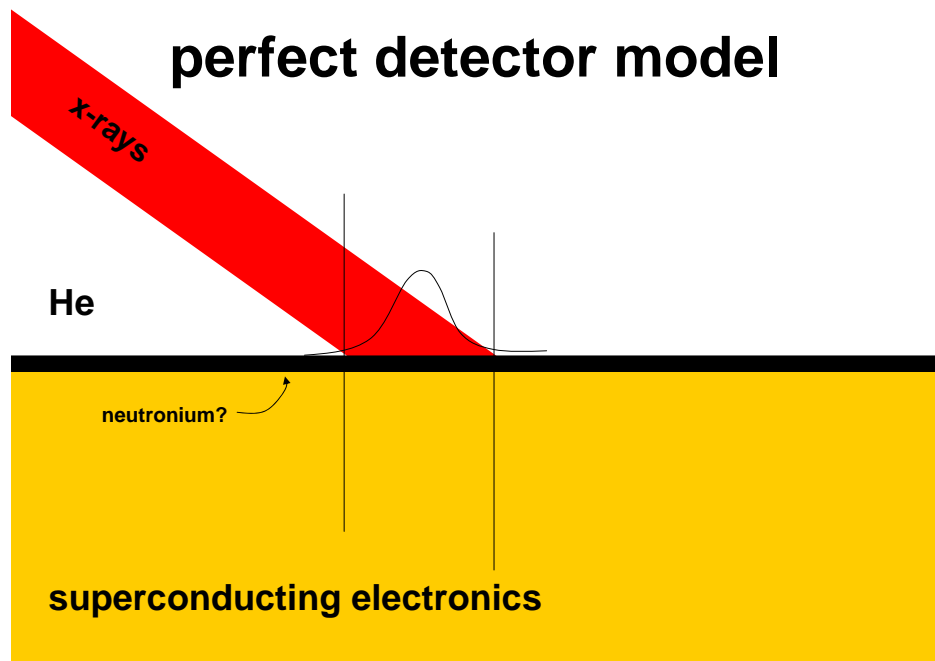
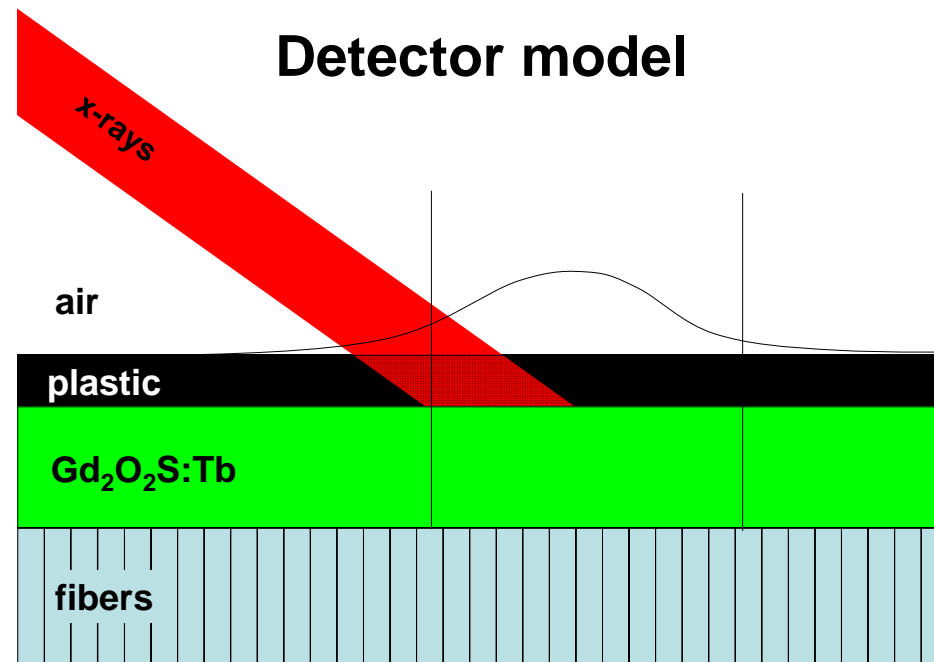
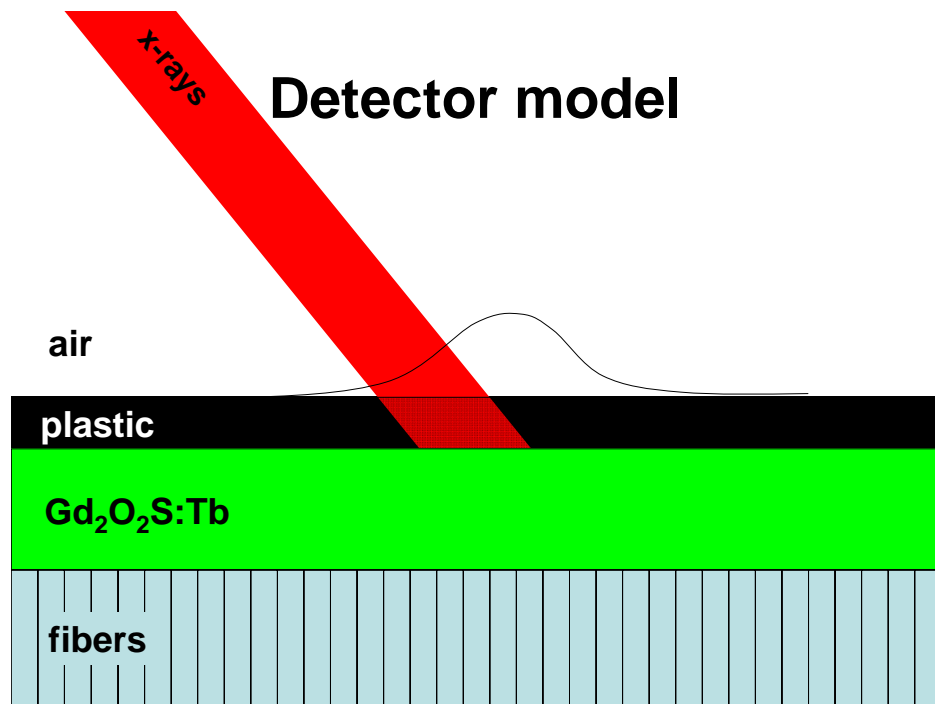
Detector model



Detector model







normal	statistic	perfect
4.0%	R _{merge}	3.9%
24.3	I/sd	28.1
3.7	I/sd (1.4 Å)	7.3
1.37 3 0.02	SDCORR	0.98 11 0.05
38.01	PADFPH	35.16
4.104	mlphare f''	3.973
0.224	FOM	0.272
0.579	FOMDM	0.563
0.7113	CC(1H87)	0.7051
15.72	R _{cryst}	16.64
18.30	R _{free}	19.38

**can we collect
shorter exposures
with
higher redundancy?**

same total dose with high and low redundancy

dataset	1	2-11	12
exposure	1.0s	0.1s	1.0s
frames	100	100 x 10	100
R _{merge}			
R _{anom}			
I/sd			
I/sd (2.0 Å)			
redundancy			
PADFPH			
FOM			
FOMDM			
CC(1H87)			

same total dose with high and low redundancy

dataset	1	2-11	12
exposure	1.0s	0.1s	1.0s
frames	100	100 x 10	100
R _{merge}	5.6%	11.2%	4.7%
R _{anom}	4.8%	4.7%	4.7%
I/sd	29.5	43.4	33.3
I/sd (2.0 Å)	23.3	29.6	25.8
redundancy	7.6	75.7	7.6
PADFPH	36.69	37.11	37.93
FOM	0.342	0.343	0.366
FOMDM	0.698	0.711	0.726
CC(1H87)	0.418	0.492	0.468

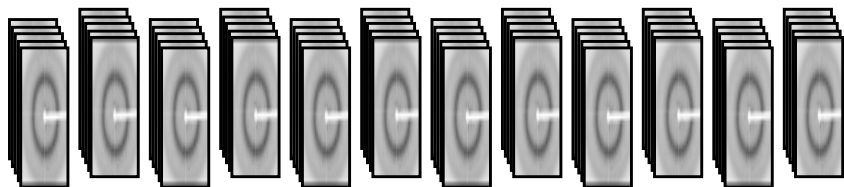
same total dose with high and low redundancy

dataset	1	2-11	12
exposure	1.0s	0.1s	1.0s
frames	100	100 x 10	100
R _{merge}	5.6%	11.2%	4.7%
R _{anom}	4.8%	4.7%	4.7%
I/sd	29.5	43.4	33.3
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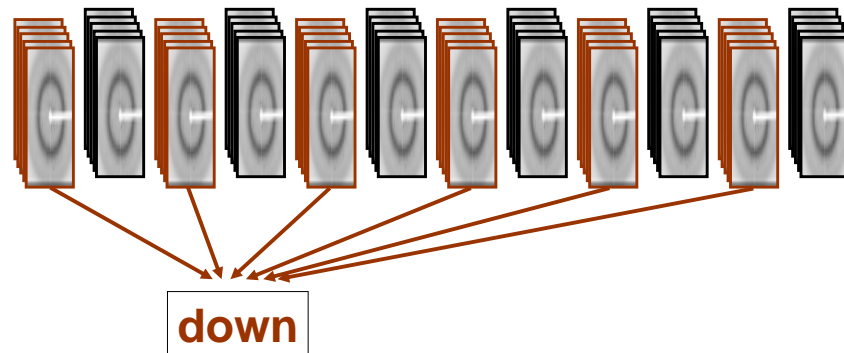
same total dose with high and low redundancy

dataset	1	2-11	12
exposure	1.0s	0.1s	1.0s
frames	100	100 x 10	100
R _{merge}	5.6%	11.2%	4.7%
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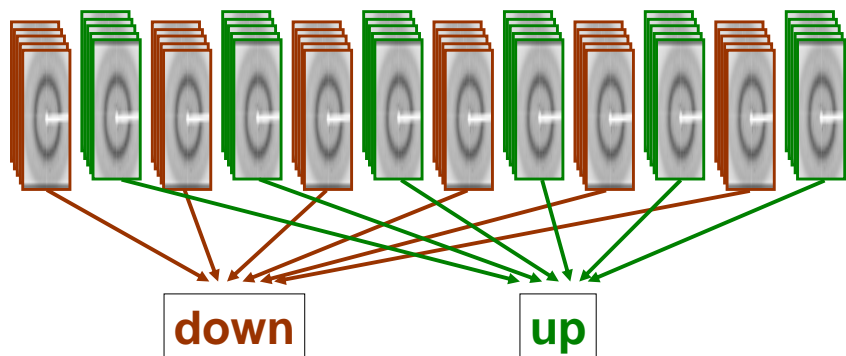
Spatial Noise



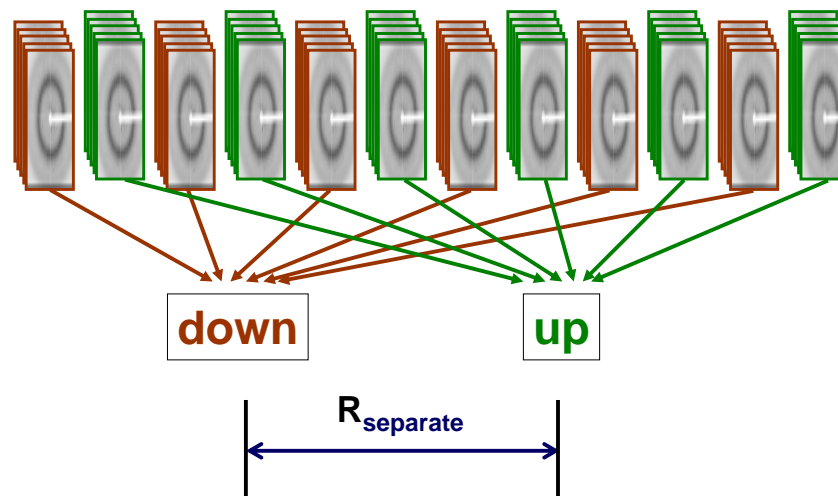
Spatial Noise



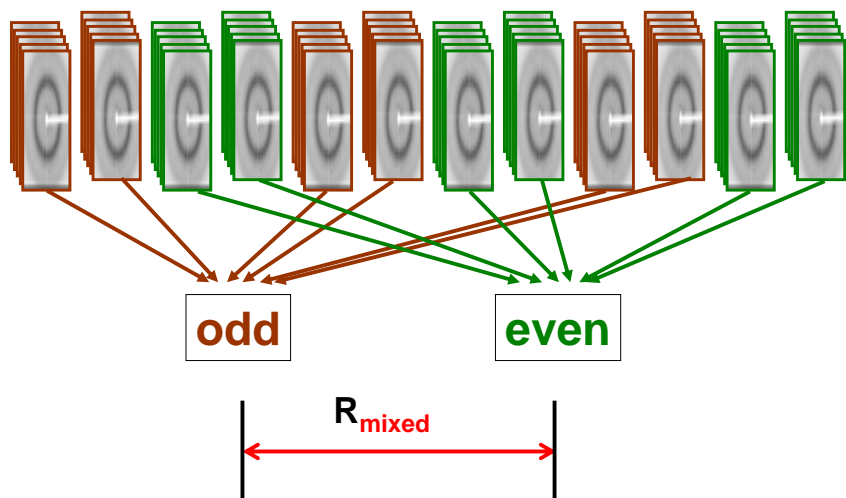
Spatial Noise



Spatial Noise



Spatial Noise



Spatial Noise

separate: 2.5%

Spatial Noise

separate: 2.5%

mixed: 0.9%

Spatial Noise

separate: 2.5%

mixed: 0.9%

$$2.5\%^2 - 0.9\%^2 = 2.3\%^2$$

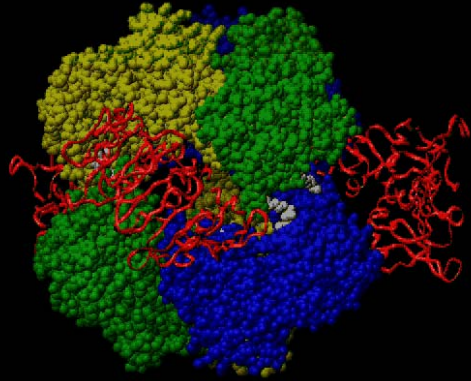
Spatial Noise

$$\text{mult} > \left(\frac{2.3\%}{\langle \Delta F/F \rangle} \right)^2$$

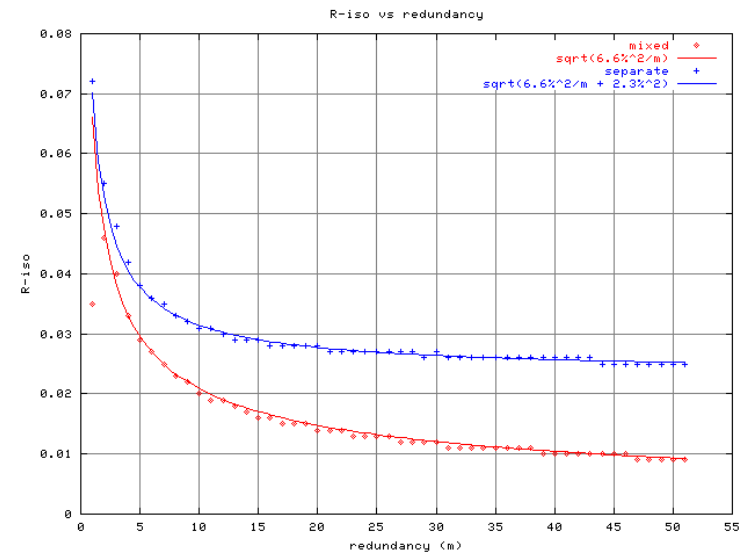
Spatial Noise

$$\text{mult} > \left(\frac{R_{\text{merge}}}{\langle \Delta F/F \rangle} \right)^2$$

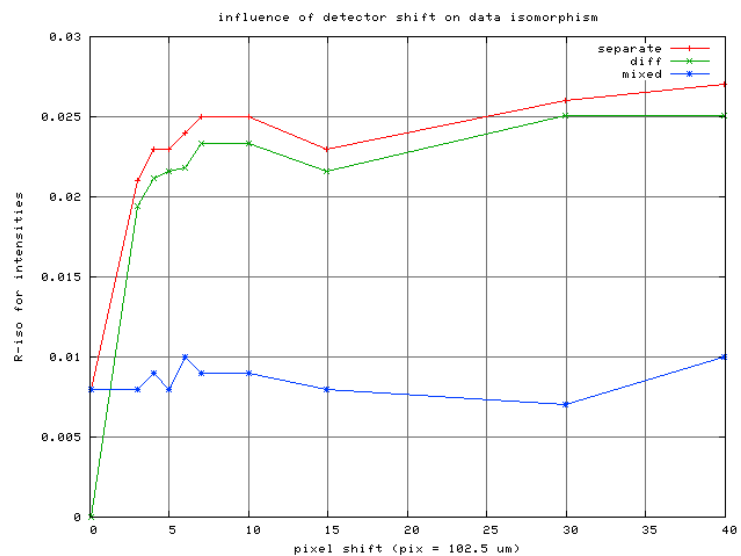
Proteins move



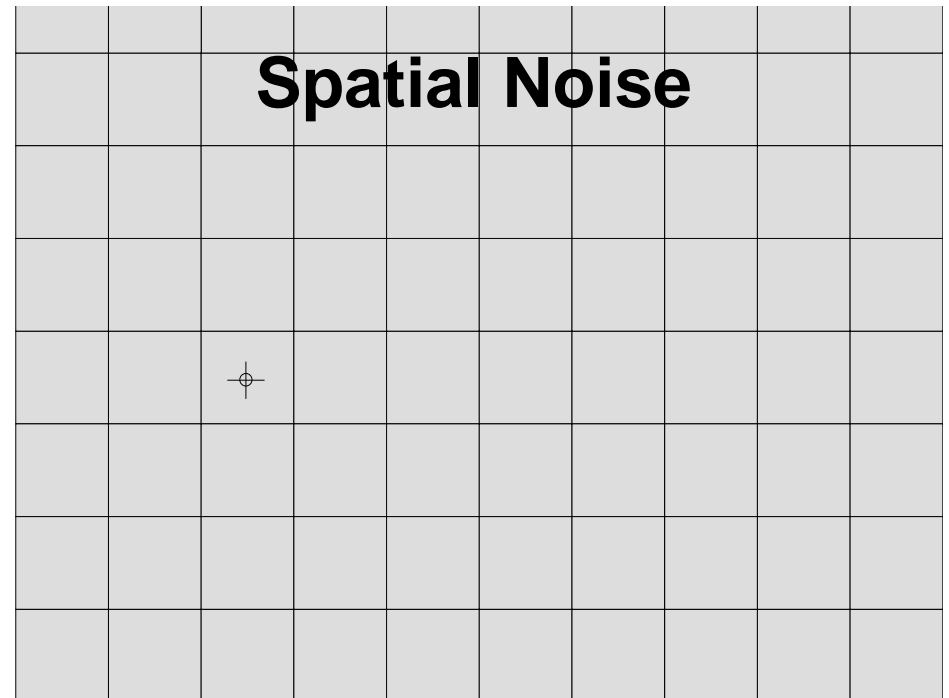
Spatial Noise



Spatial Noise



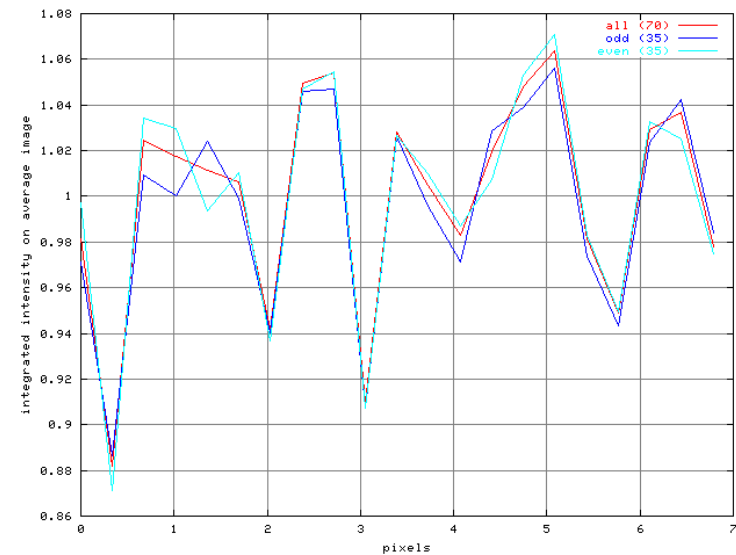
Spatial Noise



Spatial Noise



Spatial Noise



Summary

98% screening

short exposures FIRST

“dose slicing”

site damage is unpredictable

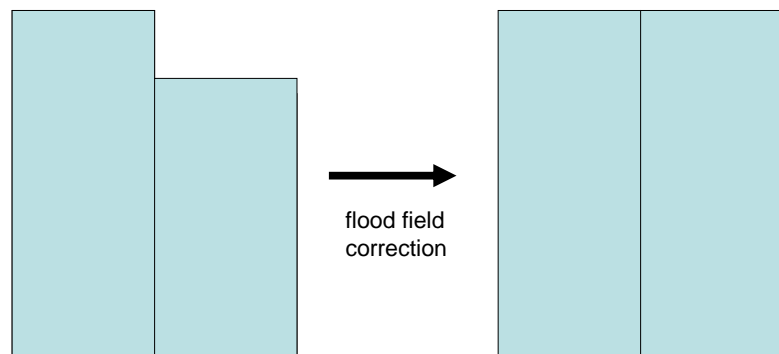
Development Snapshot

http://bl831.als.lbl.gov/~jamesh/mlfsom/development_snapshot.tar.gz

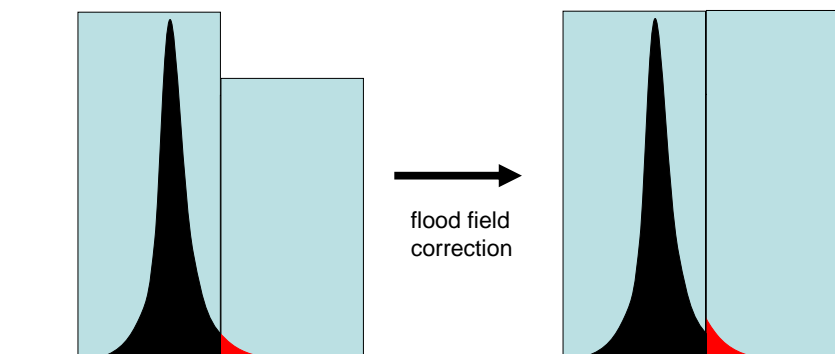
<http://bl831.als.lbl.gov/~jamesh/xtalsize.html>

http://bl831.als.lbl.gov/~jamesh/powerpoint/ACA_2009.ppt

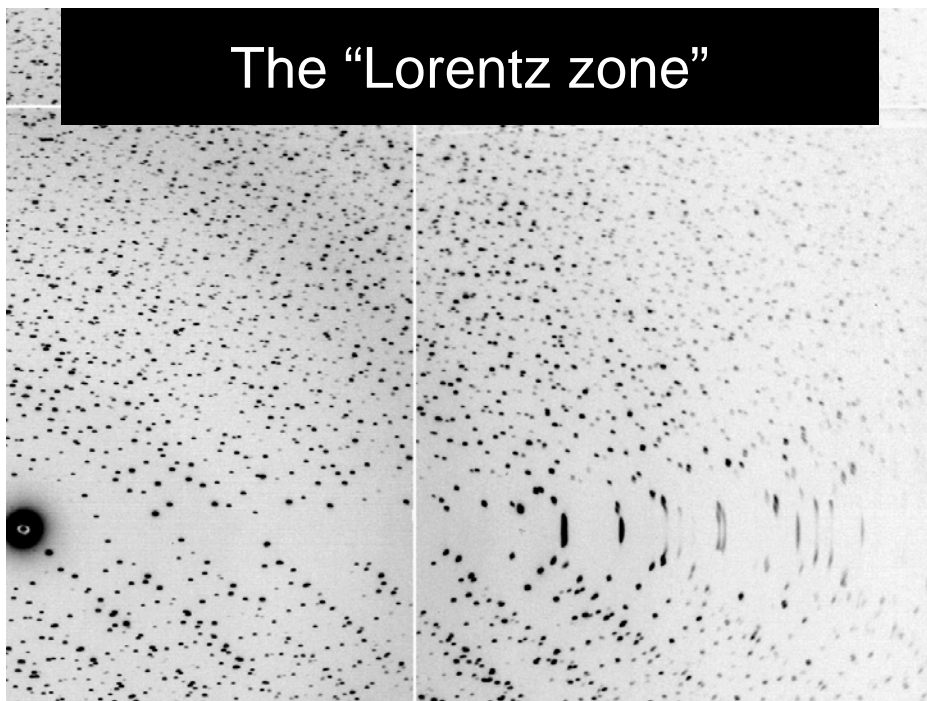
Spatial Noise



Spatial Noise

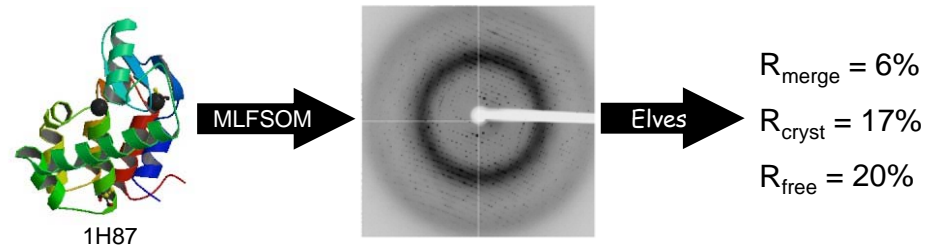


The “Lorentz zone”



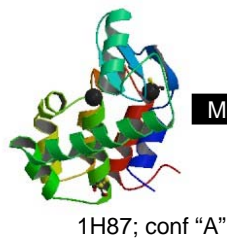
The R-factor Gap

multi-conformer PDB file

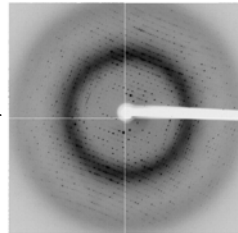


The R-factor Gap

single-conformer PDB file



MLFSOM



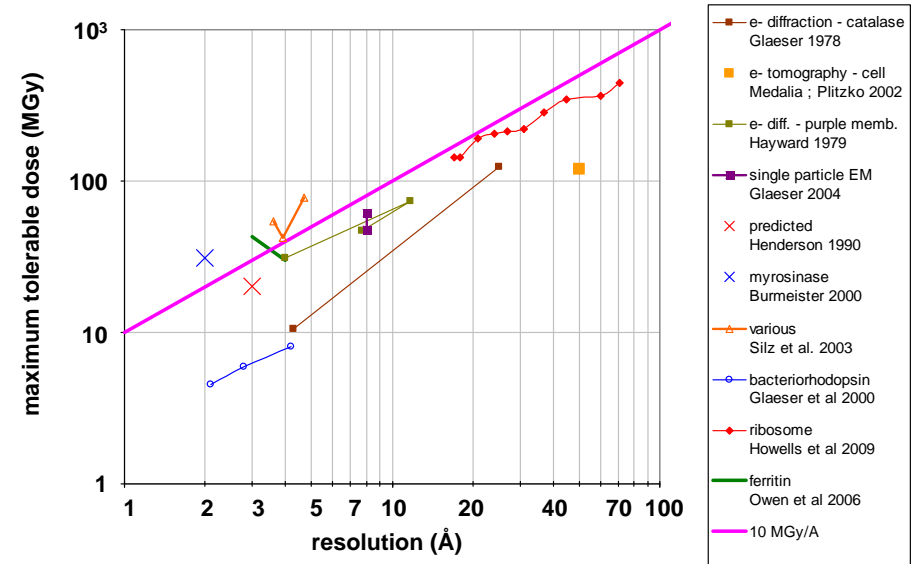
Elves

$$R_{\text{merge}} = 6\%$$

$$R_{\text{cryst}} = 7\%$$

$$R_{\text{free}} = 8\%$$

Howells *et al.* (2009)
J. Electron. Spectrosc. Relat. Phenom. **170** 4-12



Radiation Damage Model

global (lattice) damage

$$|F_{\text{used}}| = |F_{\text{undam}}| \exp(-\ln(2) \frac{\text{dose}}{d \cdot H})$$

- F_{used} - structure factor used for spot
- F_{undam} - structure factor of undamaged crystal
- dose** - absorbed dose (MGy)
- H** - 10 MGy/Å
- d** - resolution of spot (Å)

Radiation Damage Model

global (lattice) damage

$$F = F_0 \exp(-A \cdot s)$$

- F** - structure factor used for spot
- F_0 - structure factor of undamaged crystal
- A** - something Debye said was zero
- s** - 0.5/d
- d** - resolution of spot (Å)

Radiation Damage Model

global (lattice) damage

$$F = F_0 \exp(-A \cdot s)$$

- F** - structure factor used for spot
- F₀** - structure factor of undamaged crystal
- A** - something Debye said was zero
- s** - $0.5/d$
- d** - resolution of spot (Å)

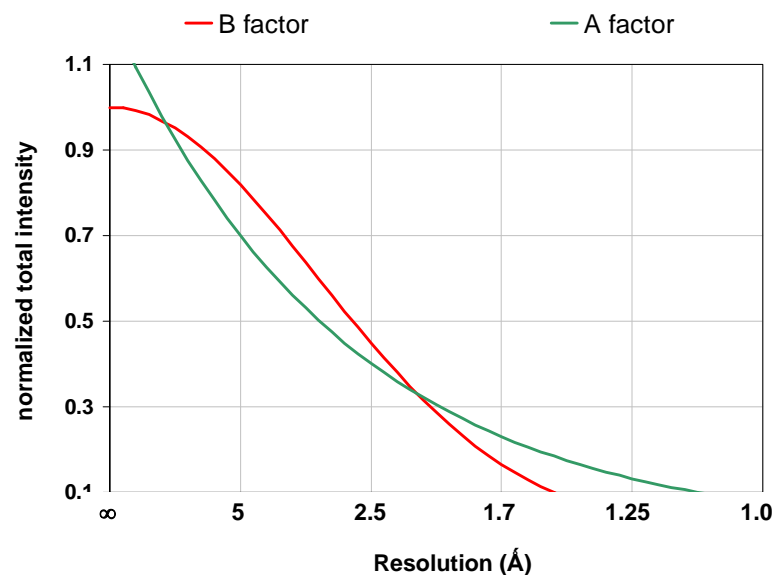
Radiation Damage Model

global (lattice) damage

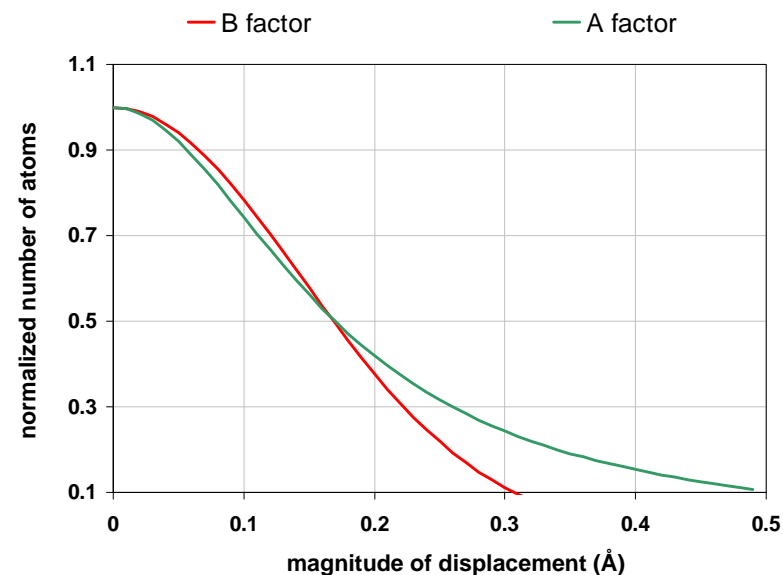
$$F = F_0 \exp(-A \cdot s - B \cdot s^2 - C \cdot s^3 - \dots)$$

- A** - something Debye said was zero
- B** - canonical Debye-Waller factor
- C** - something else Debye said was zero
- s** - $0.5/d$
- d** - resolution of spot (Å)

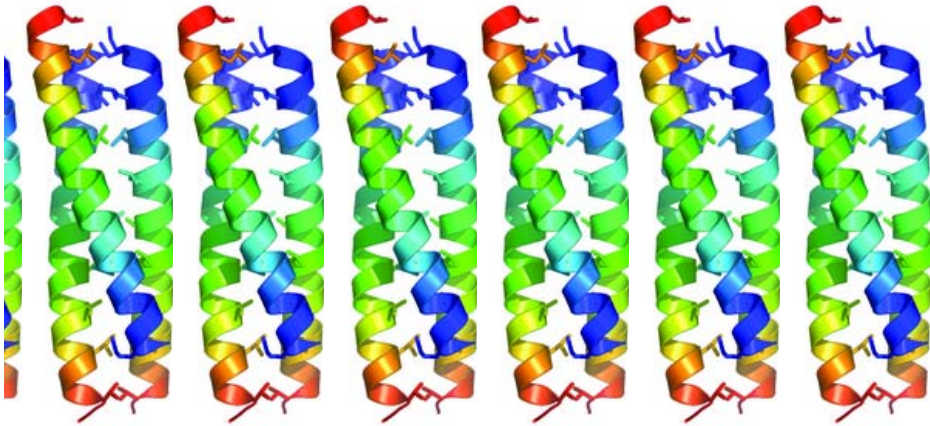
Radiation Damage Model



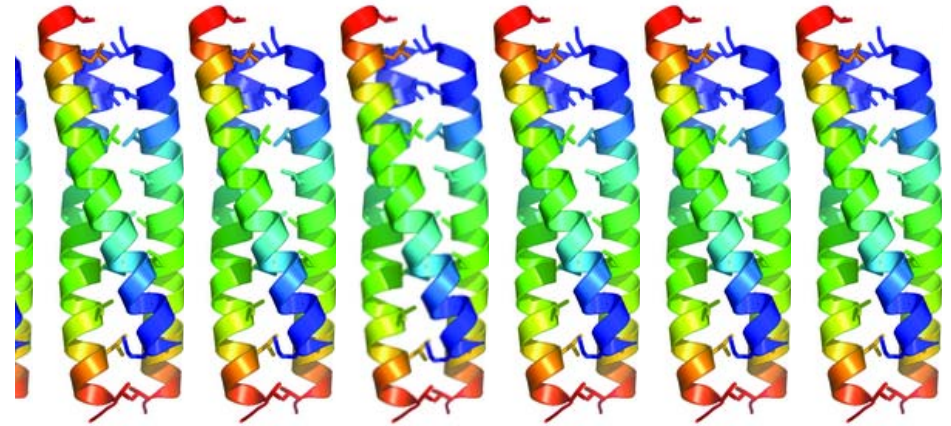
Radiation Damage Model



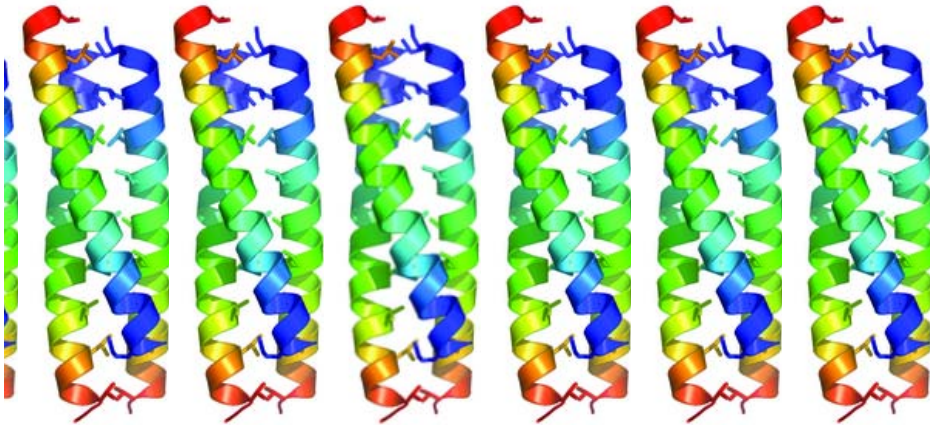
Radiation Damage Model



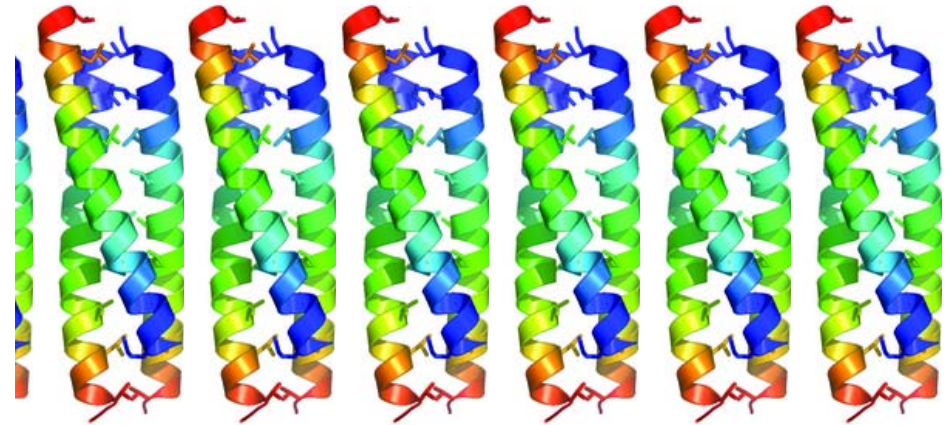
Radiation Damage Model



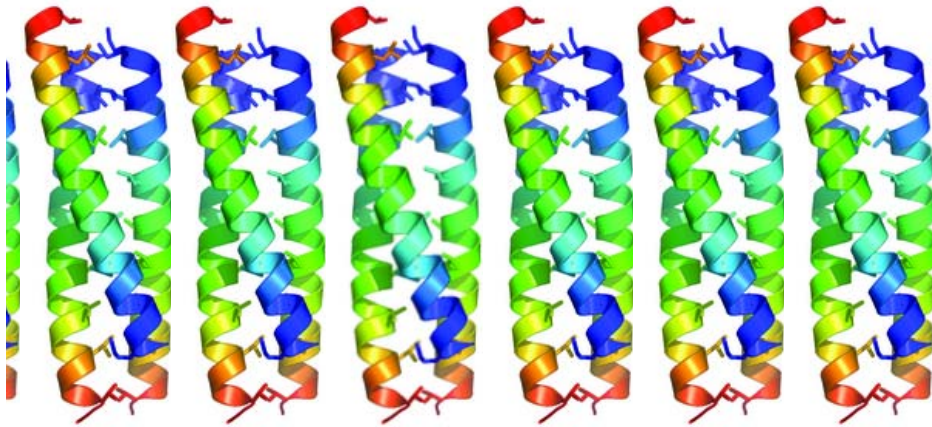
Radiation Damage Model



Radiation Damage Model

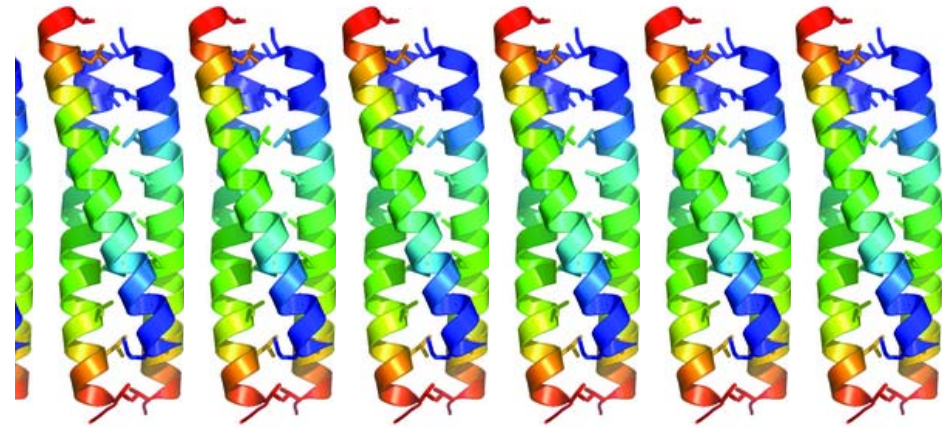


Radiation Damage Model



Kanzaki 1957

Radiation Damage Model



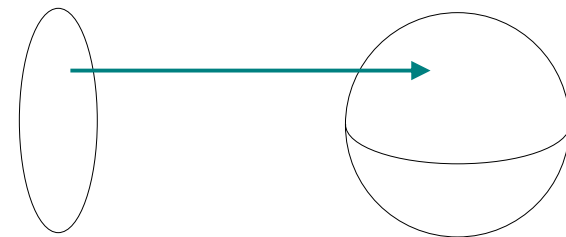
Kanzaki 1957

Radiation Damage prediction

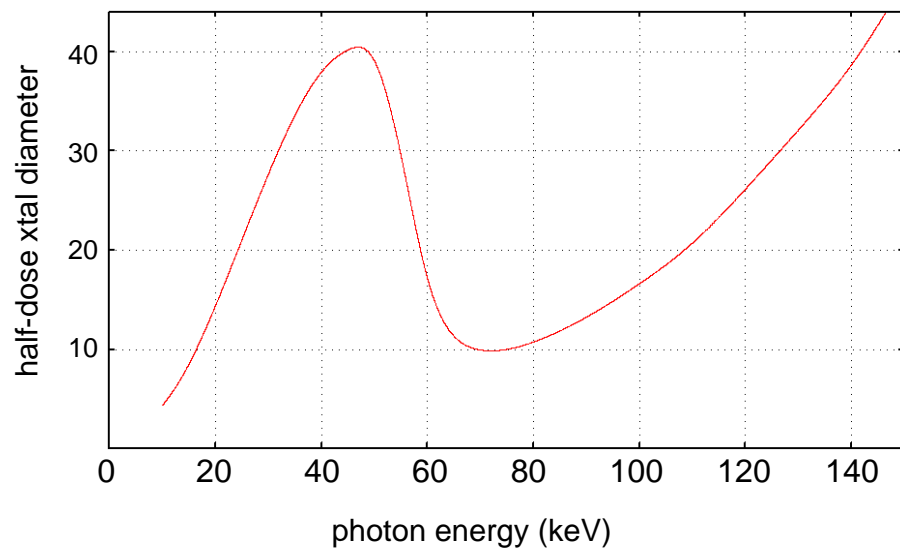
$$\frac{\text{data}}{\text{damage}} \approx k + \left(\frac{2 \text{ ps}}{\text{exposure}} \right)^3$$

particle transport simulation using MCNP

(other codes exist: EGSnrc)



Nave-Hill effect



Nave-Hill effect

